Incomplete lecture notes for the course of QUANTUM MECHANICS 20015/2016 M. Damnjanović

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## PREFACE

Tekst je i ove godine dopunjavan delovima predavanja, delimicno reorganizovan. Jedan deo zadataka sa vežbi je promenjen, pre svega zbog promene asistenta na predmetu (Marko Milivojević), ali i zbog načina selekcije. I dalje se ovo može smatrati samo pratećim materijalom predavanja, a nikako ne kompletnim i pročišćenim odrazom ispredavanih lekcija. To će, nadamo se, postati sledećih godina. Tekst dobrim delom nije proveravan nakon pisanja, a posebno ne lektorisan. Stoga treba biti skeptičan prema formulama (!), pa čak i nekim formulacijama, jer prilikom elektronskog editovanja svaka nepažnja može da bude vrlo kreativna. Dalje, nekompletan je: mada su sva predavanja obradjena, i u tom smislu tekst pokriva kurs, česte indikacije podnaslova ukazuju na projektovana proširenja neophodna za potpunije sagledavanje pojedinih tema, otkrivajući najvažnije konceptualne celine koje kursom nisu obuhvaćene. Slike i primeri koji treba da ilustruju sadržaje krajnje su redukovani. Zato molim studente da tekst bude pre svega podsetnik za ono sto treba uraditi, a da za konačnu pripremu ispita koriste dopunsku literaturu.
3.10.2015, M.D.

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## Chapter 1

## Quantum Kinematics

### 1.1 Measurement: System, State, Observable

This section introduces fundamental notions of quantum formalism, such as system, state, ensemble, measurement, etc. All these notions are common ingredients of any physical theory, but it turns out that their thorough analysis is not really necessary for classical mechanics. Further, we try to emphasize that these basic notions cannot be introduced independently, i.e. to define one of them the other ones should be also clarified. So, we start with intuitive analysis, being not essentially related to quantum mechanics, in order to single out the main features of the introduced notions. Their rigorous quantum definition will be given in the next section within three postulates, which are also mutually dependent and, together with definition of measurement, they should be understood as a whole.

### 1.1.1 Intuitive analysis

System (quantum) is an object $\boldsymbol{S}$ having some previously defined set of properties. This set is exhaustive enough to make distinction between various systems. E.g., each electron has same mass, charge, spin etc., and other particles are differing by some of them. This characterization by well defined properties is time (meaning technology) dependent: a hundred years ago, when mass and charge were the only known characteristics distinguishing elementary particles, many of them would be considered the same. Therefore, recognition of a particular system consists of a sufficient set of measurements of its properties.

The set of properties defining a system still leave some non-uniqueness: systems of the same type differ by some additional characteristics. Thus, one can gather systems with certain given values of some (or all) of these additional properties into ensemble $\rho$. We usually say that various ensembles appearing this way differ by the state (characterizing included systems). In this sense, the notions of ensemble and its state are synonyms: one says "state of the ensemble" and "ensemble in the state" (although the misleading phrases "state of the system", or "system in the state" are often used instead). An example may be an ensemble of electrons with given momentum, i.e. electrons in a state with given value of momentum. Note that this notion is based on the set of measurements of some
properties, as well as the primary notion of the system. Consequently, two ensembles are in the same state if all the measurements performed on them give the same results. Again, the notion of the state is dependent on our knowledge of the possible properties of the system. Important example is that before discovery of the spin, ensembles (i.e. states) differing only in the value of the spin were considered the same.

Intuitively, measurement is any process of determining of some property of an ensemble. In fact, the very idea of a property of the system is that it is observable, i.e. measurable. This means that for each property (observable) $A$ there is at least one measuring device (in this sense, the notions of the physical quantity and of the device measuring it may be identified), apparatus $\boldsymbol{A}$, capable to distinguish between various values of $A$ : each value corresponds to a particular position of the apparatus' pointer. In other words, in the course of the interaction of the apparatus with the measured system some of these values is realized. Two important facts should be emphasized in this context. Firstly, various members of the same ensemble may produce different pointer positions. Therefore, measurements on the ensemble necessarily have statistical nature. Secondly, before measurement is performed nothing can be said about the measured property. This may be interpreted such that the measured property does not exist without the apparatus, or even that the property is realized or imposed by measurement.

The statistical nature of the measurement is well known even within the classical framework: a measurement does not give the result with certainty, but many measurements are performed, and the result is obtained by statistical analysis. Therefore, a single measurement is meaningless. Further, since the measurement is an interaction of apparatus and system, it may change, or even destruct the system (this is particularly important for small systems), and it may be impossible to repeat the procedure on the same system. Therefore, in general, measurements are performed on the ensembles of systems. Nevertheless, as it will be stressed out in the analysis of double slit experiment (Subsection 1.1.2), the necessity for the statistical approach and ensembles in quantum mechanics stems from additional, quite substantial reason.

To summarize, measurement means measurement of some physical observable $A$ on the ensemble (in the state) $\rho$, giving as the result probability distribution of the possible values of $A$. Precisely, let $\sigma(A)=\left\{a_{1}, a_{2}, \ldots\right\}$ be the set of the possible values of $A$ (i.e. of the positions of the pointer of the apparatus, defined independently of $\rho$ ); each particular system from the measured ensemble $\rho$ interacts with apparatus successively, and due to this interaction the pointer gets a series of values $a_{i}$ from $\sigma(A)$. Let in the course of this measurement each value $a_{i}$ is pointed to altogether $N_{i}$ times. Obviously, $N=\sum_{i} N_{i}$ is the number of systems in the ensemble. Then the result of this measurement is the probability distribution $v\left(a_{i}, A, \rho\right) \stackrel{\text { def }}{=} N_{i} / N$. Since the only criterion of the validity of any physical theory (and particularly quantum mechanics) is the comparison with the experiment, the fundamental task of such a theory is to give prediction for $v\left(a_{i}, A, \rho\right)$ for each $\rho$ and $A$ in terms of its formalism. Before proceeding further in this direction, several remarks should be made.

Firstly, we note here that the (quantum) theory can be well founded only with infinite ensembles, and in the rest of the text this will be always assumed. However, in the real experiments $N$ must be finite, but large enough to provide reliable statistics, i.e. the statistics enabling comparison to the theoretical $N=\infty$ limit.

Secondly, since only the results of the measurements are verifiable, thus being the only objective of quantum mechanics, the quantum mechanical statements always refer to them. The underlying logic (this is not classical Boolean logic, as it will be clarified soon) is therefore completely subdued to this fact, with all the statements based on the results of the measurements. Considering the statements as the answers to the corresponding questions, the above formulated task of a physical theory means its ability to give the answers (i.e. predictions) to the questions which can be at least in principle answered in terms of the experimental results. Together with the elementary statements about probabilities of results, also the statements which can be reduced to the elementary ones are legitimate. For example, besides the question about probability of a single value $a_{i}$, one can search for the probability $v\left(\left\{a_{i}, a_{j}, \ldots\right\}, A, \rho\right)$ to get any of the values from the subset $\left\{a_{i}, a_{j}, \ldots\right\}$ of $\sigma(A)$; obviously, this must satisfy

$$
v\left(\left\{a_{i}, a_{j}, \ldots\right\}, A, \rho\right)=v\left(a_{i}, A, \rho\right)+v\left(a_{j}, A, \rho\right)+\ldots
$$

Similarly, mean value of the observable $A$ in the state $\rho$

$$
\begin{equation*}
\langle A\rangle_{\rho}=\sum_{i} a_{i} v\left(a_{i}, A, \rho\right) \tag{1.1}
\end{equation*}
$$

may be looked for.
For further convenience, in the case when $v\left(a_{i}, A, \rho\right)=1$ for some particular value $a_{i}$ and consequently $v\left(a_{j}, A, \rho\right)=0$ in all other cases $i \neq j$, we will say that the state $\rho$ is with the sharp value $a_{i}$ of the observable $A$.

Further, it is easy to understand that due to its infinity, each ensemble $\rho$ can be partitioned into (also infinite!) subensembles $\rho_{1}, \rho_{2}, \ldots$ in many ways. The weight $w_{I}$ of the subensemble is defined as its relative part in the ensemble. Considering the process of the partition as putting one system after another from $\rho$ into one of $\rho_{I}$, in the $N$-th step one can find the numbers $N_{i}$ of the systems in subensembles, and define $w_{I}=\lim _{N \rightarrow \infty} N_{I} / N$. Obviously, $w_{I}$ is finite, positive and $\sum_{I} w_{I}=1$. Also, having ensembles $\rho_{1}, \rho_{2}, \ldots$ (with the same type of the system) one can make their mixture just joining all of them into single ensemble $\rho$. Immediately after this, it cannot be reconstructed how the total ensemble is constructed, i.e. the corresponding partition is by no means singled out.

Being also ensemble, each subensemble is suitable for measurements. Of course, the results of the subensemble measurements are in general mutually different, and also differ from the results of the ensemble measurement. In fact, the results of the subensemble measurements $v\left(a_{i}, A, \rho_{I}\right)$ are to the ensemble one related only by

$$
\begin{equation*}
v\left(a_{i}, A, \rho\right)=\sum_{I} w_{I} v\left(a_{i}, A, \rho_{I}\right) . \tag{1.2}
\end{equation*}
$$

Still, there are ensembles giving for any observable $A$ the same results $v\left(a_{i}, A, \rho_{I}\right)=$ $v\left(a_{i}, A, \rho_{J}\right)$, and therefore $v\left(a_{i}, A, \rho\right)=v\left(a_{i}, A, \rho_{I}\right)$ in the all subensembles. These are called homogeneous ensembles or pure states. All other ensembles are called inhomogeneous, or mixed states. Pure states characterize maximally prepared ensembles. In fact, recall that the measurements required to define a particular system leave some freedom for different states which are distinguished between by some additional measurements.


Figure 1.1: Double slit experiment. (a) Device: C is source of collimated photons, $\mathrm{s}_{1}$ and $\mathrm{s}_{2}$ are slits on the screen, $X$ is detector of the particles (e.g. photo plate); $\mathrm{D}_{1}$ and $\mathrm{D}_{2}$ are detectors of the particles on the slits. (b) Results, being the probability distributions (intensities) $v(x)$ obtained by $X: v_{i}(i=1,2)$ is obtained when only $\mathrm{s}_{i}$ is opened, $v_{1 \not 22}$ is their normalized sum, while curve $v$ is obtained when both slits are opened.

If all the possible additional measurements are performed, and ensemble is made only of the systems possessing the wanted values of measured observables, the homogeneous ensemble is obviously obtained. Despite we used here the intuitive notion of the possible (additional) measurements, it is quite nontrivial, and will be neatly reconsidered within quantum mechanical formalism, yielding to the one of the most striking differences of classical and quantum physics.

### 1.1.2 Superposition: Experiments and Phenomenology

Let us consider neatly well known double slit experiments. At first we discuss only the simplest scheme, but later on it will be stressed out that this is in a sense the most general experimental situation.

The experimental device is given in Fig. 1.1, panel (a). The source (C) prepares homogeneous ensemble $\psi$ directed toward the plate. Starting from the source, systems pass through slits $\mathrm{s}_{1}$ and $\mathrm{s}_{2}$, and come to the detector. Obviously, detector is measuring device showing position $X$ of the incoming systems, i.e. the result of this measurement is the probability distribution $v(x, X, \psi)$.

If only slit $\mathrm{s}_{1}$ is opened, the result is distribution $v_{1}$ (Fig. 1.1, panel (b)), and similarly with only slit $\mathrm{s}_{2}$ distribution $v_{2}$ is obtained. Now, we make two experiments with both slits open. In the first one detectors $D_{1}$ and $D_{2}$ are switched on, providing thus the
information which slit (i.e. path) is chosen. The resulting distribution ${ }^{1}$ is $v_{1 \not 22}=v_{1}+v_{2}$. On the contrary, in the second experiment the detectors $D_{1}$ and $D_{2}$ are turned off. Now, when the path of the system is not known, the experimental distribution is $v$.

In the terms of classical theory, the first distribution $v_{1 \nLeftarrow 2}$ corresponds to the ensemble of corpuscules, when intensity is sum of the intensities obtained with one slit closed. The second one $v$ is like interference of waves (of course, we did not analyze here the conditions enabling distribution $v$, like distance between slits, etc.), when the intensities are proportional to the square of fields: $v_{1} \sim\left|\boldsymbol{E}_{1}\right|^{2}, v_{2} \sim\left|\boldsymbol{E}_{2}\right|^{2}$, meaning that $v \sim\left|\boldsymbol{E}_{1}+\boldsymbol{E}_{2}\right|^{2}$. However, two experiments show that the same system exhibits both corpuscular and wave behavior, depending on the experimental situation. This wave-corpuscle dualism essentially means that system is neither corpuscle nor wave, but some of its properties, being manifested in appropriate circumstances, can be interpreted in terms of these classical (perhaps intuitive, or, more likely, familiar after courses of classical mechanics) prototypes of behavior. More important is to realize that the interference corresponds to the lack of information on the path of systems. In fact, we can say that the resulting probability is caused by the interference of possible paths, and not particles (which is excluded by the experiments in which one particle comes out after another from source). Thus, if it not measured (by $D_{1}$ and $D_{2}$ in the above experiment), even the trajectory of a particle does not exist; only superposition of all possible paths corresponds to experimental results. Finally, note that the path is not known without additional measurement (by $D_{1}$ and $D_{2}$ ) because the prepared initial state does not prefer either of two possibilities allowed by the slits on the screen. Thus, for such a state and such screen there are two equally probable paths. However, quite different situations may be easily constructed: e.g. initial state allows only one path (beam directed toward one slit only), the paths are not equally probable, or more than two possible paths.

### 1.1.3 Pure States

The described experiments, and many others performed with pure states can be explained consistently only if the superposition principle is introduced: if two pure states of a system are detected, then any linear combination of these states can be realized, too. Obviously this implies that the pure states are vectors in some vector space, which will be called state space. Moreover, the superposition of the states with the same sharp value of some particular property, will be the state with the same value of that property. Otherwise even the notion of the system would be lost, since the superposition of the states of the same system (defined by the fixed values of the system defining properties) could change the value of these properties, and hereby the system.

## POSTULATE 1. - STATES

Each pure state of the system $S$ is represented by a unit vector in the state space $\mathcal{S}$ of $S$, and each such a vector represents a pure state of $S$. All the vectors differing by a phase factor only correspond to the same physical state.

[^0]
### 1.1.4 Observables

To complete the description of the measurement process, one must include the property which is measured. Intuitively, it can be assumed that all the directly measured quantities are real, despite the fact that we can combine them to get complex or tensor quantities convenient for some specific purpose. This is to be understood in the sense that outcomes of the measurements are reals.

Further, let us single out some particular value $a$ of the physical property $A$. Recall that any superposition of the pure states with the sharp value $a$ of $A$ results in the state with the same sharp value of $A$. This means that in the state space $\mathcal{S}$ there is a subspace $\mathcal{S}_{a}(A)$ including all such pure states. Consequently, taking into account all the possible measured values of $A$, the observable $A$ introduces a decomposition of the state space into these subspaces: $\mathcal{S}=\oplus_{a} \mathcal{S}_{a}(A)$. These two entities (real values and subspaces) can be interpreted as the spectrum and eigen decomposition of $\mathcal{S}$, i.e. they define uniquely a hermitian operator in $\mathcal{S}$ by its spectral form:

$$
\begin{equation*}
A=\sum_{a} a P_{a}(A) \tag{1.3}
\end{equation*}
$$

$P_{a}(A)$ denotes the eigenprojector of $A$ for the eigenvalue $a$, i.e. the projector onto the eigenspace $\mathcal{S}_{a}(A)$.

## POSTULATE 2. - OBSERVABLES

Each observable biuniquely corresponds to a hermitian operator in the system state space $\mathcal{S}$.

Thus, within the quantum formalism, physical observables are represented by the hermitian operators. On the other hand, this also means that whatever hermitian operator is considered (and in the infinite dimensional Hilbert spaces there are really strange ones), there is at least in principle possibility to construct the measuring device corresponding to it.

It is obvious that the eigenvectors of the various physical observables play an important role in the quantum formalism, as well as in the concrete calculations. It is convenient to denote the eigenvector corresponding to the eigenvalue $a$ of $A$ as $|a, \lambda\rangle$, where the counter $\lambda$ distinguishes between various (linearly independent) eigenvectors for the same eigenvalue. Thus an eigenbasis of $A$ is

$$
\left\{|a, \lambda\rangle\left|a \in \sigma(A), \lambda=1, \ldots,\left|\mathcal{S}_{a}(A)\right|\right\}\right.
$$

Having an eigenbasis of $A$, one can use it to represent vectors and operators in $\mathcal{S}$ by the columns and matrices. Such a representation will be called $A$-representation.

Particularly, if $A$ is complete or nondegenerate observable, i.e. if its eigenspaces are one-dimensional, $|a\rangle$ uniquely determines the physical state. If $A$ is incomplete, since the commuting observables have common eigenbasis, the notation can be refined using instead of $\lambda$ the eigenvalues of another commutative observable $B$. In fact, let us consider a basis $|a, \lambda\rangle$ in $\mathcal{S}_{a}(A)$, which is also an eigenbasis of $B$ (such a basis exits): $B|a, \lambda\rangle=b_{\lambda}|a, \lambda\rangle$. The eigenvalues $a$ and $b_{\lambda}$ characterizing the common eigenvectors are called compatible. If
all the corresponding eigenvalues $b_{\lambda}$ are different, then these eigenvalues completely characterize this basis, and dropping the redundant label $\lambda$, the vectors are naturally denoted as $|a, b\rangle$. If this is the case for all the eigenvalues of $A$, the common eigenbasis is unique. In general, the set of the observables $\left\{A^{(1)}, \ldots, A^{(N)}\right\}$ having uniquely determined common eigenbasis $\left|a^{(1)}, \ldots, a^{(N)}\right\rangle$ is called complete set of compatible observables (CSCO). This means that these vectors are unique ones satisfying the system of eigenequations:

$$
A^{(i)}\left|a^{(1)}, \ldots, a^{(N)}\right\rangle=a^{(i)}\left|a^{(1)}, \ldots, a^{(N)}\right\rangle, \quad i=1, \ldots, N
$$

### 1.1.5 Probability distribution

To complete the quantum formalism, it remains to interrelate states and observables (already incorporated into formalism) to get the prediction for result of the measurement of a particular observable on the system in some particular state. To this end we use spectral decomposition (1.3) to find the components from the eigenspaces of $A$ of the pure state $|\psi\rangle:|\psi\rangle=\sum_{a} P_{a}(A)|\psi\rangle=\sum_{a}\left|\psi_{a}\right\rangle$. Since $|\psi\rangle$ is unit vector, the square norms of these components are summed in 1: $\sum_{a} \|\left|\psi_{a}\right\rangle \|^{2}=1$, as it is expected from any probability distribution. This hints that the probability to find $a$ is $\|\left|\psi_{a}\right\rangle \|^{2}$. A straightforward generalization is that the probability of the results from the interval $\left[a_{1}, a_{2}\right]$ is $\sum_{a \in\left[a_{1}, a_{2}\right]} \|\left|\psi_{a}\right\rangle \|^{2}$, giving a justification of:

## POSTULATE 3. - PROBABILITIES

When the measurement of the observable $A$ is performed on the system in the pure state $|\psi\rangle$, the value from the interval $\left[a_{1}, a_{2}\right]$ is obtained with the probability

$$
v\left(\left[a_{1}, a_{2}\right], A,|\psi\rangle\right)=\| P_{\left[a_{1}, a_{2}\right]}(A)|\psi\rangle \|^{2} .
$$

Later on this expression will be used also in the equivalent forms:

$$
\begin{equation*}
v\left(\left[a_{1}, a_{2}\right], A,|\psi\rangle\right)=\langle\psi| P_{\left[a_{1}, a_{2}\right]}(A)|\psi\rangle=\operatorname{Tr} P_{\left[a_{1}, a_{2}\right]}(A)|\psi\rangle\langle\psi| . \tag{1.4}
\end{equation*}
$$

For the observables with purely discrete spectrum within considered interval, $P_{\left[a_{1}, a_{2}\right]}(A)$ is sum $\sum_{a \in\left[a_{1}, a_{2}\right]} P_{a}(A)$ of the eigenprojectors related to the enclosed eigenvalues. However, if $\left[a_{1}, a_{2}\right]$ encloses only continual spectrum of $A$, there is no eigenvector for a particular continual "eigen-value" $a$ (precisely, the point of the continual spectrum) and therefore $P_{a}(A)$ vanishes, as well as the related probability, $v(a, A,|\psi\rangle)=0$. Still, the projector $P_{\left[a_{1}, a_{2}\right]}(A)$ is well defined with help of the spectral measure $E_{a}(A)$ associated to $A$ : $P_{\left[a_{1}, a_{2}\right]}(A)=E_{a_{2}}(A)-E_{a_{1}}(A)$; in terms of the (generalized) eigenvectors, for a complete observable $A$ (generalization is obvious), this is $P_{\left[a_{1}, a_{2}\right]}(A)=\int_{a_{1}}^{a_{2}} \mathrm{~d} a|a\rangle\langle a|$. This corresponds to the intuitively expected finite probability to obtain result in some finite interval, and naturally invokes probability density at any point $a \in\left[a_{1}, a_{2}\right]$. Indeed, using $A$-representation $|\psi\rangle=\int \mathrm{d} a \psi(a)|a\rangle$ of the state $|\psi\rangle$, the first of (1.4) gives for the interval around $a$ the probability: $v(a-\Delta a / 2, a+\Delta a / 2], A,|\psi\rangle)=\int_{a-\Delta a / 2}^{a+\Delta a / 2}\left|\psi\left(a^{\prime}\right)\right|^{2} \mathrm{~d} a^{\prime}$. Its differential defines the probability density in $a$ :

$$
\begin{equation*}
\tilde{v}\left([a, A,|\psi\rangle)=\lim _{\Delta a \rightarrow 0} \frac{v([a-\Delta a / 2, a+\Delta a / 2], A,|\psi\rangle)}{\Delta a}=|\psi(a)|^{2} .\right. \tag{1.5}
\end{equation*}
$$

The postulate clearly introduces the basic elements of quantum logic. Since it involves the state which is measured, and the projector defining the expected event, one can generalize the notion of the postulate: referring to the projector as to the event (thus the subspaces are events, while the elementary events are one dimensional subspaces, i.e. rays) the expression $v(P,|\psi\rangle)=\| P|\psi\rangle \|^{2}$ is the probability of that event when the system is prepared in the state $|\psi\rangle$.

Note that each pure state $|\phi\rangle$ defines ray projector $P=|\phi\rangle\langle\phi|$. Then, there exists observable $A$ with the spectral decomposition $A=\alpha P+\sum_{a}^{\prime} a P_{a}(A)$ (summation over the eigenvalues different from $\alpha$ ). The probability of the result $\alpha$ in the measurement of $A$ in the initial state $|\psi\rangle$ is easily found by (1.4). Since it can be interpreted as the probability to find state $|\phi\rangle$ when the system is prepared in the state $|\psi\rangle$, it is called transition probability from $|\psi\rangle$ to $|\phi\rangle$ :

$$
\begin{equation*}
v(1,|\phi\rangle\langle\phi|,|\psi\rangle)=v_{|\psi\rangle \rightarrow|\phi\rangle}=|\langle\phi \mid \psi\rangle|^{2} . \tag{1.6}
\end{equation*}
$$

### 1.1.6 Mixed State

Immediately after introducing the notion of the ensemble, it has been pointed out that besides the homogeneous ones, there are also the inhomogeneous ones. While the former are described by the state vectors, the latter have not been incorporated into the formalism. To describe mixed states, we again rely on the obvious facts concerning the results of measurements.

We consider an inhomogeneous ensemble $\rho$ obtained by mixing, i.e. joining, of the homogeneous ensembles: $N_{i}$ replicas of the homogeneous ensemble described by $\left|\psi_{i}\right\rangle$ is taken together, for $i=1, \ldots, n$. Denoting by $N=\sum_{i} N_{i}$ the total number of the joined ensembles, the weight of each homogeneous component is given by $\omega_{i}=N_{i} / N$; naturally, weights $\omega_{i}$ satisfy $0 \leq \omega_{i} \leq 1$ and $\sum \omega_{i}=1$. The probability distribution resulting in the measurement of the observable $A$ is obviously the average of the distributions $v\left(\left[a_{1}, a_{2}\right], A,\left|\psi_{i}\right\rangle\right)$ obtained in the measurements on the involved pure states:

$$
\begin{equation*}
v\left(\left[a_{1}, a_{2}\right], A, \rho\right)=\sum_{i} \omega_{i} v\left(\left[a_{1}, a_{2}\right], A,\left|\psi_{i}\right\rangle\right)=\operatorname{Tr} P_{\left[a_{1}, a_{2}\right]}(A) \rho, \tag{1.7}
\end{equation*}
$$

where

$$
\begin{equation*}
\rho=\sum_{i} \omega_{i}\left|\psi_{i}\right\rangle\left\langle\psi_{i}\right| . \tag{1.8}
\end{equation*}
$$

Comparing with (1.4), one finds that again $P_{\left[a_{1}, a_{2}\right]}(A)$ comprises the information on the measuring observable, while the remaining part, $\rho$, gives all available information about the system, i.e. ensemble. Hence, $\rho$ is the entity describing the ensemble. For example, using $A$ instead of its eigenprojectors in (1.7) one finds the meanvalue:

$$
\begin{equation*}
\langle A\rangle=\operatorname{Tr} \rho A \tag{1.9}
\end{equation*}
$$

Being a convex linear combination of the ray projectors $\left|\psi_{i}\right\rangle\left\langle\psi_{i}\right|, \rho$ is obviously hermitian, positive and with unit trace. Thus any mixed state is described by a statistical
operator. Note that the orthogonality of the states $\left|\psi_{i}\right\rangle$ has not been assumed, and therefore (1.8) is not a spectral form of $\rho$. Nevertheless, for any statistical operator there is an orthonormal eigenbasis $|r, \lambda\rangle$. However, the spectral decomposition

$$
\begin{equation*}
\rho=\sum_{r \lambda} r|r, \lambda\rangle\langle r, \lambda| \tag{1.10}
\end{equation*}
$$

has the same form as (1.8), revealing that any statistical operator describes some mixture. In addition, we see that the same mixture is obtained in at least two different ways from pure states: spectral form and (1.8). Consequently, it cannot be reconstructed how the mixture is made, and the question of the pure components of a mixture is in general meaningless. Instead, one can show (see Exercise 1.1) only that the involved pure states are from the range of $\rho$.

Exercise 1.1: Show that the states $\left|\psi_{i}\right\rangle$ of (1.8) span the range of $\rho$.
From the construction it is obvious that the mixtures are most general states of physical systems. In addition, Gleason's theorem [?] shows that in any separable Hilbert space (with dimension greater than 2) the probability measures over the subspaces ${ }^{2}$ biuniquelly correspond to the statistical operators by (1.7). Particularly, pure states are also mixtures with a single homogeneous ensemble involved, i.e. the statistical operator corresponding to the pure state $|\psi\rangle$ is according to (1.8) the ray projector $|\psi\rangle\langle\psi|$. The ray projectors are exactly the intersection of the sets of the projectors and statistical operators: projectors are positive operators, and their trace is the dimension of their range (thus, only the ray projectors are statistical operators); on the other hand, a statistical operator is idempotent if and only if it has one nondegenerate eigenvalue 1 , and the other one 0 , with degeneracy $|\mathcal{S}|-1$ (thus, it is ray projector). For any ray projector $P$ there is a state $|\psi\rangle$ such that $P=|\psi\rangle\langle\psi|$, and all the vectors $\mathrm{e}^{\mathrm{i} \phi}|\psi\rangle$ differing by an arbitrary phase give the same ray projector. This means that the pure physical states are in biunique correspondence to the ray projectors, i.e. idempotent mixtures. We see that the statistical operator description is free of the phase ambiguity characteristic for the vector notation.

The sharp value $a$ of the observable $A$ in the mixed state $\rho$ is by definition obtained when $\operatorname{Tr} P_{a}(A) \rho=1$. Using any basis $|a, \lambda\rangle$ in $\mathcal{S}_{a}(A)$, the eigenprojector becomes $P_{a}(A)=$ $\sum_{\lambda}|a, \lambda\rangle\langle a, \lambda|$, and $1=\operatorname{Tr} P_{a}(A) \rho=\sum_{\lambda}\langle a, \lambda| \rho|a, \lambda\rangle$. Since $\rho$ is positive operator, this means that the trace over the eigenspace $\mathcal{S}_{a}(A)$ is equal to the complete trace, and further that for any vector $|\psi\rangle$ orthogonal to $\mathcal{S}_{a}(A)$ one has $\langle\psi| \rho|\psi\rangle=0$, i.e. that orthocomplement of $\mathcal{S}_{a}(A)$ is a subspace in the null-space $\mathcal{S}_{0}(\rho)$ of $\rho$. In other words, the sharp value $a$ of $A$ means that $\rho$ is some mixture of the eigenstates of $A$ for the eigenvalue $a$. Particularly, if $a$ is nondegenerate, then the mixture $\rho=|a\rangle\langle a|$ is in fact the pure state $|a\rangle$.

For each pure state there is a CSCO for which it is a common eigenvector. Therefore, the measurement of these observables reveal their sharp values. On the contrary, there is not proper mixed state with this property. Therefore, pure states correspond to the

[^1]maximal possible information about system. This can be seen also with help of the entropy, which is defined by:
\[

$$
\begin{equation*}
S=-k \operatorname{Tr} \rho \ln \rho, \tag{1.11}
\end{equation*}
$$

\]

with $k$ being Boltzman constant. Calculating trace in the eigenbasis of $\rho$, with help of the spectral form (1.10) one finds $S=-k \sum_{i} r_{i} \ln r_{i}$. Since $1 \geq r_{i} \geq 0$, this expression is always nonnegative. It takes its minimal value $S=0$ for all the states with a single nonvanishing (and automatically nondegenerate) eigenvalue $r_{1}=1$. Obviously, these states are pure states. In accordance with the fact that the pure states are maximally prepared ones, this shows consistency of the above definition of entropy.

In the equilibrium statistical physics of special importance is canonical distribution described by the canonical state

$$
\begin{equation*}
\rho_{\mathrm{can}} \stackrel{\text { def }}{=} Z^{-1} \mathrm{e}^{-H / k T}, \quad Z=\operatorname{Tr}^{-H / k T} \tag{1.12}
\end{equation*}
$$

Here, $T$ is temperature of the ensemble, $H$ is hamiltonian (operator of energy) and $Z$ is called statistical sum or partition function.

### 1.1.7 Measurement

Since all notions of quantum formalism are based on the results of measurements, a precise consideration of the measurement process itself is necessary.

Intuitively, measurement process is a series of acts of interactions of the apparatus with the members of ensemble, each of them resulting in a particular position of the pointer of apparatus. Such positions correspond to various eigenvalues of the measured observable, and the distribution of these outcomes over ensemble coincides with the postulated quantum prediction (1.4). Besides these characteristics which must be satisfied by any measurement device, various measurements may affect measured system in different ways (even destroy it!) depending on the type of the involved interaction. However, it is convenient to extract only the minimal process leading to the results, and ignore all other details of the particular measurement.

According to the instant to which the results refer, measurements may be classified into retrospective and predictive. In the former one, the obtained results are valid immediately before measurement, giving no information about the future (and even existence) of the system. On the contrary, a predictive measurement is not related to the history of the system, and the obtained results refer to the forthcoming instant. Of course, a measurement may be simultaneously retrospective and predictive. Particularly important predictive measurements are selective measurements, enabling to separate subensembles corresponding to different outcomes. Such measurements are unavoidable in the preparations of the ensembles with wanted properties. Also, this type of measurement is used in the following idealization, necessary to understand the mentioned minimal change that ensemble undergoes in any measurement.

Definition 1.1 Ideal measurement of the observable $A$ is an interaction of the apparatus with ensemble in the arbitrary state $|\psi\rangle$ such that:
(i) Each eigenvalue a of $A$ is obtained with the quantum probability:

$$
v(a, A,|\psi\rangle)=\langle\psi| P_{a}(A)|\psi\rangle ;
$$

(ii) Immediately after the measurement the subensemble selected by the eigenvalue a has sharp value a of $A$;
(iii) If any observable $B$ compatible with $A$ had sharp value $b$ in $|\psi\rangle$, this sharp value is retained immediately after the measurement.

In fact, the requirement (i) assures that the considered process is really a quantum measurement, while (ii) formalize the intuitively necessary property of selective measurement: in the repeated measurement of $A$ the selected subensembles show only the value by which they are selected. However, (iii) enables to follow the change of state during the ideal measurement.

Theorem 1.1 (Collapse of the state, Projection postulate) The ideal measurement of the observable $A=\sum_{a} a P_{a}(A)$ is performed. Then:
(i) If the initial state of the system is $|\psi\rangle$, the state of the subensemble corresponding to the selective measurement of the eigenvalue $a$ is

$$
\begin{equation*}
\left|\psi_{a}^{\prime}\right\rangle=P_{a}(A)|\psi\rangle / \sqrt{\langle\psi| P_{a}(A)|\psi\rangle} . \tag{1.13a}
\end{equation*}
$$

(ii) If the system is initially in the state $\rho$, its state immediately after the (nonselective) measurement is

$$
\begin{equation*}
\rho^{\prime}=\sum_{a} P_{a}(A) \rho P_{a}(A) \tag{1.13b}
\end{equation*}
$$

-Proof: (i) We expand the initial state in the form $|\psi\rangle=\sum_{a} P_{a}(A)|\psi\rangle$, and use the normalized components $P_{a}(A)|\psi\rangle / \sqrt{\langle\psi| P_{a}(A)|\psi\rangle}$ to construct the projector $Q=\sum_{a} \frac{P_{a}(A)|\psi\rangle\langle\psi| P_{a}(A)}{\langle\psi| P_{a}(A)|\psi\rangle}$. Note that the denominator is $v(a, A,|\psi\rangle)$. Obviously, $P_{a}(A) Q=Q P_{a}(A)=\frac{P_{a}(A)|\psi\rangle\left\langle\langle | P_{a}(A)\right.}{v(a, A,|\psi\rangle)}$, i.e. $Q$ commutes with each spectral projector of $A$, and the product is the ray projector on the corresponding component of $|\psi\rangle$. The first conclusion also means that $Q$ is compatible with $A$. Further, in the initial state it has sharp value 1: $v(1, Q,|\psi\rangle)=\sum_{a} \frac{\left.\langle\psi| P_{a}(A)|\psi\rangle\langle\psi| P_{a}(A)|\psi\rangle\right\rangle}{v(a, A,|\psi\rangle)}=1$. According to the third characteristics of the ideal measurement, after the ideal measurement $Q$ retains this sharp value, meaning that in each subensemble its eigenvalue 1 value is detected with certainty. Particularly, in the subensemble selected by the eigenvalue $a$ of $A$, it has this sharp value, as well as $P_{a}(A)$. Altogether, this subensemble has sharp value 1 of $Q P_{a}(A)$. As this is the ray projector, the corresponding state is necessarily (1.13a).
(ii) Let $\rho=\sum_{i} \omega_{i}\left|\psi_{i}\right\rangle\left\langle\psi_{i}\right|$ be one of the decompositions of the mixed state onto the pure components. This means that the ensemble $\rho$ one can be divided into the subensembles $\left|\psi_{i}\right\rangle$, each weighted by $\omega_{i}$. After the selective measurements of different eigenvalues of $A$ on the pure components $\left|\psi_{i}\right\rangle$, one gets the subensembles $\frac{P_{a}(A)\left|\psi_{i}\right\rangle}{\sqrt{v\left(a, A, \mid \psi_{i}\right)}}$ for each $a$ and each $i$. The number of subsystems in each of these subensembles is $\omega_{i} v\left(a, A,\left|\psi_{i}\right\rangle\right) N$, where $N$ is the total number of systems in the ensemble $\rho$ and $v\left(a, A,\left|\psi_{i}\right\rangle\right)=\left\langle\psi_{i}\right| P_{a}(A)\left|\psi_{i}\right\rangle$. Hence, the resulting state is the mixture of all the subensembles, with weights $\omega_{i} v\left(a, A,\left|\psi_{i}\right\rangle\right)$, i.e. $\rho^{\prime}=\sum_{i} \omega_{i} \sum_{a} v\left(a, A,\left|\psi_{i}\right\rangle\right) \frac{P_{a}(A)\left|\psi_{i}\right\rangle\left\langle\psi_{i}\right| P_{a}(A)}{v\left(a, A,\left|\psi_{i}\right\rangle\right)}=\sum_{a} P_{a}(A) \rho P_{a}(A)$.

The theorem determines the change of state during the ideal measurement. Any other nonideal measurement may be understood as the minimal one followed by some dynamical process which is essentially irrelevant for the results, but cannot be avoided due to the specific interaction of the system with concrete equipment.

Exercise 1.2: The space of all linear operators in $\mathcal{S}$ is the direct product $\mathcal{S} \otimes \mathcal{S}^{*}$, with naturally defined scallar product $(A, B) \stackrel{\text { def }}{=} \operatorname{Tr} A^{\dagger} B$. This space is usually called superspace, while the operators acting in it are called superoperators. Show that for each $A$ the mapping $\Pi_{A}: \rho \rightarrow \rho^{\prime}$ defined by (1.13b) is a superprojector.

Exercise 1.3: Find the states that are not changed in the ideal measurement of $A$.
The subensemble resulting in the selective measurement of an eigenvalue of an observable (or some set of compatible observables) is homogeneous for any pure or mixed measured state, if and only if this eigenvalue is nondegenerate. Such measurements are called complete.

### 1.1.8 Interference: Superposition and Mixture

Now the formalism is built up sufficiently to enable interpretation of the phenomena observed in the double-slit experiments. The source has prepared state with a sharp value of non-degenerate observable $A$, so that the emitted beam is a pure state $|\psi\rangle$. However, screen with slits serve as a measuring device of an observable $B$, with two emphasized eigenvalues (corresponding to slits). This observable is again assumed complete thus the slits correspond to pure eigenstates $\left|\psi_{1}\right\rangle$ and $\left|\psi_{2}\right\rangle$. However, $A$ and $B$ are not compatible, and particularly, initial state is not eigenstate of $B$, but has equal projections $\left\langle\psi_{1} \mid \psi\right\rangle=$ $\left\langle\psi_{2} \mid \psi\right\rangle$, meaning by Eq. (1.6) that the probabilities to detect two paths after screen are equal. Finally, we measure position of the particles, which means the third complete observable $X$.

Thus, in the case of switched off $\mathrm{D}_{1}$ and $\mathrm{D}_{2}$, the state after screen is superposition $\left|\psi^{\prime}\right\rangle=\left(\left|\psi_{1}\right\rangle+\left|\psi_{2}\right\rangle\right) / \sqrt{2}$. The resulting probability (1.4) is
$v\left(x, X,\left|\psi^{\prime}\right\rangle\right)=\frac{\left|\left\langle x \mid \psi_{1}\right\rangle+\left\langle x \mid \psi_{2}\right\rangle\right|^{2}}{2}=\frac{\psi_{1}^{*}(x) \psi_{1}(x)+\psi_{2}^{*}(x) \psi_{2}(x)+\psi_{1}^{*}(x) \psi_{2}(x)+\psi_{2}^{*}(x) \psi_{1}(x)}{2}$,
When the detectors are switched on, each chosen path effectively makes a subensemble $\left(\left|\psi_{1}\right\rangle\right.$ or $\left.\left|\psi_{2}\right\rangle\right)$, and the state after screen is, according to (1.13b), mixture: $\rho=\left(\left|\psi_{1}\right\rangle\left\langle\psi_{1}\right|+\right.$ $\left.\left|\psi_{2}\right\rangle\left\langle\psi_{2}\right|\right) / 2$. In this case (1.7) gives

$$
v(x, X, \rho)=\psi_{1}^{*}(x) \psi_{1}(x)+\psi_{2}^{*}(x) \psi_{2}(x) .
$$

Thus we see that $v_{1 \nsim 2}$ is obtained directly in this case, while in the previous one there are additional two terms describing the interference.

Of course these results may be easily generalized. Let $|\psi\rangle=\sum_{i} \psi_{i}|i\rangle$. Then, the probability to detect value $a$ of the observable $A$ is

$$
\begin{equation*}
v(a, A,|\psi\rangle)=\sum_{i}\left|\psi_{i}\right|^{2} v(a, A,|i\rangle)+\sum_{i \neq j} \psi_{i} \psi_{j}^{*}\langle a \mid i\rangle\langle j \mid a\rangle . \tag{1.14}
\end{equation*}
$$

The first sum is exactly the probability obtained with help of the mixture $\rho=\sum_{i}\left|\psi_{i}\right|^{2}|i\rangle\langle i|$, while the second describes the interference. Also, $v(a, A,|\psi\rangle)=\left|\sum_{i} \psi_{i}\langle a \mid i\rangle\right|^{2}$ shows that the probability amplitude $\langle a \mid \psi\rangle$ is superposition of the contributing probability amplitudes $\langle a \mid i\rangle$. In other words, when the path is not known, the probability amplitudes of the possible paths are superposed, while for known paths we combine their probabilities.

When a mixed state $\rho=\sum_{k} \omega_{k}\left|\psi_{k}\right\rangle\left\langle\psi_{k}\right|$ is expressed as matrix in the eigenbasis $|a\rangle$ of an arbitrary complete observable $A$, then the matrix elements $\rho_{a b}=\sum_{k} \omega_{k}\left\langle a \mid \psi_{k}\right\rangle\left\langle\psi_{k} \mid b\right\rangle$ relate the mixed pure states to the measurement of $A$. At first, the diagonal elements $\rho_{a a}=\sum_{k} \omega_{k} v\left(a, A,\left|\psi_{k}\right\rangle\right)$ obviously show probability to detect system in the state $|a\rangle$, with clear contributions from the mixed pure states. Therefore they are called populations (of the states $|a\rangle$ ). The nondiagonal elements, according to (1.14), are the averaged (over the mixed pure components) interference terms of the states $|a\rangle$ and $|b\rangle$ for the system being in each of the pure component $\left|\psi_{k}\right\rangle$. Thus their nonvanishing values reveal that although the state is mixture, there is some remained interference from the coherent subensembles; this explains why these elements are called coherences.

### 1.2 Quantization and State Spaces

### 1.2.1 Canonical Quantization

Classical description of any physical system is easily established, despite possibly unsatisfactory predictions. It is therefore natural to attempt to build the quantum description using the same ideas, or more precisely, to incorporate somehow the classical notions within quantum formalism, i.e. to find an algorithm to construct quantum formalism starting from the classical theory. To this end, the notions of the classical states and variables should be carefully reconsidered, and compared to their quantum analogues.

Within classical theory a system is described in terms of phase space and the space of variables. The states are points in the phase space, which is a manifold factorized onto the configuration space (generally, this is not vector space) and the momentum space. Usually, the dimensions of the both factors are the same, and proportional to the number of the particles in the system (mathematically precise, the phase space is cotangent bundle of the configuration manifold). Any physical quantity, called classical variable, is a function over the state space, $f(\boldsymbol{q}, \boldsymbol{p})$. Thus, all the variables are functions of the coordinates and momenta, which are therefore called fundamental variables. In accordance with the profound continuity principle of physics, these functions are differentiable. Note that in each point of the phase space, all the variables have well defined values, or, using quantum vocabular, each variable $f$ has the sharp value $f(\boldsymbol{q}, \boldsymbol{p})$ in any classical state. Therefore, for each pair of classical states there are variables having different values in them, i.e. distinguishing between these states. A particular choice of such variables is the fundamental set $\{\boldsymbol{q}, \boldsymbol{p}\}$.

Structurally, the variables form a vector space, since linear combination of differentiable functions is also differentiable. In addition, pointwise multiplication $h(\boldsymbol{q}, \boldsymbol{p}) \stackrel{\text { def }}{=}$ $f(\boldsymbol{q}, \boldsymbol{p}) g(\boldsymbol{q}, \boldsymbol{p})$ of variables defines a commutative multiplication $h=f g$, and the space of variables becomes an abelian associative algebra. It is also a Lie algebra, with the Poisson
bracket $[a, b] \stackrel{\text { def }}{=} \sum_{i}\left(\frac{\partial a}{\partial q_{i}} \frac{\partial b}{\partial p_{i}}-\frac{\partial a}{\partial p_{i}} \frac{\partial b}{\partial q_{i}}\right)$ as the Lie product.

On the other side, the set of quantum observables is real space of all the hermitian operators acting in the state space $\mathcal{S}$. It is associative algebra with respect to the operator multiplication (composition). However, the crucial difference is noncommutativity of this algebra, implying that no state is a common eigen state for each observable, i.e. there is no state with a sharp value for any observable. To understand the difficulty and get a hint how to override it, we start with coordinate. If the corresponding observable $\hat{q}$ has to retain the basic idea of the classical coordinate, it must have the continual spectrum $\sigma(q)=\mathbb{R}$. The corresponding eigenvectors $|q\rangle$ (i.e. $\hat{q}|q\rangle=q|q\rangle$ ), being the states with the sharp value $q$ of $\hat{q}$, resemble to the classical states with well defined coordinate. However, there a is continuum of the classical states with sharp value $q$, distinguished by the value of the momentum. In analogy with the phase space, one may try to assume that each eigenvalue $q$ of $\hat{q}$ is degenerate, and that all the degenerate vectors should be labelled by an additional value of the momentum. Accordingly, the complete characterization of these eigenvectors would be $|q, p\rangle$, with the simultaneously satisfied $\hat{q}|q, p\rangle=q|q, p\rangle$ and $\hat{p}|q, p\rangle=q|q, p\rangle$, implying that $\hat{q}$ and $\hat{p}$ commute. But this contradicts to the experiments, showing that well collimated beam of particles, although having more or less sharp value of momenta, is far from having sharp position and classical trajectory. Thus, the idea of definition of the state by the values of the observables on them must be weaken: we have to build the state space with either coordinate or momentum eigenvectors only. This opens the question if these observables are complete.

Further, it seems natural to retain the functional relations between the classical variables in the quantum formalism, since these are well established experimentally and therefore formalism independent. In fact, this cannot be avoided, since the variables, like the angular momentum, are even defined as the functions of the fundamental ones. Thus, the quantum observables that are meaningful in the classical physics (like angular momentum, energy, etc.; in the Section 4.5, spin as a nonclassical degree of freedom, will be discussed) should be the same functions of the observales of coordinates and momenta. However, the product of the hermitian operators is hermitian if and only if they commute, implying that incompatibility of the observables suppress this simple analogy. In addition, commutation relations between the operators has already proved to be essential in the view of the simultaneous measurements. Since the commutator (multiplied by imaginary unit) is a Lie product in the space of observables, it has to be related to the classical Lie product, i.e. to the Poisson bracket.

Thus, the quantization procedure has to resolve the described problems, but also to provide the uniqueness of the quantum formalism.

## POSTULATE 4. - QUANTIZATION

The quantum description of a physical system is obtained from the classical one by the following correspondence rules:

1. Linear combination of the variables corresponds to the same linear combination of the observables;
2. The product $A B$ of two variables corresponds to the symmetrized product $(\hat{A} \hat{B}+\hat{B} \hat{A}) / 2$ of the observables;
3. The quantization is continual: any series $\left\{\hat{A}_{i}\right\}$ of observables corresponding to the convergent series $\left\{A_{i}\right\}$ of variables with limit $A$, is convergent with the limit $\hat{A}$ being the quantized variable $A$.
4. The basic set of variables corresponds to the basic set of observables, in a way that Poisson bracket $[A, B]_{P Z}$ of any variable $A$ with each basic variable $B$ becomes the commutator $\frac{1}{\mathrm{i} \hbar}[A, B]$.

The first rule is obvious, while the second one introduces the minimal change of the product to the symmetrized product, to regain the hermitian observable. Next rule, asserting that the limit and the quantization are mutually commuting operations, enables to quantize the variables which are transcendent functions of the fundamental set, using their series over the fundamental set.

The last rule is more subtle. Its role is to enable algebraic construction of the state space. To this end, the fundamental set of observables should be defined first, and the set of observables of coordinates and momenta is a natural choice by the analogy with the classical mechanics. Though this is not sufficient for the theories involving nonclassical degrees of freedom (such as spin or other interior degrees to be introduced much later), as far as quantization of the classical theories is considered, this approach is effective. Namely, in such situations all quantum observables are functions of coordinates and momenta. Accordingly, possible degeneracy of the observable of the coordinate would introduce another observables completing it to a CSCO. However, the last rule requires that classical analog $A$ of an opservable $\hat{A}$ from such a CSCO, thus compatible with $\hat{Q}$, satisfies $[A, Q]_{P Z}=-\frac{\mathrm{d} A}{\mathrm{~d} P}=0$. This means that $A$ is a function of $Q$ only, $A=A(Q)$, and therefore $\hat{A}$ is the same function $\hat{A}(\hat{Q})$ of $\hat{Q}$, which is a contradiction: no function of $\hat{Q}$ can lower the degeneracy of $\hat{Q}$. Imaginary unit in front of the commutator has been anticipated as a way to obtain the hermitian Lie product, while the Planck's constant $\hbar$ with the dimension of action is necessary to compensate the dimension of the product of the coordinate and the momentum in the Poisson bracket. Its value $\hbar=1.0545710^{-34} \mathrm{Js}$ is determined by many very precise experiments. Note that for the coordinates and the momenta this reads:

$$
\begin{equation*}
\left[\hat{Q}_{i}, \hat{Q}_{j}\right]=0, \quad\left[\hat{P}_{i}, \hat{P}_{j}\right]=0, \quad\left[\hat{Q}_{i}, \hat{P}_{j}\right]=\mathrm{i} \hbar \delta_{i j} \tag{1.15}
\end{equation*}
$$

In general, canonically conjugate observables $\hat{A}$ and $\hat{B}$ satisfy $[\hat{A}, \hat{B}]=\mathrm{i} \hbar$.

Nevertheless, to enable later inclusion of the nonclassical degrees of freedom, the fundamental set of observables is to be defined in a more general way.
Definition 1.2 The fundamental set is any set $F=\left\{\hat{A}_{1}, \ldots, \hat{A}_{f}\right\}$ of observables such that:

1. the state space $\mathcal{S}$ is irreducible with respect to $F$, i.e. besides the trivial ones in $\mathcal{S}$ there is no other common invariant subspaces;
2. There is at least one subset in F being CSCO, such that the whole common eigen basis of this CSCO can be generated from a single common eigen vector by the functions of the fundamental observables;
3. the space $\mathcal{S}$ in which such fundamental set is defined is unique up to isomorphism.

The first requirement is necessary since if the fundamental set is reducible with $\mathcal{S}_{1}$ and $\mathcal{S}_{2}$ being the common invariant subspaces, then the formalism can be completely developed in the two invariant subspaces independently. More striking is the consequence that the dynamics of the system, i.e. the change of the state during the time would be restricted to these subspaces, meaning that there would be in fact two physical descriptions of the same system. This intuitive paradox is prevented by the last requirement.

The second requirement is essential for the construction of the state space. This will be enlighten in the next section, when the state spaces of the elementary systems will be explicitly constructed as the Lebesgue space over the classical configuration space, with help of the family of the operator functions mapping one coordinate eigen vector into the others.

CORRESPONDENCE Q2C FOR $\hbar \rightarrow 0$

### 1.3 Construction of the state space

In this section the postulate of quantization will be utilized to construct the state space of the systems with classical analogs. Here come systems with $N$ particles in $D$ dimensions.

### 1.3.1 Single one-dimensional particle

This simplest system classically is described by the two-dimensional phase space with the coordinates $\{x, p\}$ being the fundamental variables with the commutators $[x, x]_{\mathrm{PZ}}=$ $[p, p]_{\mathrm{PZ}}=0$ and $[x, p]_{\mathrm{PZ}}=1$. According to the quantization postulate, in the quantum formalism there is unique state space $\mathcal{H}$ with the opservables of coordinate $\hat{x}$ and momentum $\hat{p}$, such that $[\hat{x}, \hat{x}]=[\hat{p}, \hat{p}]=0$ and $[\hat{x}, \hat{p}]=\mathrm{i} \hbar$.
Exercise 1.4: Show that the differential operator $p=-\mathrm{i} \hbar \frac{\mathrm{d}}{\mathrm{d} x}$ is linear and Hermitian in the space of all differentiable wavefunctions, which vanish at both ends of an interval $(a, b)$.

Exercise 1.5: Show that $[x, p]=\mathrm{i} \hbar,\left[x^{n}, p\right]=\mathrm{i} \hbar n x^{n-1},[A, p]=\mathrm{i} \hbar \frac{d A}{d x},\left[x, p^{n}\right]=\mathrm{i} \hbar n p^{n-1},[x, B]=\mathrm{i} \hbar \frac{d B}{d p}$, where $p=-\mathrm{i} \hbar \frac{\mathrm{d}}{\mathrm{d} x}, A=A(x)$ and $B=B(p)$ are differentiable functions in $x$ and $p$, respectively.

An immediate consequence of these relations is:

Theorem 1.2 The observable $\hat{x}$ has pure continual spectrum $C \sigma(\hat{x})=\mathbb{R}$.

- Proof: Let us define the one-parameter family of the unitary operators $\hat{U}(q) \xlongequal{\text { def }} \mathrm{e}^{-\frac{1}{\hbar} q \hat{p}}$, for each real $q$. Using the exercise 1.5 , it is easy to show that $[\hat{x}, \hat{U}(q)]=\sum_{n=0}^{\infty} \frac{\left(-\frac{i}{\hbar} q\right)^{n}}{n!}\left[\hat{x}, \hat{p}^{n}\right]=\sum_{n=0}^{\infty} \frac{\left(-\frac{i}{\hbar} q\right)^{n} n i \hbar}{n!} \hat{p}^{n-1}=$ $q \hat{U}(q)$. Further, if $x \in \mathbb{R}$ is in the continual spectrum of $\hat{x}$, then the (generalized) eigen vector $|x\rangle$ exists such that $\hat{x}|x\rangle=x|x\rangle$. But then, $\hat{x}(\hat{U}(q)|x\rangle)=([\hat{x}, \hat{U}(q)]+\hat{U}(q) \hat{x})|x\rangle=\hat{U}(q)(q+\hat{x})|x\rangle=$ $(x+q)(\hat{U}(q)|x\rangle)$. This means that $\hat{U}(q)|x\rangle$ is also eigen vector for the eigenvalue $x+q$. Therefore, if any real number $x$ is in the spectrum, then all reals are also in the spectrum, since $q$ is arbitrary.

This is in accordance with the intuitive notion of coordinate (and momentum also): the sharp value $x$ of the particle coordinate, i.e. completely localized particle, should be described by the vector $|x\rangle$, corresponding thus to the continual spectrum. Note the appearance of the translational operator family $\hat{U}(q)$, which, as mentioned in the definition of the FSO generates the whole eigen basis of $\hat{x}$ from each of its vectors. In fact, the nonuniqueness of the phase factor is resolved by convention

$$
\begin{equation*}
|x\rangle=U(x)|0\rangle, \forall x \in \mathbb{R} \tag{1.16}
\end{equation*}
$$

Hence, one can start with the basis $|x\rangle$, make the space over it, and check the other parts of the quantization procedure. At first, recall that $\hat{x}$ is complete observable; if not, it could be completed to CSCO, but then the canonical commutator rules would allow only functions of $\hat{x}$ as the compatible with $\hat{x}$ observables. And such observables cannot be used to complete $\hat{x}$. Further, the uniqueness of the obtained space follows from the fact that the Heissenberg Lie algebra (vector space over $\{\hat{x}, \hat{p}, \hat{I}\}$ with the commutators $[\hat{x}, \hat{p}]=$ const and $[\hat{x}, \hat{I}]=[\hat{I}, \hat{p}]=0$ ) is solvable, meaning that its irreducible representations are either one-dimensional or infinite dimensional one.

## Coordinate representation

Thus, the task of the construction of the space is fulfilled. Each vector can be expanded over the obtained basis $|\psi\rangle=\int_{\mathbb{R}} \psi(x)|x\rangle \mathrm{d} x$. The corresponding Fourier coefficients $\psi(x)=\langle x \mid \psi\rangle$ give the coordinate representation of the state, which is called wave function. Instead of the point in the phase space, the state is therefore described by the functions. The density of probability to find the particle at $x$ is $\psi^{*}(x) \psi(x)$. The finite norm functions, i.e. the functions over $\mathbb{R}$ satisfying $\int_{\mathbb{R}}|\psi|^{2} \mathrm{~d} x<\infty$ form the Lebesgue space $\mathcal{L}(\mathbb{R})$. Note that the vectors $|x\rangle$ are not from this space, since their coordinate representation is Dirac $\delta$-function: $\left|x^{\prime}\right\rangle=\int_{\mathbb{R}} \delta\left(x-x^{\prime}\right)|x\rangle \mathrm{d} x$. In the coordinate representation the observable $\hat{x}$ becomes the multiplicative operator: since $\hat{x}|\psi\rangle=\int_{\mathbb{R}} \psi(x) \hat{x}|x\rangle \mathrm{d} x=\int_{\mathbb{R}}(x \psi(x))|x\rangle \mathrm{d} x$, the function $\psi(x)$ is mapped by $\hat{x}$ to $x \psi(x)$. Therefore, the domain of the coordinate $\hat{x}$ is the subspace in $\mathcal{L}(\mathbb{R})$ with the functions satisfying that also $\int_{\mathbb{R}} x^{2}|\psi|^{2} \mathrm{~d} x<\infty$. Finally, since $\hat{p}=\mathrm{i} \hbar\left[\frac{\mathrm{d} \hat{U}(q)}{\mathrm{d} q}\right]_{q=0}$, the coordinate representation of $\hat{p}$ is found to be the derivative operator

$$
\begin{equation*}
\hat{p}=-\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} x} . \tag{1.17}
\end{equation*}
$$

Namely, $\hat{p}|\psi\rangle=\mathrm{i} \hbar\left[\frac{\mathrm{d} \hat{U}(q)}{\mathrm{d} q}\right]_{q=0} \int_{\mathbb{R}} \psi(x)|x\rangle \mathrm{d} x=\mathrm{i} \hbar\left[\frac{\mathrm{d}}{\mathrm{d} q} \int_{\mathbb{R}} \psi(x)|x+q\rangle \mathrm{d} x\right]_{q=0}=$ $\mathrm{i} \hbar \int_{\mathbb{R}}\left[\frac{\mathrm{d} \psi(x-q)}{\mathrm{d} q}\right]_{q=0}|x\rangle \mathrm{d} x=-\mathrm{i} \hbar \int_{\mathbb{R}} \frac{\mathrm{d} \psi(x)}{\mathrm{d} x}|x\rangle \mathrm{d} x$. Its domain is comprised by the functions
from $\mathcal{L}(\mathbb{R})$ for which also the image is in $\mathcal{L}(\mathbb{R})$, i.e. such that $\int_{\mathbb{R}}\left|\frac{d \psi(x)}{d x}\right|^{2} d x<\infty$.

## Momentum representation

The eigenproblem of the momentum operator (1.17) is easily solved. Its spectrum is continual and nondegenerate, $C \sigma(\hat{p})=\mathbb{R}$, with the corresponding eigenvectors (generalized, distributions) normalized to Dirac's $\delta$-function

$$
\begin{equation*}
|p\rangle=\int_{\mathbb{R}}\langle x \mid p\rangle|x\rangle \mathrm{d} x, \quad\langle x \mid p\rangle=\psi_{p}(x)=\frac{1}{\sqrt{2 \pi}} \mathrm{e}^{\mathrm{i} p x}, \quad\left\langle p \mid p^{\prime}\right\rangle=\delta\left(p-p^{\prime}\right) \tag{1.18}
\end{equation*}
$$

known as plane waves.
As the spectrum is nondegenerate, momentum operator is itself a CSCO, enabling to represent vectors and operators in momentum representation. As for the states, they are again represented by functions, $\psi(p)=\langle\psi \mid p\rangle$. While the momentum operator in this representation is the multiplicative operator, $\hat{p} \psi(p)=p \psi(p)$, the coordinate operator is the derivative, $\hat{x} \psi(p)=\mathrm{i} \hbar \frac{\mathrm{d}}{\mathrm{d} x} \psi(p)$.

Clearly, the form of the plane waves reveals that the transition between coordinate and momentum representation is realized by the Fourier transformation:

$$
\begin{equation*}
\psi(p)=\frac{1}{\sqrt{2 \pi}} \int_{\mathbb{R}} \mathrm{e}^{-\mathrm{i} p x} \psi(x) \mathrm{d} x, \quad \psi(x)=\frac{1}{\sqrt{2 \pi}} \int_{\mathbb{R}} \mathrm{e}^{\mathrm{i} p x} \psi(p) \mathrm{d} p . \tag{1.19}
\end{equation*}
$$

### 1.3.2 Three dimensional particle, many particles

According to the Fatu-Fubini theorem, Lebesgue space over $\mathbb{R}^{n}$ is direct product of the Lebesgue spaces over each $\mathbb{R}: \mathcal{L}(\mathbb{R} \oplus \cdots \oplus \mathbb{R})=\mathcal{L}(\mathbb{R}) \otimes \cdots \otimes \mathcal{L}(\mathbb{R})$. Thus, each degree of freedom described by the real axis contributes by its Hilbert space $\mathcal{L}(\mathbb{R})$, and the total state space is product of these. For the discrete degrees of freedom, like spin, the corresponding space is finite dimensional.

This gives straightforward algorithm to construct state spaces of arbitrary systems. For a single particle, as its configurational space is three-dimensional (corresponding to three degrees of freedom) state space is $\mathcal{S}=\mathcal{L}\left(\mathbb{R}^{3}\right)=\mathcal{L}(\mathbb{R}) \otimes \mathcal{L}(\mathbb{R}) \otimes \mathcal{L}(\mathbb{R})=\mathcal{L}^{3}(\mathbb{R})$. Analogously, for $N$ such particles, the state space becomes $\mathcal{S}=\mathcal{L}\left(\mathbb{R}^{3 N}\right)=\mathcal{L}^{3 N}(\mathbb{R})$.

In analogy to the single one-dimensional particle, according to (1.15), there is common eigenbasis $\left|Q_{11}, \ldots, Q_{13}, \ldots, Q_{N 3}\right\rangle$ of all the coordinate operators $Q_{n i}(n=1, \ldots, N, i=$ $1,2,3)$. This basis defines the coordinate representation. Momentum operators of the $n$-th particle, expressed in the convenient vector form, are represented as:

$$
\boldsymbol{p}_{n}=-\mathrm{i} \hbar \boldsymbol{\nabla}_{n}=-\mathrm{i} \hbar\left(\frac{\partial}{\partial Q_{n x}} \boldsymbol{e}_{x}+\frac{\partial}{\partial Q_{n y}} \boldsymbol{e}_{y}+\frac{\partial}{\partial Q_{n z}} \boldsymbol{e}_{z}\right) .
$$

Exercise 1.6: The translation operator $U(\boldsymbol{a})$ is defined to be such that $U(\boldsymbol{a}) \psi(\boldsymbol{r})=\psi(\boldsymbol{r}-\boldsymbol{a})$. Show that $U(\boldsymbol{a})$ is unitary and express it in terms of $\boldsymbol{p}=-\mathrm{i} \hbar \boldsymbol{\nabla}$.

### 1.3.3 Relative particles

### 1.3.4 Subsystems: reduced states and correlation

As discussed in subsection 1.3.2, the system consisted of two subsystems (particles, degrees of freedom like coordinates) is described by the state space $\mathcal{S}=\mathcal{S}_{1} \otimes \mathcal{S}_{2}$, where $\mathcal{S}_{1}$ and $\mathcal{S}_{2}$ are the subsystems' state spaces. This space is spanned by the states being direct products of the single-particle states, which are called non-correlated or separable states, to distinguish from the correlated states, being their nontrivial linear combinations.

It is intuitively plausible that when such composite system is in the non-correlated state $|\psi\rangle=\left|\psi^{(1)}\right\rangle \otimes\left|\psi^{(2)}\right\rangle$, the subsystems are in the factor states $\left|\psi^{(1)}\right\rangle$ and $\left|\psi^{(2)}\right\rangle$, respectively. However, it is not obvious what is the subsystem state in the case of the correlated composite state.

As usual, these questions will be clarified through analysis of measurements. Particularly, as a state of a subsystem is manifested only in subsystem measurements, it should give correct results (1.4) in all the measurements of the first subsystem observables $A^{(1)}$. In the composite space, such observables induce composite observables $A^{(1)} \otimes \mathbf{1}^{(2)}$. Suppose that $A^{(1)}$ is a complete observable in $\mathcal{S}^{(1)}$, with the eigenbasis $|a\rangle$, while $|b\rangle$ is an arbitrary basis in $\mathcal{S}^{(2)}$. Then, any composite state $|\psi\rangle$ may be expanded over non-correlated basis $|a\rangle \otimes|b\rangle$ as

$$
\begin{equation*}
|\psi\rangle=\sum_{a b} \psi_{a b}|a\rangle \otimes|b\rangle=\sum_{a}|a\rangle \otimes\left|\psi_{a}\right\rangle, \quad\left|\psi_{a}\right\rangle=\langle a \mid \psi\rangle=\sum_{b} \psi_{a b}|b\rangle \tag{1.20}
\end{equation*}
$$

Here, $\langle a \mid \psi\rangle$ is the partial scalar product of the composite state $|\psi\rangle$ and the first subsystem state $|a\rangle$, giving a state of the second subsystem. In the selective measurement of $A^{(1)} \otimes \mathbf{1}$, the state of the subensemble corresponding to the pointer value $a$ is obtained by the projection with $P_{a}^{(1)} \otimes \mathbf{1}$, since no requirement on the second subsystem is imposed. The result is $|a\rangle \otimes\left|\psi_{a}\right\rangle$, i.e. the product of the chosen eigenvector with the partial scalar product of the state with that eigenvector.

Of course, the partial scalar product may be defined also with the states of the second system. Using it, one further defines the partial trace of an operator. In fact, if the composite operator has in the Dirac notation the form $A=\sum_{a a^{\prime} b b^{\prime}} A_{a^{\prime} b^{\prime}}^{a b}|a b\rangle\left\langle a^{\prime} b^{\prime}\right|$, than its partial traces over two spaces, called the reduced operators, are:

$$
\begin{align*}
& A_{1}=\operatorname{Tr}_{2} A=\sum_{b}\langle b| A|b\rangle=\sum_{a a^{\prime}}\left(\sum_{b} A_{a^{\prime} b}^{a b}\right)|a\rangle\left\langle a^{\prime}\right|,  \tag{1.21a}\\
& A_{2}=\operatorname{Tr}_{1} A=\sum_{a}\langle a| A|a\rangle=\sum_{b b^{\prime}}\left(\sum_{a} A_{a b^{\prime}}^{a b}\right)|b\rangle\left\langle b^{\prime}\right| . \tag{1.21b}
\end{align*}
$$

It is easy to show that like the usual trace, the partial ones are independent on the used bases. Further, if $A$ is hermitian, positive or statistical operator, both its partial traces are of the same types.

Exercise 1.7: Show that partial trace of the hermitian operator is hermitian. Analogously for positive and statistical operators.

After this technical preparation, we come back to the question of subsystem states. On the composite system in the most general (mixed) state $\rho$, the subsystem observable $A^{(1)}=\sum_{a} P_{a}^{(1)}$ is measured. The probability of the results is

$$
\begin{array}{r}
v\left(a, A^{(1)} \otimes \mathbf{1}, \rho\right)=\operatorname{Tr}\left(P_{a}^{(1)} \otimes \mathbf{1}\right) \rho=\sum_{a^{\prime} b}\left(\left\langle a^{\prime}\right| \otimes\langle b|\right)\left(P_{a}^{(1)} \otimes \mathbf{1}\right) \rho\left(\left|a^{\prime}\right\rangle \otimes|b\rangle\right) \\
=\sum_{a^{\prime}}\left\langle a^{\prime}\right| P_{a}^{(1)}\left(\sum_{b}\langle b| \rho|b\rangle\right)\left|a^{\prime}\right\rangle=\operatorname{Tr} P_{a}^{(1)} \rho_{1}=v\left(a, A^{(1)}, \rho_{1}\right) . \tag{1.23}
\end{array}
$$

This shows that the reduced density matrix $\rho_{1}$ (and analogously defined $\rho_{2}$ ) completely determines the results of the subsystem measurements. In this sense, the reduced statistical operators describe the states of the subsystem.

It is important to note that even in the case of the pure composite state $|\psi\rangle$ (i.e. $\rho=$ $|\psi\rangle\langle\psi|$, the reduced states are not pure. The obtained mixed states in these cases are called the mixtures of the second kind to distinguishes from the usual mixtures obtained by the mixing of the different ensembles. Only for non-correlated composite states $\left|\psi^{(1)}\right\rangle \otimes\left|\psi^{(2)}\right\rangle$ reduced states are pure states $\rho_{i}=\left|\psi^{(i)}\right\rangle\left\langle\psi^{(i)}\right|$.

When initial state is separated, $\rho^{\prime}=\rho_{1} \otimes \rho_{2}$, the reduced states are just the factors $\rho_{1}$ and $\rho_{2}$ (this includes also the pure states: for $|\psi\rangle=\left|\psi^{(1)}\right\rangle \otimes\left|\psi^{(2)}\right\rangle$, the subsystems are in the factor states $\left|\psi^{(1)}\right\rangle$ and $\left|\psi^{(2)}\right\rangle$ ) as intuitively guessed at the beginning. This is the reason that such states sometimes are called classical.

Note that the reduced operators $\rho_{1}$ and $\rho_{2}$ do not completely describe the whole composite system. Indeed, they do not determine the composite state $\rho$, since different composite states $\rho$ and $\rho^{\prime}$ may have the same both reduced states. For example, for each $\rho$ giving $\rho_{1}$ and $\rho_{2}$, also the separable (or noncorrelated) state $\rho^{\prime}=\rho_{1} \otimes \rho_{2}$ has the same reduced operators. The difference between $\rho$ and $\rho^{\prime}$ reflects the interrelation of the subsystem states in the composite state, and it is erased by the partial traces. In general, this interrelation is called correlation or entanglement, and the information on it is contained in $\rho-\rho^{\prime}$, although there are some other derived quantities which can be used to measure it.

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### 1.4 Uncertainty relations

Every observable has a sharp value in each of its eigenstates, but this is not the case otherwise. Dispersion of the observable is statistical measure of the non-sharpness:

$$
\begin{equation*}
\Delta_{\psi}(A)=\sqrt{\left\langle A^{2}\right\rangle_{\psi}-\langle A\rangle_{\psi}^{2}}=\sqrt{\left\langle\left(A-\langle A\rangle_{\psi}\right)^{2}\right\rangle_{\psi}}, \tag{1.24}
\end{equation*}
$$

vanish in the eigen states of $A$, and the eigen states are the only nondispersive states for $A$. On the other hand, observable $A$ has nonvanishing dispersion in a general state, pointing to the nontrivial probability distribution, i.e. to an uncertainty of the prediction of particular values of the pointer.

For the pair of observables, $A$ and $B$, the product of the dispersions $\Delta_{\psi}(A) \Delta_{\psi}(B)$ may be considered as a measure of the uncertainty of the prediction of the pairs of the two pointers (for $A$ and $B$ ) values in the state $\psi$. Both observables have sharp values only if $\psi$ is their common eigen state. Thus, if they are incompatible, such states may not exist at all. So, it becomes obvious that in general there is an uncertainty of the measurement results of the two observables, and that it is related to their compatibility.

Theorem 1.3 (Heisenberg uncertainty relation) For any state (pure or mixed) $\psi$ and any pair of observables $A$ and $B$ the uncertainty of the probability distributions $v(A, \psi)$ and $v(B, \psi)$ has lower bound:

$$
\begin{equation*}
\Delta_{\psi}(A) \Delta_{\psi}(B) \geq \frac{1}{2}\left|\langle[A, B]\rangle_{\psi}\right| \tag{1.25}
\end{equation*}
$$

- Proof: For a pure state $|\psi\rangle$ define the auxiliary operators $A^{\prime}=A-\langle A\rangle_{\psi}, B^{\prime}=B-\langle B\rangle_{\psi}$, and vectors $|u\rangle=A^{\prime}|\psi\rangle,|v\rangle=B^{\prime}|\psi\rangle$. In the Schwartz inequality $\||u\rangle\| \| \||v\rangle \| \geq|\langle u \mid v\rangle|$, on the left is $\Delta_{\psi}(A) \Delta_{\psi}(B)$, while on the right there is the mean $\left|\left\langle A^{\prime} B^{\prime}\right\rangle_{\psi}\right|=\left|\left\langle\frac{A^{\prime} B^{\prime}+B^{\prime} A^{\prime}}{2}\right\rangle_{\psi}+\left\langle\frac{A^{\prime} B^{\prime}-B^{\prime} A^{\prime}}{2}\right\rangle_{\psi}\right|$. Recall that the obtained symmetrized and antisymmetrized products are hermitian and scew-hermitian operators, and their means are real and pure imaginary, respectively. Thus, the absolute value of the whole expression is greater than that of its imaginary part. Altogether, $\Delta_{\psi}(A) \Delta_{\psi}(B) \geq\left|\left\langle A^{\prime} B^{\prime}\right\rangle_{\psi}\right| \geq \frac{1}{2}\left|\left\langle\left[A^{\prime}, B^{\prime}\right]\right\rangle_{\psi}\right|=\frac{1}{2}\left|\langle[A, B]\rangle_{\psi}\right|$. I
Exercise 1.8: Prove (1.25) for the mixed states.
In particular, canonically conjugated observables $A$ and $B$ satisfy Heisenberg relation:

$$
\begin{equation*}
\Delta_{\psi}(A) \Delta_{\psi}(B) \geq \frac{\hbar}{2} \tag{1.26}
\end{equation*}
$$

The states realizing the minimal uncertainty $|\langle[A, B]\rangle| / 2$ are called minimal wave packets. To find them it should be examined when the inequalities used in the proof are equalities. Thus, the first one, the Schwartz inequality, degenerates into equality if and only if the related vectors are collinear, that is when $A^{\prime}|\psi\rangle=\alpha B^{\prime}|\psi\rangle$. The second inequality becomes an equality when the mean of the symmetrized product vanishes, i.e. when $0=\langle\psi|\left(A^{\prime} B^{\prime}+B^{\prime} A^{\prime}\right)|\psi\rangle=\left(\alpha^{*}+\alpha\right)\left\langle B^{\prime 2}\right\rangle$. Thus, $\alpha$ must be pure imaginary.
Exercise 1.9: Show that the coordinate-momentum minimal wave packet is the Gaussian $\psi(x)=$ $\mathrm{e}^{-\left(\frac{x-a}{2 f}\right)^{2}+\frac{\mathrm{i}}{\hbar} b x}$. What are the $a, b$ and $f$ ?

Using the uncertainty relation, one can estimate delocalization of a free particle having mean energy $\langle H\rangle$. Hamiltonian of such a particle has kinetic part only, $H=\frac{\hat{p}^{2}}{2 m}$. Therefore, for the one-dimensional particle: $\langle H\rangle=\frac{\left\langle\hat{\hat{p}}^{2}\right\rangle}{2 m}=\frac{(\Delta \hat{p})^{2}}{2 m}+\frac{\langle\hat{p})^{2}}{2 m} \geq \frac{\hbar^{2}}{8 m(\Delta \hat{x})^{2}}$, while for three dimensional particle this gives $\langle H\rangle \geq \frac{\hbar^{2}}{8 m}\left(\frac{1}{(\Delta \hat{x})^{2}}+\frac{1}{(\Delta \hat{y})^{2}}+\frac{1}{(\Delta \hat{z})^{2}}\right)$. Obviously, for better localization, more energy is needed.

For commutative observables $\Delta_{\psi}(A) \Delta_{\psi}(B) \geq 0$. In fact, $\Delta_{\psi}(A) \Delta_{\psi}(B)=0$ for the common eigenstates, being nondispersive for both observables, in accordance with the introduced notion of compatibility. In addition, noncommutative observables may have some common eigenstates (but not a common eigen basis!). In such states $\Delta_{\psi}(A) \Delta_{\psi}(B)=$ 0.

## Chapter 2

## Quantum Dynamics

The main task of the physics is to predict events in the future on the basis of the known facts in the past. Unlike to the kinematics, which hardly relate the system itself but only the geometry of the neighboring space, the dynamics dominantly involve the system itself, its structure, mutual interactions of its components and possible external influences.

### 2.1 A classification of the physical systems

The classical mechanics solves the dynamical problem within canonical formalism. The obtained canonical equations involve the hamiltonian variable. It arises as the quantity governing the dynamics: through the variational principle, it singles out the geodesic of the motion. Therefore, to enable the direct canonical quantization, one should understand the domain of validity of the classical canonical formalism, and particularly the conditions enabling the hamiltonian approach. To this end we briefly classify physical systems according to their relationship to the surrounding. In fact, just the possibility to define, i.e. recognize, a particular physical system is neatly related to the interaction of the system with its surrounding.

Starting from the simplest, and increasing the physical complexity of the systems we point out the following classes:
(i) System of free particles: there is no interaction between the subsystems (particles), i.e. classical Hamiltonian is sum of the one particle hamiltonians; the other properties are the same as for the next class.
(ii) Isolated system: although the subsystems interact mutually, they do not interact with the surrounding; thus the hamiltonian involves only system coordinates and momenta, preserving space-time homogeneity.
(iii) Conservative system: it may interact with the surrounding, and the Hamiltonian may involve also some parameters characterizing the surrounding (external field breaking space homogeneity); still it cannot exchange the energy, and the hamiltonian is time independent, i.e. time is homogeneous (e.g. particles in the conservative fields).
(iv) Dynamically separated system: this is a system which interact with the surrounding in a very general manner, but still its dynamic may be singled out from the dynamics of the surrounding; the Hamiltonian is dependent both on the time and on the external field parameters, but the system parameters are not involved in the equation of motion of the surrounding (no feedback), which allows the separation from the surrounding (e.g. electron in the accelerator).
(v) Kinematically separated system: only some degrees of freedom in some large system are used to single out kinematically independent subsystem; the dynamical equations may not be written independently for this part (e.g. atom as a subsystem in a molecule).

Obviously hamiltonian approach may be formulated for the classes (i)-(iv), which are idealized to some extent. While negligible feedback may be very realistic approximation of (iv), isolated systems and free particles models are applicable only when some aspects of dynamics are relevant (note that such ideal systems could not be even detected).

### 2.2 Postulate of Evolution

The states of physical systems are changed during the time (in fact, not only the state, but also the system itself may be changed, but such generalizations will not be studied here). Just these changes are the motion or the evolution of the system, and the task of the dynamics is to find the law of the motion. According to the experience (this means experiments) and elementary intuition, there are some basic properties of the evolution of the quantum systems, which are comprised in the following

## POSTULATE 5. - EVOLUTION

The evolution of the physical system during the time interval $t \in\left[t_{\mathrm{i}}, t_{\mathrm{f}}\right]$ is a process in which the initial state $\left|\psi, t_{\mathrm{i}}\right\rangle$ is transformed to the final state $\left|\psi, t_{\mathrm{f}}\right\rangle$ such that:
(i) the change is causal, meaning that to each initial state $\left|\psi, t_{\mathrm{i}}\right\rangle$ uniquely corresponds single final state $\left|\psi, t_{f}\right\rangle$;
(ii) the superposition of the states is conserved, i.e. if $\left|\psi, t_{\mathrm{i}}\right\rangle=\alpha_{1}\left|\psi_{1}, t_{\mathrm{i}}\right\rangle+$ $\alpha_{2}\left|\psi_{2}, t_{\mathrm{i}}\right\rangle$, then $\left|\psi, t_{\mathrm{f}}\right\rangle=\alpha_{1}\left|\psi_{1}, t_{\mathrm{f}}\right\rangle+\alpha_{2}\left|\psi_{2}, t_{\mathrm{f}}\right\rangle ;$
(iii) the number of systems within ensemble is conserved;
(iv) the change is continual with respect to time.

The properties of the evolution emphasized in the postulate are sufficient to describe it formally. Indeed, the causality means that the evolution is a map $U:\left|\psi, t_{\mathrm{i}}\right\rangle \rightarrow\left|\psi, t_{\mathrm{f}}\right\rangle$, i.e. an operator in the state space $\mathcal{H}$. Then the next property essentially means that this operator $U$ is linear. To understand the importance of the third property, recall that the probability to get the value $a$ of the observable $A$ in the state at $t_{\mathrm{i}}$ is $v\left(a, A,\left|\psi, t_{\mathrm{i}}\right\rangle\right)=$
$N_{a}^{(\mathrm{i})} / N=\left\langle\psi, t_{\mathrm{i}}\right| P_{a}(A)\left|\psi, t_{\mathrm{i}}\right\rangle$. All these probabilities are summed in 1 , giving that $\left|\psi, t_{\mathrm{i}}\right\rangle$ is normalized. Assuming that this measurement is predictive, it is possible to measure $A$ on the same ensemble but in $t_{\mathrm{f}}$. The result $v\left(a, A,\left|\psi, t_{\mathrm{f}}\right\rangle\right)=N_{a}^{(\mathrm{f})} / N$, with unchanged $N$ implies that again $\left\langle\psi, t_{\mathrm{f}} \mid \psi, t_{\mathrm{f}}\right\rangle=1$, i.e. that the evolution does not change the norm of the state. Note that this property characterize non-relativistic systems. Hence, the linear operator $U$ has to conserve the norm of the vectors. The last condition significantly restricts the operators suitable for the description of the evolution: the Wigner's theorem shows that such operators may be either unitary or antiunitary. Bearing in mind the already established linearity, it turns out that $U$ is unitary operator, or precisely the family of the operators $U\left(t_{\mathrm{i}}, t_{\mathrm{f}}\right)$. To include the last natural continuity principle, we reparameterize this family in the form $U\left(t-t_{\mathrm{i}}, t_{\mathrm{i}}\right)$, gaining the explicit function of the elapsed time $t-t_{\mathrm{i}}$ for $t \in\left[t_{\mathrm{i}}, t_{\mathrm{f}}\right]$. Of course the continuity is to be understood in the very broad sense, i.e. as the differentiability of some sufficient order.

Summarizing all the consequences of the postulate, we obtain the law of the motion:

$$
\begin{equation*}
|\psi, t\rangle=U\left(t-t_{\mathrm{i}}, t_{\mathrm{i}}\right)\left|\psi, t_{\mathrm{i}}\right\rangle . \tag{2.1}
\end{equation*}
$$

The evolution operator $U\left(t-t_{\mathrm{i}}, t_{\mathrm{i}}\right)$ completely governs the dynamics of the system, and therefore it comprises all the relevant dynamical characteristics. In the classical mechanics, the Hamiltonian is the variable defining the dynamics, and therefore it is expected to relate the evolution operator with the Hamiltonian observable obtained by the quantization postulate. To establish this relation, we use some obvious properties of the evolution. At first, for $t=t_{\mathrm{i}}$ evolution is the identity map: $U\left(0, t_{\mathrm{i}}\right)=I$. Further, if an arbitrary instant $t_{1}$ in the interval $\left(t_{\mathrm{i}}, t_{\mathrm{f}}\right)$ is singled out, the evolution may be seen as the composition of the evolution from $t_{\mathrm{i}}$ to $t_{1}$ and from $t_{1}$ to $t_{\mathrm{f}}$ : $U\left(t_{\mathrm{f}}-t_{\mathrm{i}}, t_{\mathrm{i}}\right)=$ $U\left(t_{\mathrm{f}}-t_{1}, t_{1}\right) U\left(t_{1}-t_{\mathrm{i}}, t_{\mathrm{i}}\right)$. Finally, let us fix the initial time $t_{\mathrm{i}}$ and find the infinitesimal increment $\mathrm{d} U\left(t-t_{\mathrm{i}}, t_{\mathrm{i}}\right)$ of $U$ in $t=t_{\mathrm{i}}$. By definition this is the difference $\mathrm{d} U\left(0, t_{\mathrm{i}}\right)=$ $U\left(\mathrm{~d} t, t_{\mathrm{i}}\right)-U\left(0, t_{\mathrm{i}}\right)$. On the other side, as a differential, it is proportional to $\mathrm{d} t$; this way the observable $H\left(t_{\mathrm{i}}\right)$ is defined by $\mathrm{d} U\left(0, t_{\mathrm{i}}\right)=-\frac{\mathrm{i}}{\hbar} H\left(t_{\mathrm{i}}\right) \mathrm{d} t$. Hence, $-\frac{\mathrm{i}}{\hbar} H\left(t_{\mathrm{i}}\right) \mathrm{d} t=U\left(\mathrm{~d} t, t_{\mathrm{i}}\right)-I$. The differential at some later instant $t>t_{\mathrm{i}}$ can be found with help of the composition rule of the evolutions from $t_{\mathrm{i}}$ to $t$ and from $t$ to $t+\mathrm{d} t$ : $\mathrm{d} U\left(t-t_{\mathrm{i}}, t_{\mathrm{i}}\right)=U\left(\mathrm{~d} t+t-t_{\mathrm{i}}, t_{\mathrm{i}}\right)-$ $U\left(t-t_{\mathrm{i}}, t_{\mathrm{i}}\right)=(U(\mathrm{~d} t, t)-I) U\left(t-t_{\mathrm{i}}, t_{\mathrm{i}}\right)$. Using the derived expression for the differential of $U$ at $t=t_{\mathrm{i}}$, one gets that the bracket equals $-\frac{i}{\hbar} H(t) \mathrm{d} t$. This gives the first order differential equation for $U$ and the initial condition:

$$
\begin{equation*}
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} U\left(t-t_{0}, t_{0}\right)=H(t) U\left(t-t_{0}, t_{0}\right), \quad U\left(0, t_{0}\right)=I \tag{2.2}
\end{equation*}
$$

Therefore, this is the equation for the evolution operator $U\left(t-t_{0}, t_{0}\right)$, in its differential form. Note that this two parameter family is defined through the one parameter family of the operator $H(t)$. Besides it has the dimension of energy, the hermitian operator function $H(t)$ determines the dynamics of the system; it should be expected that this is Hamiltonian observable, as it will be shown later on.

The integral form of (2.2) is obtained by formal integration. One gets $U\left(t-t_{0}, t_{0}\right)=$ $I-\frac{\mathrm{i}}{\hbar} \int_{t_{0}}^{t} H\left(t_{1}\right) U\left(t_{1}-t_{0}, t_{0}\right) \mathrm{d} t_{1}$, and the iterative substitution on the right of this formal
solution, gives the Dyson series:

$$
\begin{equation*}
U\left(t-t_{0}, t_{0}\right)=\sum_{n=0}^{\infty}\left(-\frac{\mathrm{i}}{\hbar}\right)^{n} \int_{t_{0}}^{t} \mathrm{~d} t_{1} \ldots \int_{t_{0}}^{t_{n-1}} \mathrm{~d} t_{n} H\left(t_{1}\right) \cdots H\left(t_{n}\right) . \tag{2.3}
\end{equation*}
$$

Note that the operators from the family $H(t)$ in various instants may be noncommutative, $\left[H\left(t_{i}\right), H\left(t_{j}\right)\right] \neq 0$. Therefore, in (2.3) appears the time ordered product: it is implicitly assumed that $t_{0} \leq t_{n} \leq \cdots \leq t_{1} \leq t$. Instead, one can use explicit time ordering $T$, which rearranges any product of the operators in the product ordered with the increased time:

$$
\begin{equation*}
T\left[A\left(t_{i_{1}}\right) \cdots A\left(t_{i_{n}}\right)\right] \stackrel{\text { def }}{=} A\left(t_{1}\right) \cdots A\left(t_{n}\right) \quad\left(t_{1} \geq \cdots \geq t_{n}\right) \tag{2.4}
\end{equation*}
$$

In (2.3) the time ordering is implicitly given through the upper bound of the integrals over $t_{i}$. On the other side, if all the integrations in $t_{n}, \ldots, t_{1}$ were performed from $t_{0}$ to $t$, the total integration domain would be the $n$-dimensional cube with the length $t-t_{0}$. In this cube, the domain $t_{1} \leq \cdots \leq t_{n}$ is only one $n!$-th part; precisely, to any of $n$ ! possible ordering of the instances $t_{1}, \ldots, t_{n}$, i.e. to any permutation of these instances, corresponds the sub-domain of the same volume. Thus, the integration in (2.3) may be also performed over the cube, but with explicit time ordering, and in all the sub-domains the same value is obtained, altogether $n$ ! times:

$$
\begin{equation*}
U\left(t-t_{0}, t_{0}\right)=\sum_{n=0}^{\infty} \frac{\left(-\frac{\mathrm{i}}{\hbar}\right)^{n}}{n!} \int_{t_{0}}^{t} \mathrm{~d} t_{1} \ldots \int_{t_{0}}^{t} \mathrm{~d} t_{n} T\left[H\left(t_{1}\right) \cdots H\left(t_{n}\right)\right] . \tag{2.5}
\end{equation*}
$$

In the case when the operators $H(t)$ and $H\left(t^{\prime}\right)$ commute for any instants $t$ and $t^{\prime}$, the time ordering may be omitted; each of the integrals becomes independent, and the last equation is easily integrated to

$$
\begin{equation*}
U\left(t-t_{0}, t_{0}\right)=\mathrm{e}^{-\frac{\mathrm{i}}{\hbar} \int_{t_{0}}^{t} H\left(t^{\prime}\right) \mathrm{d} t^{\prime}} \tag{2.6}
\end{equation*}
$$

For the time independent Hamiltonian $H(t)=H$, being an example of this type, we get:

$$
\begin{equation*}
U\left(t-t_{0}, t_{0}\right)=\mathrm{e}^{-\frac{\mathrm{i}}{\hbar} H\left(t-t_{0}\right)} \tag{2.7}
\end{equation*}
$$

### 2.3 Schrödinger equation

To get explicitly the law of motion of the states, one combine the equations (2.1) and (2.2) to get the famous time dependent Schrödinger equation:

$$
\begin{equation*}
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t}|\psi, t\rangle=H(t)|\psi, t\rangle, \quad|\psi, t\rangle_{t=t_{0}}=\left|\psi, t_{0}\right\rangle . \tag{2.8}
\end{equation*}
$$

This is the first order differential equation, with the initial condition given on the right. Its coordinate representation

$$
\begin{equation*}
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} \psi(\boldsymbol{r}, t)=H(\boldsymbol{r}, \boldsymbol{p}, t) \psi(\boldsymbol{r}, t), \quad \psi(\boldsymbol{r}, t)_{t=t_{0}}=\psi\left(\boldsymbol{r}, t_{0}\right) \tag{2.9}
\end{equation*}
$$

is known as the time dependent wave equation.

### 2.3.1 Conservative systems

In a conservative system the Hamiltonian is constant with respect to time, and the evolution is described by (2.7). When it acts on a hamiltonian eigen vector $\left|E, t_{0}\right\rangle$ (for the eigen value $E$ ), only the time dependent phase factor appears:

$$
|E, t\rangle=U\left(t-t_{0}\right)\left|E, t_{0}\right\rangle=\mathrm{e}^{-\mathrm{i} E\left(t-t_{0}\right) / \hbar}\left|E, t_{0}\right\rangle .
$$

This means that the eigenvectors of the Hamiltonian are stationary states in the physical sense, i.e. such a physical state is not changed during the time. On the other side, for any stationary state, i.e. state such that $|\psi, t\rangle=f(t)\left|\psi, t_{0}\right\rangle$, the Schrödinger equation reads $(\mathrm{i} \hbar \dot{f} / f)\left|\psi, t_{0}\right\rangle=H\left|\psi, t_{0}\right\rangle$; since the right side is time independent, the bracket must be a constant, meaning that $\left|\psi, t_{0}\right\rangle$ (and therefore $|\psi, t\rangle$ ) is an eigen state of the Hamiltonian.

This observation suffices to reduce the evolution problem of the conservative systems to the time independent Hamiltonian eigen problem. Indeed, let $|E, \lambda\rangle$ be an eigen basis of $H$ :

$$
\begin{equation*}
H|E, \lambda\rangle=E|E, \lambda\rangle \tag{2.10}
\end{equation*}
$$

If system in $t_{0}$ is described by the vector $|E, \lambda\rangle$, in the later instant $t$ the state vector is $\mathrm{e}^{-\mathrm{i} E\left(t-t_{0}\right) / \hbar}|E, \lambda\rangle$. Then, a quite general state can be expanded over the Hamiltonian eigen basis at any instant $t$ in the form: $|\psi, t\rangle=\sum_{E, \lambda} c_{E \lambda}(t)|E, \lambda\rangle$. Note that the time dependence is completely comprised in the expansion coefficients. On the other hand, since the evolution preserves the superposition (by the postulate),

$$
|\psi, t\rangle=U\left(t-t_{0}, t_{0}\right)\left|\psi, t_{0}\right\rangle=\sum_{E, \lambda} c_{E \lambda}\left(t_{0}\right) U\left(t-t_{0}, t_{0}\right)|E, \lambda\rangle=\sum_{E, \lambda} c_{E \lambda}\left(t_{0}\right) \mathrm{e}^{-\mathrm{i} E\left(t-t_{0}\right) / \hbar}|E, \lambda\rangle,
$$

giving that the evolution is defined by

$$
\begin{equation*}
c_{E \lambda}(t)=c_{E \lambda}\left(t_{0}\right) \mathrm{e}^{-\mathrm{i} E\left(t-t_{0}\right) / \hbar} . \tag{2.11}
\end{equation*}
$$

Hence, solving the eigen problem of the Hamiltonian, the evolution is completely determined. In other words, the solutions of (2.8) are completely determined (through (2.11)) by the solutions of (2.10), which is therefore called time independent Scrödinger equation.

These conclusions are important for the characterization of the states. In fact, if $H$ is degenerate observable, then using any $\operatorname{CSCO}\{H, A, B, \ldots\}$ containing $H$ and the observables commuting with $H$, one gets the stationary basis $|E, a, b, \ldots\rangle$, which is called complete classification of the states. Here, the eigenvalues $a, b$ etc. of the observables $A, B$ etc. take the role of the degeneracy counter $\lambda$ from (2.10). Such a states have sharp values of these observables. Consequently, these values are conserved during the evolution, i.e. they are integrals of motion, also called good quantum numbers.

### 2.3.2 Evolution of the mixed states

The evolution of the mixed states is easily derived from the evolution of the pure ones. In fact, let at $t_{0}$ a mixed state is given in its most general form $\rho\left(t_{0}\right)=\sum_{i} \omega_{i}\left|\psi_{i}, t_{0}\right\rangle\left\langle\psi_{i}, t_{0}\right|$. Afterwards mixture evolves together with its pure components $\left|\psi_{i}, t\right\rangle$, meaning:

$$
\begin{equation*}
\rho(t)=U\left(t-t_{0}, t_{0}\right) \rho\left(t_{0}\right) U^{\dagger}\left(t-t_{0}, t_{0}\right) . \tag{2.12}
\end{equation*}
$$

The differential form of this equation is obtained as derivative in $t$ of (2.12), with the substitution of (2.2) and its adjoint. This analogue of the Schrödinger equation for mixed states is called Liouville equation:

$$
\begin{equation*}
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} \rho(t)=[H(t), \rho(t)], \quad \rho(t)_{t=t_{0}}=\rho\left(t_{0}\right) \tag{2.13}
\end{equation*}
$$

Exercise 2.1: Prove Hellman-Feynman Theorem: for the hamiltonian parameterized by a real parameter $a$ the normalized eigenstates $(H(a)|E(a)\rangle=E(a)|E(a)\rangle)$ satisfy

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} a} E(a)=\langle E(a)| \frac{\mathrm{d}}{\mathrm{~d} a} H(a)|E(a)\rangle . \tag{2.14}
\end{equation*}
$$

Exercise 2.2: Prove that for any observable $A$, matrix elements of the commutator $[H, A]$ in the (normalized) stationary basis satisfy

$$
\begin{equation*}
\langle n, \lambda|[H, A]\left|n^{\prime}, \lambda^{\prime}\right\rangle=\left(E_{n}-E_{n^{\prime}}\right)\langle n, \lambda| A\left|n^{\prime}, \lambda^{\prime}\right\rangle . \tag{2.15}
\end{equation*}
$$

In particular, taking the hamiltonian in the form $H=T+U(\boldsymbol{r})$, find the matrix elements of the momentum and force (gradient of the potential) proving virial theorem.

### 2.3.3 Evolution of the probabilities and means

The evolution of the system causes the change of the observable quantities, such as probabilities of the measurement results or mean values. Even for the observables being constant in time, the measurement results may vary due to the evolution of the states.

Given the observable $A(t)$ which may be explicit function of time (e.g. if it is related to some variable external field), its mean in the state $\rho(t)$ during the time is given as $\langle A\rangle_{t}=\operatorname{Tr} \rho(t) A(t)$. Deriving this in $t$, and applying again (2.2) one gets: $\mathrm{i} \hbar \frac{\mathrm{d}}{\mathrm{d} t}\langle A\rangle_{t}=$ $\operatorname{Tr}[H(t), \rho(t)] A(t)+\mathrm{i} \hbar \operatorname{Tr} \rho(t) \frac{\mathrm{d}}{\mathrm{d} t} A(t)$. Expanding commutator, this becomes:

$$
\begin{equation*}
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t}\langle A\rangle_{t}=\langle[A(t), H(t)]\rangle_{t}+\mathrm{i} \hbar\left\langle\frac{\partial}{\partial t} A\right\rangle_{t} \tag{2.16}
\end{equation*}
$$

The derivative of a classical variable in $t$ is based on the notion of its value in the close instants. In the quantum sense, the observable itself may have no particular value, and obtains it in the measurement process. Therefore the derivatives must refer not only to the parametric dependence of $A$ of time, but to the both quantities, observable and state, participating in the measurement process. In this sense, since (2.16) is fulfilled for each state, it can be used to define the time derivative of the observable, and get the classical like equation of motion:

$$
\begin{equation*}
\left\langle\frac{\mathrm{d}}{\mathrm{~d} t} A\right\rangle_{t} \stackrel{\text { def }}{=} \frac{\mathrm{d}}{\mathrm{~d} t}\langle A\rangle_{t}, \quad \mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} A=[A(t), H(t)]+\mathrm{i} \hbar \frac{\partial}{\partial t} A . \tag{2.17}
\end{equation*}
$$

Also, it is straightforward to find how the probability of a particular measurement result depends on time:

$$
v_{t}(a, A,|\psi\rangle)=\| P_{a}(A) U\left(t-t_{0}, t_{0}\right)|\psi\rangle \|^{2} .
$$

Analogously, given a state $|\phi\rangle$ fixed as a common eigenvector of a CSCO, the transition probability from arbitrary state $|\psi\rangle$ to $|\phi\rangle$ is

$$
\left.v_{t}(|\psi\rangle \rightarrow|\phi\rangle)=\left|\langle\phi| U\left(t-t_{0}, t_{0}\right)\right| \psi\right\rangle\left.\right|^{2} .
$$

### 2.3.4 Probability current and continuity equation

Assuming that the one particle Hamiltonian has the form $H=\boldsymbol{p}^{2} / 2 m+V(\boldsymbol{r}, t)$, the Schrödinger equation in the coordinate representation and its complex conjugate are:

$$
\begin{aligned}
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} \psi(\boldsymbol{r}, t) & =-\frac{\hbar^{2}}{2 m} \Delta \psi(\boldsymbol{r}, t)+V(\boldsymbol{r}, t) \psi(\boldsymbol{r}, t) \\
-\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} \psi^{*}(\boldsymbol{r}, t) & =-\frac{\hbar^{2}}{2 m} \Delta \psi^{*}(\boldsymbol{r}, t)+V(\boldsymbol{r}, t) \psi^{*}(\boldsymbol{r}, t) .
\end{aligned}
$$

Multiplying the first of them by $\psi^{*}(\boldsymbol{r}, t)$ and the second one by $\psi(\boldsymbol{r}, t)$, and summing them, one gets the continuity equation:

$$
\begin{equation*}
\frac{\partial \rho}{\partial t}+\operatorname{div} \boldsymbol{j}=0 \tag{2.18}
\end{equation*}
$$

Here, $\rho(\boldsymbol{r})=\psi^{*}(\boldsymbol{r}, t) \psi(\boldsymbol{r}, t)=\tilde{v}(\boldsymbol{r}, \hat{\boldsymbol{r}},|\psi\rangle)$ is the density of the probability (1.5) of the particle occurrence in $r$, while

$$
\begin{equation*}
\boldsymbol{j}=\frac{\hbar}{2 m \mathrm{i}}\left(\psi^{*}(\boldsymbol{r}, t) \nabla \psi(\boldsymbol{r}, t)-\psi(\boldsymbol{r}, t) \nabla \psi^{*}(\boldsymbol{r}, t)\right)=\frac{1}{m} \operatorname{Re}\left(\psi^{*}(\boldsymbol{r}, t) \frac{\hbar}{\mathrm{i}} \nabla \psi(\boldsymbol{r}, t)\right) . \tag{2.19}
\end{equation*}
$$

is the probability current.
Exercise 2.3: ${ }^{\circ}$ Show that the probability current of the free particle in the state $\psi_{\mathrm{pw}}(\boldsymbol{r})=C \mathrm{e}^{\mathrm{i} k \cdot \boldsymbol{r}}$ is

$$
\begin{equation*}
j_{\mathrm{pw}}=\frac{\hbar}{m} \boldsymbol{k}|C|^{2} . \tag{2.20}
\end{equation*}
$$

### 2.3.5 Uncertainty relation time-energy

Time and energy are related to some extent analogously to coordinate and momenta. Therefore, an uncertainty relation involving these quantities may be expected. Still, since there is no observable of time, such relation must be derived indirectly.

We say that $A$ is appropriate observable for the state $|\psi\rangle$ if it is not explicitly time dependent and $\frac{\mathrm{d}}{\mathrm{d} t}\langle A\rangle_{|\psi\rangle} \neq 0$. For such an observable and state, the minimal time $\tau_{A}=\Delta_{|\psi\rangle}(A) /\left|\frac{\mathrm{d}}{\mathrm{d} t}\langle A\rangle_{|\psi\rangle}\right|$ must pass until the change of mean overwhelms the dispersion $\Delta_{|\psi\rangle}(A)$.

Theorem 2.1 (Time-energy uncertainty relation) Any non-stationary state $|\psi, t\rangle$ and any appropriate observable $A$ satisfy:

$$
\begin{equation*}
\tau_{A} \Delta_{|\psi\rangle}(H) \geq \frac{1}{2} \hbar . \tag{2.21}
\end{equation*}
$$

$=$ Proof: According to (1.25), $\left.\Delta_{\psi}(A) \Delta_{\psi}(H) \geq \frac{1}{2}\left|\langle[A, H]\rangle_{\psi}\right|=\frac{1}{2}\left|i \hbar \frac{\mathrm{~d}}{\mathrm{~d} t}\langle A\rangle_{|\psi\rangle}\right|=\frac{\hbar}{2} \right\rvert\, \frac{\mathrm{d}}{\mathrm{d} t}\langle A\rangle_{|\psi\rangle}$, and the relation looked for follows when $\Delta(A)$ is expressed in terms of $\tau_{A}$.

Since this inequality holds for all the appropriate observables, also the infimum $\tau=$ $\inf \left\{\tau_{A} \mid A\right.$ appropriate $\}$ satisfies $\tau \Delta_{|\psi\rangle}(H) \geq \frac{1}{2} \hbar$. Hence, $\tau$ is the best estimate independent of the particular observable.

In the scattering experiments, the target energy is measured by means of beams of particles. The latter are free particles, with hamiltonian $H=p^{2} / 2 m$, well collimated with the dispersion $\Delta p \neq 0$. The beam is in average localized at $\langle x\rangle$ with the dispersion $\Delta x$, and the mean value evolves. Hence $x$ is appropriate observable, and the measurement has duration of at least $\tau_{x}$.

Completely different approach to the time-energy uncertainty relation appears on the context of unstable states. In fact, when the quantum system, like atom, is excited from the ground state, the interaction with the surrounding results usually in the deexcitation. Therefore the excited states have their lifetime, and $\tau_{1 / 2}$ is the half lifetime of such a state. On the other hand, the excited state energy cannot be exactly found in the experiments, but with some error $\Delta E$. The phenomenological uncertainty relation $\tau_{1 / 2} \Delta E \geq \hbar$ is experimentally found:

### 2.4 Simple systems

### 2.4.1 One-Dimensional systems

### 2.4.2 Harmonic oscillator

The hamiltonian of the one-dimensional harmonic oscillator is

$$
\begin{equation*}
H=\frac{1}{2} \frac{p^{2}}{m}+\frac{1}{2} m \omega^{2} q^{2} \tag{2.22}
\end{equation*}
$$

Introducing new observables $Q=q \sqrt{m \omega / \hbar}$ and $P=p / \sqrt{m \omega \hbar}=-\mathrm{i} \partial_{Q}$ (note that $P$ and $Q$ are not canonically conjugated, since $[Q, P]=\mathrm{i}$, one gets

$$
\begin{equation*}
H=\hbar \omega H^{\prime}, \quad H^{\prime}=\frac{1}{2}\left(P^{2}+Q^{2}\right) \tag{2.23}
\end{equation*}
$$

Exercise 2.4: (Linear Harmonic Oscillator) Find eigenfunctions and energy spectrum of the Hamiltonian $H=\frac{p_{x}^{2}}{2 m}+\frac{m \omega^{2}}{2} x^{2}$.

Besides direct solution (Exercise 2.4) of the appearing hypergeometric equation, the eigen problem of the hamiltonian may be solved with help of the operators:

$$
\begin{equation*}
a^{ \pm}=\frac{1}{\sqrt{2}}(Q \mp \mathrm{i} P)=\frac{1}{\sqrt{2}}\left(Q \mp \frac{\mathrm{~d}}{\mathrm{~d} Q}\right) . \tag{2.24}
\end{equation*}
$$

The following significant properties of these operators directly follow from the definition:

$$
\begin{equation*}
\left(a^{-}\right)^{\dagger}=a^{+}, \quad\left[a^{-}, a^{+}\right]=1 \tag{2.25}
\end{equation*}
$$

To this pair of operators the third one, occupancy number operator is accompanied:

$$
\begin{equation*}
n=a^{+} a^{-} . \tag{2.26}
\end{equation*}
$$

It is manifestly positive, and, since $n|x\rangle=0$ iff $0=\langle x| n|x\rangle=\| a^{-}|x\rangle \|$, its kernel is $\operatorname{ker}(n)=\operatorname{ker}\left(a^{-}\right)$. The following commutator relation is easily verified:

$$
\begin{equation*}
\left[n, a^{ \pm}\right]= \pm a^{ \pm} \tag{2.27}
\end{equation*}
$$

Exercise 2.5: Show that $\left[a^{\mp},\left(a^{ \pm}\right)^{s}\right]= \pm s\left(a^{ \pm}\right)^{s-1}$, and (??? PROVERITI FORMULU)

$$
\left[\left(a^{-}\right)^{m},\left(a^{+}\right)^{p}\right]=\sum_{s=1}^{\min \{m, p\}} \frac{p!}{(p-s)!} \frac{m!}{(m-s)!}\left(a^{+}\right)^{p-s}\left(a^{-}\right)^{m-s} .
$$

Exercise 2.6: Show that any monomial over $a^{+}$and $a^{-}$may be expressed as polynomial with ordered terms $\left(\left(a^{+}\right)^{x}\right.$ before $\left.\left(a^{-}\right)^{y}\right)$.

The introduced triple of the operators $a^{-}, a^{+}$and $n$ is very important in the quantum theory of many particles. At first we note that $a^{ \pm}$are by (2.24) biuniquely related to the fundamental set of observables $\{x, p\}$, making possible to express all the opservables in terms of these two. The Exercise 2.5 then asserts that any observable may be expanded in the series over $a^{-}$and $a^{+}$with the terms being monomials of the type $\left(a^{+}\right)^{p}\left(a^{-}\right)^{m}$, as shown in the Exercise 2.6. Particularly important are spectral properties of $n$.

Theorem 2.2 The spectrum of occupancy number operator is purely discrete $\sigma(n)=$ $\{0,1,2, \ldots\}$ and nondegenerate, with all the eigenvectors $|i\rangle$ determined by the vacuum state $|0\rangle$ as

$$
\begin{equation*}
|i\rangle=\frac{1}{\sqrt{i!}}\left(a^{+}\right)^{i}|0\rangle \tag{2.28}
\end{equation*}
$$

-Proof: As a positive operator, $n$ must have an eigenvalue $x>0$. Let $|x\rangle$ be the corresponding eigenvector: $n|x\rangle=x|x\rangle$. The commutator relation (2.27) gives $n\left(a^{ \pm}|x\rangle\right)=\left(a^{ \pm} n \pm a^{ \pm}\right)|x\rangle=(x \pm$ 1) ( $a^{ \pm}|x\rangle$ ). Obviously, together with $x$ the eigenvalues are also $x-1$ (for $\left.a^{-}|x\rangle\right), x-2$ (for $\left.\left(a^{-}\right)^{2}|x\rangle\right)$, etc. This series would become negative at some point. The only way to avoid such contradiction is that $x$ is integer itself: then one eigenvector is $\left(a^{-}\right)^{x}|x\rangle \sim|0\rangle$. The vacuum vector $|0\rangle$ satisfies $n|0\rangle=0$, meaning that it is from the kernel of $n$, implying $0=\langle 0| n|0\rangle=\langle 0| a^{+} a^{-}|0\rangle=\| a^{-}|0\rangle \|^{2}$. Thus, $a^{-}|0\rangle=0$, and for $j>x$ negative part of the spectrum is not generated as $\left(a^{-}\right)^{j}|j\rangle=0$. The series is then cut here, since all the remaining members vanish, restricting the spectrum of $n$ to nonnegative integers. Assume that this spectrum is not simple; the degeneracy can be removed by a compatible observable. As $q$ and $p$ are functions of $a^{+}$and $a$, any observable has a form $A=\sum_{p m} a_{p m}\left(a^{+}\right)^{p}\left(a^{-}\right)^{m}$. The compatibility with $n$ imposes condition

$$
0=[A, n]=\sum_{p m} a_{p m}\left[\left(a^{+}\right)^{p}\left(a^{-}\right)^{m}, a^{+} a^{-}\right]=\sum_{p m}(m-p) a_{p m}\left(a^{+}\right)^{p}\left(a^{-}\right)^{m},
$$

requiring $a_{p m}=0$ unless $m=p$. But then $A=\sum_{s} b_{s} n^{s}$, i.e. it is function of $n$ and cannot remove the degeneracy. Hence, $n$ is complete observable. Finally, it is clear that $|i\rangle=c_{i}\left(a^{+}\right)^{i}|0\rangle$, and only the normalizing coefficient in (2.28) is to be found. Assuming that the basis $\{|i\rangle\}$ is orthonormal, the squared
norm of $a^{+}|i\rangle=c|i+1\rangle$ is $\langle i| a^{-} a^{+}|i\rangle=i+1$, i.e. $a^{+}|i\rangle=\sqrt{i+1}|i+1\rangle$, verifying the proposed result. !

Now it is easy to solve the eigen problem of the harmonic oscillator. Namely, since obviously $H^{\prime}=n+\frac{1}{2}$, i.e.

$$
\begin{equation*}
H=\hbar \omega\left(n+\frac{1}{2}\right) \tag{2.29}
\end{equation*}
$$

the eigen vectors of $H$ are exactly the same vectors (2.28) as for $n$, while the spectrum of $H$ is shifted upward for one half, and then multiplied by $\hbar \omega$. Thus the spectrum consist of equidistant eigenvalues, the difference between successive ones being $\hbar \omega$. The remarkable features of this spectrum allow us to consider that the eigen state $|i\rangle$ describes a system with $i$ free (noninteracting) quasi-particles: each of them contributes by the single quantum $\hbar \omega$. Operator $a^{+}$, shifting from $|i\rangle$ to $|i+1\rangle$, acts as the creation operator of these quasi particles, while its adjoint $a^{-}$is the annihilation operator. In this context, the eigen value $i$ of $n$ is the number of the quanta in the state $|i\rangle$, and $n$ is called occupation number operator. Usually the annihilation operator $a^{-}$is denoted by $a$, while the creation one $a^{+}$is according to the first property simply its adjoint $a^{\dagger}$.

To get the eigenfunctions in the $Q$-representation $\psi_{i}(Q)=\langle Q \mid i\rangle$, we note that the vacuum state is singled out by $a^{-}|0\rangle=0$, which is by (2.24) the differential equation $\frac{\mathrm{d} \psi}{\mathrm{d} Q}+Q \psi=0$, with the solution $\psi_{0}(Q)=\pi^{-\frac{1}{4}} \mathrm{e}^{-\frac{1}{2} Q^{2}}$. Then applying (2.28) one recognizes Hermite function $\psi_{i}(Q)$ of the order $i$ :

$$
\begin{equation*}
\psi_{i}(Q)=\frac{1}{\sqrt{\sqrt{\pi} 2^{i} i!}} \mathrm{e}^{-\frac{1}{2} Q^{2}} H_{i}(Q) \tag{2.30}
\end{equation*}
$$

Exercise 2.7: Prove (2.30). Hint: show the operator identity: $\left(Q-\frac{\mathrm{d}}{\mathrm{d} Q}\right)=\left(-\mathrm{e}^{\frac{1}{2} Q^{2}} \frac{\mathrm{~d}}{\mathrm{~d} Q} \mathrm{e}^{-\frac{1}{2} Q^{2}}\right)$.
Exercise 2.8: Generalize the above consideration to the system with $d$ harmonic degrees of freedom; its hamiltonian is $H=\sum_{i=1}^{d}\left(\frac{1}{2} \frac{p_{i}^{2}}{m_{i}}+\frac{1}{2} m_{i} \omega_{i}^{2} q_{i}^{2}\right)$.

Exercise 2.9: Find the mixed state describing harmonic oscillator in the thermodynamic equilibrium with a heat reservoir at temperature $T$. Discuss the mean energy in the limits of zero and very high temperature.

### 2.5 Pictures

The unitary transformations preserve the complete structure of the quantum mechanics. Mathematically, this refers to the invariance of linear combination and norm of vectors, as well as of the operator product and commutator. These leads to the physically obvious independence of the measurement results on the choice of the basis in the state space. Nevertheless, a particular choice of the basis may be more convenient either for conceptual or practical reasons. Here we will consider three such possibilities, called pictures, related to the dynamics of the system, which proved to be important for various purposes.

In fact, the first one has been introduced in the derivation of the dynamical law. Throughout the section 2.3 it has been tacitly assumed that the states are evolving, while
the observables remain unchanged at least for the conservative systems. This picture is used to define the other ones. Indeed, any dynamical picture will be defined by an unitary operator function $T(t)$ over time; in such a picture state and observable are in the instant $t$ given by their Scrödinger values $\left(|\psi, t ; \mathrm{S}\rangle\right.$ and $\left.A_{\mathrm{S}}(t)\right)$ and the picture operator $T(t)$ :

$$
\begin{equation*}
|\psi, t ; T\rangle=T(t)|\psi, t ; \mathrm{S}\rangle, \quad A_{T}(t)=T(t) A_{\mathrm{S}}(t) T^{-1}(t) . \tag{2.31}
\end{equation*}
$$

Hence, the Schrödinger picture is obtained with $T(t)=I$, with the evolution law for states given by (2.1) or (2.8), while the observables are considered time independent as far as the system evolution is regarded, although an explicit time dependence $A(t)$ may be imposed by the surrounding (e.g. the apparatus is moving).

### 2.5.1 Heisenberg's picture

Within this evolution picture the picture operator is inverse to the evolution: $T(t)=$ $U^{-1}(t)$. The equations of motion

$$
\begin{align*}
|\psi, t ; \mathrm{H}\rangle & =|\psi, 0 ; \mathrm{S}\rangle=|\psi, 0 ; \mathrm{H}\rangle  \tag{2.32a}\\
A_{\mathrm{H}}(t) & =U^{-1}(t) A_{\mathrm{S}}(t) U(t), \quad \mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} A_{\mathrm{H}}(t)=\left[A_{\mathrm{H}}(t), H(t)\right]+\mathrm{i} \hbar \frac{\partial}{\partial t} A_{\mathrm{H}}(t) \tag{2.32b}
\end{align*}
$$

are easily derived from the Scrödinger's one. Note that the states are time independent, while the observables evolve. The observables are changed for two different reasons. The evolution of the system, incorporated through the picture itself (by the definition on the right of (2.32b)) results in the first (commutator) term in the differential form at the right of (2.32b). Possible explicit (external influence) changes of the measuring apparatus gives the last contribution in $(2.32 \mathrm{~b})$, where one should be warned that the partial time derivative is defined as $\frac{\partial}{\partial t} A_{\mathrm{H}}(t) \stackrel{\text { def }}{=} U^{-1}(t)\left[\frac{\partial}{\partial t} A_{\mathrm{S}}(t)\right] U(t)$.

### 2.5.2 Dirac's picture

This picture, also called the interaction picture, gives suitable point of view for the physical problems with Hamiltonian containing two parts, $H=H_{0}+H^{\prime}$; in fact, usually it is assumed that $H^{\prime}$ is small with respect to $H_{0}$, but for purpose of introducing Dirac picture this is not important. Then the picture is characterized by the operator $T(t)=U_{0}^{-1}(t)$, where $U_{0}$ is defined as the evolution realized by the first part of the Hamiltonian, i.e. satisfying (2.2) for $H_{0}$; defining the remaining evolution by $U=U_{0} U^{\prime}$, from (2.2) one gets:

$$
\begin{array}{ll}
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} U_{0}\left(t-t_{0}, t_{0}\right)=H_{0}(t) U_{0}\left(t-t_{0}, t_{0}\right), & U_{0}\left(0, t_{0}\right)=I, \\
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} U^{\prime}\left(t-t_{0}, t_{0}\right)=H_{D}^{\prime}(t) U^{\prime}\left(t-t_{0}, t_{0}\right), & U^{\prime}\left(0, t_{0}\right)=I . \tag{2.34}
\end{array}
$$

It must be understood that although these equations apparently decouple the dynamics related to $H_{0}$ and $H^{\prime}$, the second one, referring to $H^{\prime}$ implicitly involves $H_{0}$ through the

Dirac picture of $H^{\prime}$. Both the evolutions of the states and observables are nontrivial; from (2.31) follows

$$
\begin{equation*}
|\psi, t ; \mathrm{D}\rangle=U^{\prime}(t)|\psi, 0 ; \mathrm{D}\rangle, \quad A_{\mathrm{D}}(t)=U_{0}^{-1}(t) A_{\mathrm{S}}(t) U_{0}(t) \tag{2.35}
\end{equation*}
$$

The differential form of these equations is easily found:

$$
\begin{array}{r}
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t}|\psi, t ; \mathrm{D}\rangle=H_{\mathrm{D}}^{\prime}(t)|\psi, t ; \mathrm{D}\rangle \\
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} A_{\mathrm{D}}(t)=\left[A_{\mathrm{D}}(t), H_{0 \mathrm{D}}\right]+\mathrm{i} \hbar \frac{\partial}{\partial t} A_{\mathrm{D}}(t) \tag{2.37}
\end{array}
$$

with $\frac{\partial}{\partial t} A_{\mathrm{D}}(t) \stackrel{\text { def }}{=} U_{0}^{-1}(t)\left[\frac{\partial}{\partial t} A_{\mathrm{S}}(t)\right] U_{0}(t)$.

### 2.6 Propagator and Green function

Retarded propagator is two instance $\left(t_{2}, t_{1}\right)$ function derived from evolution operator. It vanishes for $t_{2}<t_{1}$, while otherwise it is matrix element of the evolution operator in the coordinate representation:

$$
\begin{align*}
K\left(\boldsymbol{r}_{2}, t_{2} ; \boldsymbol{r}_{1}, t_{1}\right) & =\left\langle\boldsymbol{r}_{2}, t_{2}\right| U\left(t_{2}-t_{1}, t_{1}\right)\left|\boldsymbol{r}_{1} 2, t_{1}\right\rangle \theta\left(t_{2}-t_{1}\right)  \tag{2.38}\\
& =\theta\left(t_{2}-t_{1}\right) \sum_{n, \lambda} \phi_{n \lambda}^{*}\left(\boldsymbol{r}_{1}\right) \mathrm{e}^{-\mathrm{i} E_{n}\left(t_{2}-t_{1}\right)} \phi_{n \lambda}\left(\boldsymbol{r}_{2}\right)
\end{align*}
$$

Using eigenbasis of the Hamiltonian, $\phi_{n}\left(\boldsymbol{r}_{2}\right)=\left\langle\boldsymbol{r}_{2} \mid E_{n}\right\rangle$, satisfying Schrödinger equation $\left(\mathrm{i} \hbar \frac{\partial}{\partial t_{2}}-H\left(\boldsymbol{r}_{2}, \boldsymbol{p}_{2}\right)\right) \mathrm{e}^{-\mathrm{i} E_{n}\left(t_{2}-t_{1}\right)} \phi_{n}\left(\boldsymbol{r}_{2}\right)=0$, one gets (step function of time is a constant for Hamiltonian operator, its derivative over $t_{2}$ is $\delta\left(t_{2}-t_{1}\right)$, and $\sum_{n, \lambda} \phi_{n \lambda}^{*}\left(\boldsymbol{r}_{1}\right) \phi_{n \lambda}\left(\boldsymbol{r}_{2}\right)=$ $\left.\delta\left(\boldsymbol{r}_{2}-\boldsymbol{r}_{1}\right)\right)$ :

$$
\begin{equation*}
\left(\mathrm{i} \hbar \frac{\partial}{\partial t_{2}}-H\left(\boldsymbol{r}_{2}, \boldsymbol{p}_{2}\right)\right) K\left(\boldsymbol{r}_{2}, t_{2} ; \boldsymbol{r}_{1}, t_{1}\right)=\mathrm{i} \hbar \delta\left(t_{2}-t_{1}\right) \delta\left(\boldsymbol{r}_{2}-\boldsymbol{r}_{1}\right) . \tag{2.39}
\end{equation*}
$$

So, it turns out that propagator is Green function of the Schrödinger equation.
The importance of the (retarded) propagator stems from the fact that knowing wave function at instance $t_{1}$, in the later instance $t_{2}$ the evolved state $\left|\psi, t_{2}\right\rangle=U\left(t_{2}-t_{1}, t_{1}\right)$ is in the coordinate representation obtained as

$$
\begin{equation*}
\psi\left(\boldsymbol{r}_{2}, t_{2}\right)=\int \mathrm{d} r_{1}^{3}\left\langle\boldsymbol{r}_{2}\right| U\left(t_{2}-t_{1}, t_{1}\right)\left|\boldsymbol{r}_{1}\right\rangle \psi\left(\boldsymbol{r}_{1}, t_{1}\right)=\int \mathrm{d} r_{1}^{3} K\left(\boldsymbol{r}_{2}, t_{2} ; \boldsymbol{r}_{1}, t_{1}\right) \psi\left(\boldsymbol{r}_{1}, t_{1}\right) \tag{2.40}
\end{equation*}
$$

This resembles classical Huygens' principle, where each point of the wavefront becomes a secondary wave source. Note further that the first part of the above is valid even when $t_{2}<t_{1}$; in that case, one can introduce advanced propagator $\left\langle\boldsymbol{r}_{2}\right| U\left(t_{2}-t_{1}, t_{1}\right)\left|\boldsymbol{r}_{1}\right\rangle \theta\left(t_{1}-t_{2}\right)$.

The propagator techniques are well developed, and essentially invoke another Feynman's path integral interpretation of quantum mechanics, introduced in the next section,

### 2.7 Feynman's path integral approach

Feynman showed that the quantum dynamics may be understood in a conceptually simple way $^{1}$, giving direct insight to the difference between classical and quantum treatment; in addition it offers a shortcut to many important techniques of field theory (both in solid state and particle physics).

The Feynman's approach is based on the expression for the probability amplitude to detect a particle, initially being (in $t_{0}=0$ ) at $x_{0}$, at $x=x_{N}$ in the later instance $t=t_{N}$ : $\left\langle q_{N}, t_{N} \mid q_{0}, 0\right\rangle$. The time interval $[0, t]$ may be divided in $N$ equal parts $\Delta t=t_{N} / N$, and the amplitude becomes:

$$
\left\langle q_{N} \mid q_{0}, 0\right\rangle=\left\langle q_{N}\right| \underbrace{U(\Delta t) \cdots U(\Delta t)}_{N}\left|q_{0}\right\rangle .
$$

Namely, assuming that the system is conservative the evolution operator $U(t)=\mathrm{e}^{-\frac{\mathrm{i}}{\hbar} H t}$ is written as $U(\Delta t)^{N}$. Further for each $t_{i}=i \Delta t$ we introduce identity in the form $\int \mathrm{d} q_{i}\left|q_{i}\right\rangle\left\langle q_{i}\right|$, giving

$$
\begin{align*}
\left\langle q_{N} \mid q_{0}\right\rangle & =\int \mathrm{d} q_{1} \cdots \mathrm{~d} q_{N}\left\langle q_{N}\right| U(\Delta t)\left|q_{N-1}\right\rangle \cdots\left\langle q_{i}+1\right| U(\Delta t)\left|q_{i}\right\rangle \cdots\left\langle q_{1}\right| U(\Delta t)\left|q_{0}, 0\right\rangle \\
& =\int \mathrm{d} q_{1} \cdots \mathrm{~d} q_{N} K(N, N-1) \cdots K(1,0) \tag{2.41}
\end{align*}
$$

(he $K(i, i-1)$ is short for the propagator $\left.K\left(q_{i}, t_{i} ; q_{i-1}, t_{i-1}\right)\right)$.
Further we assume that the potential is coordinate dependent only, i.e. $H=\frac{\hat{p}^{2}}{2 m}+$ $V(\hat{q})$, giving $\left\langle q_{i}+1\right| U(\Delta t)\left|q_{i}\right\rangle=\left\langle q_{i}+1\right| \mathrm{e}^{\left.-\frac{1}{\hbar} \frac{\hat{p}^{2}}{2 m}+V(\hat{q})\right) \Delta t}\left|q_{i}\right\rangle$. Note that $[\hat{p} \Delta t, V \Delta t] \sim$ $V^{\prime}(\hat{q}) \Delta^{2} t$, and in the limit $\Delta t=0$ (i.e. for infinite $N$ ) this comutator vanishes, meaning that $\mathrm{e}^{-\frac{i}{\hbar}\left(\frac{\hat{p}^{2}}{2 m}+V(\hat{q})\right) \Delta t} \rightarrow \mathrm{e}^{-\frac{i}{\hbar} \frac{\hat{p}^{2}}{2 m}} \mathrm{e}^{\frac{\mathrm{i}}{\hbar} V(\hat{q}) \Delta t}$. Thus, in the same limit the considered amplitude becomes: $\left\langle q_{i}+1\right| U(\Delta t)\left|q_{i}\right\rangle=\left\langle q_{i}+1\right| \mathrm{e}^{-\frac{i}{\hbar} \frac{\hat{p}^{2}}{2 m}}\left|q_{i}\right\rangle \mathrm{e}^{\frac{\mathrm{i}}{\hbar} V(\hat{q}) \Delta t}$. Introducing the identity in the form of the momentum eigen decomposition, $I=\int \mathrm{d} p|p\rangle\langle p|$, with plane waves $|p\rangle=$ $\frac{1}{\sqrt{2 \pi}} \int \mathrm{~d} q \mathrm{e}^{\frac{\mathrm{i}}{\hbar} p q}|q\rangle$, one gets $\left\langle q_{i}+1\right| U(\Delta t)\left|q_{i}\right\rangle=\frac{1}{2 \pi} \int \mathrm{~d} \mathrm{e}^{\frac{\mathrm{i}}{\overline{( }\left(p \frac{q_{i+1}-q_{i}}{\Delta t}-\frac{p^{2}}{2 m}\right) \Delta t}} \mathrm{e}^{\frac{\mathrm{i}}{\hbar} V\left(q_{i}\right) \Delta t}$. The last integral may be calculated ${ }^{2}$, giving $\left\langle q_{i}+1\right| U(\Delta t)\left|q_{i}\right\rangle=\sqrt{\frac{m \hbar e^{-i \pi / 2}}{2 \pi \Delta t}} \mathrm{e}^{\frac{1}{\hbar}\left[\frac{m}{2}\left(\frac{q_{i+1}-q_{i}}{\Delta t}\right)^{2}-V\left(q_{i}\right)\right] \Delta t}$. It is easy to recognize action $\int_{0}^{\Delta t} \mathrm{~d} t L(q, \dot{q})$ in the exponent, where the classical lagrangian is $L(q, \dot{q})=\frac{m}{2} \dot{q}^{2}-V(q)$ (recall that $\Delta t$ is infinitesimal). Altogether, making the product, the looked for amplitude becomes:

$$
\begin{equation*}
\left\langle q_{N} \mid q_{0}, 0\right\rangle=\int \mathcal{D} q \mathrm{e}^{\frac{i}{\hbar} \int_{0}^{t} L(q, \dot{q}) \mathrm{d} t}, \quad \mathcal{D} q=\lim _{N \rightarrow \infty}\left(\frac{m N \hbar}{2 \pi \mathrm{i} t}\right)^{N / 2} \prod_{i=1}^{N-1} \mathrm{~d} q_{i} \tag{2.42}
\end{equation*}
$$

[^2]The integral is the sum of the contributions of all possible trajectories. The infinitesimal volume $\mathcal{D} q$ is up to the phase equal to the mathematically well founded Wienner measure in the space of trajectories. The discrepancy may be formally cured by so called Wick's rotation, making time imaginary (i.e. if it is used instead of $t$ ). Anyway, as the probabilities are observable quantities, the phase is not relevant. Therefore, (2.42) is the path or functional integral formulation of quantum dynamics. It reveals that the probability amplitude of transition from $q_{0}$ to $q_{N}$ is sum of contributions over all possible paths connecting these two positions. Each path $q_{0} \rightarrow q_{N}$ is weighted by its amplitude equal to $\mathrm{e}^{\frac{i}{\hbar} S\left(q_{0} \rightarrow q_{N}\right)}$.

In the classical mechanics only the extremal path is realized. On the other hand, in the quantum mechanics all the paths are realized, but with different weights corresponding to the action along paths. As the action is in the exponent, the classical path corresponds to the stationary, i.e. minimal phase. The weights of the other paths are dumped, as the action is increased, with more emphasized oscillatory character of the integrand. There is also quasi classical limit (of quantum mechanics), where the paths with the action $\int_{0}^{t} L(q, \dot{q}) \mathrm{d} t$ much greater than $\hbar$ are taken into account only (this corresponds to retaining of the leading terms in the expansion of the exponent in (2.42)).

Expansion of the action functional around the classical path (being the minimum of the functional) must be quadratic in the parameters introduced in the space of trajectories. Therefore, the leading terms in such a perturbative approach (the method of stationary phase) are Gaussian integrals. This makes the formalism not only conceptually important, but also efficient in various problems, despite the fact that only Gaussian (besides some trivial ones) path integrals are exactly calculated.

## Chapter 3

## Galilean Transformations

The space-time symmetries are important in many aspects of system dynamics. Their most significant manifestations are the conservation laws in the isolated systems. In fact, these conservation laws are the evidence of the symmetries. When the conserved quantities are generators of the Lie-algebra of the symmetry group, they explicitly refer only to the connected component of the group. The other components give rise to conserved parities.

Non-relativistic and relativistic mechanics assume different geometry of the spacetime. This is manifested through different space-time symmetries, i.e. transformations leaving space-time unchanged, usually called relativity principle: Galilean one, relevant for non-relativistic mechanics, assumes space-time invariance with respect to the Galilean group, in contrast to the Poincare group related to the Einstein relativity.

### 3.1 Galilean Group

The Galilean group $\mathcal{G}$ consists of the space-time transformations $g(\boldsymbol{\phi}, \boldsymbol{a}, \boldsymbol{v}, \tau)$ acting on the space time point $(\boldsymbol{r}, t)$ as follows:

$$
\begin{equation*}
g(\boldsymbol{\phi}, \boldsymbol{a}, \boldsymbol{v}, \tau)(\boldsymbol{r}, t)=\left(R_{\phi} \boldsymbol{r}+\boldsymbol{a}-\boldsymbol{v} t, t+\tau\right) . \tag{3.1}
\end{equation*}
$$

Here, $\phi$ is the 3-dimensional real vector defining the rotation $R_{\phi}$ around $\phi$ for the angle $\phi$ equal to the length of $\boldsymbol{\phi}$. Also $\boldsymbol{a}$ and $\boldsymbol{v}$ are 3-dimensional real vectors, specifying the translation and boost, while $\tau$ is a real scalar, defining the time translation. The action on the momentum $\boldsymbol{p}=m \frac{\mathrm{~d} r}{\mathrm{~d} t}$ is derived directly:

$$
\begin{equation*}
g(\phi, \boldsymbol{a}, \boldsymbol{v}, \tau) \boldsymbol{p}=R_{\phi} \boldsymbol{p}-m \boldsymbol{v} \tag{3.2}
\end{equation*}
$$

All these transformations will be specified more precisely later on, and now we examine the structural properties of $\mathcal{G}$. The consecutive application of two group elements gives the multiplication law:

$$
\begin{equation*}
g(\boldsymbol{\phi}, \boldsymbol{a}, \boldsymbol{v}, \tau) g\left(\boldsymbol{\phi}^{\prime}, \boldsymbol{a}^{\prime}, \boldsymbol{v}^{\prime}, \tau^{\prime}\right)=g\left(\boldsymbol{\phi}^{\prime \prime}, \boldsymbol{a}+R_{\phi} \boldsymbol{a}^{\prime}-\boldsymbol{v} \tau^{\prime}, R_{\phi} \boldsymbol{v}^{\prime}+\boldsymbol{v}, \tau+\tau^{\prime}\right), \tag{3.3}
\end{equation*}
$$

where $\boldsymbol{\phi}^{\prime \prime}$ is defined such that $R_{\phi^{\prime \prime}}=R_{\phi} R_{\phi^{\prime}}$. Also, the requirement $g(\boldsymbol{\phi}, \boldsymbol{a}, \boldsymbol{v}, \tau)(\boldsymbol{r}, t)=$ $(\boldsymbol{r}, t)$ gives the identity element, and then, putting it on the right of the (3.3), one solves in the primed parameters to get the inverse of $g(\boldsymbol{\phi}, \boldsymbol{a}, \boldsymbol{v}, \tau)$ :

$$
\begin{equation*}
e=g(0,0,0,0), \quad g^{-1}(\boldsymbol{\phi}, \boldsymbol{a}, \boldsymbol{v}, \tau)=g\left(-\boldsymbol{\phi},-R_{\phi}^{-1}(\boldsymbol{a}+\tau \boldsymbol{v}),-R_{\phi}^{-1} \boldsymbol{v},-\tau\right) \tag{3.4}
\end{equation*}
$$

Galilean group is connected ten dimensional Lie group. It is easy to show that when all but one arguments are set to 0 , one gets the following subgroups, which are themselves significant in further considerations:

$$
\begin{align*}
\mathrm{SO}(3, \mathbb{R}) & =\{g(\boldsymbol{\phi}, 0,0,0)\}, & & \text { Rotational group; }  \tag{3.5a}\\
\boldsymbol{T}_{3}^{\mathrm{S}} & =\{g(0, \boldsymbol{a}, 0,0)\}, & & \text { Spatial translational group; }  \tag{3.5b}\\
\boldsymbol{T}_{3}^{\mathrm{B}} & =\{g(0,0, \boldsymbol{v}, 0)\}, & & \text { Boost group; }  \tag{3.5c}\\
\boldsymbol{T}_{1}^{\mathrm{T}} & =\{g(0,0,0, \tau)\}, & & \text { Time translational group. } \tag{3.5d}
\end{align*}
$$

The first two subgroups are built of pure spatial transformations, and together they form the Euclidean group $\mathrm{E}(3)=\boldsymbol{T}_{3}^{\mathrm{S}} \wedge \mathrm{SO}(3, \mathbb{R})$ (semi-direct product, with translations as the invariant subgroup). The last subgroup of the time translations, i.e. the transformations connecting systems with different origin of the time scale, is purely related to time. The third one, subgroups of the boosts, which are transformations interrelating the instantly coinciding coordinate systems moving uniformly each with respect to another, involves both time and spatial coordinates.

Besides these transformations which are continual functions of the parameters, there are also discrete transformations of the spatial, time and total reversal, $\mathcal{I}_{\mathrm{S}}, \mathcal{I}_{\mathrm{T}}$ and $\mathcal{I}=$ $\mathcal{I}_{\mathrm{S}} \mathcal{I}_{\mathrm{T}}:$

$$
\begin{equation*}
\mathcal{I}_{\mathrm{S}}(\boldsymbol{r}, t)=(-\boldsymbol{r}, t), \quad \mathcal{I}_{\mathrm{T}}(\boldsymbol{r}, t)=(\boldsymbol{r},-t), \quad \mathcal{I}(\boldsymbol{r}, t)=(-\boldsymbol{r},-t) . \tag{3.6}
\end{equation*}
$$

Together with the continual transformations, they form extended Galilean group $\overline{\mathcal{G}}$, with $\mathcal{G}$ being its invariant subgroup (as the connected component of the identity) and three additional cosets: $\overline{\mathcal{G}}=\mathcal{G}+\mathcal{I}_{S} \mathcal{G}+\mathcal{I}_{\mathrm{T}} \mathcal{G}+\mathcal{I} \mathcal{G}$. It is important to note that there is no possibility to verify directly the space-time invariance with respect to these transformations, and therefore they are not a priori involved in the relativity principle.

### 3.2 Galilean transformations of the classical variables

The Galilean transformations are defined by the action in the space-time, and afterward the action on the momenta is derived. This action can be extended to the variables, i.e. functions on the phase space. In the space of variables Galilean group is realized by the canonical transformations, being themselves functions of the variables. This is important in the view of the forthcoming quantization procedure. To begin with, we introduce Poisson bracket action of one variable onto another: to each variable $G$ corresponds the supervariable $\tilde{G}$, which is the linear operator (in the space of variables) acting on the arbitrary variable $A$ as

$$
\begin{equation*}
\tilde{G} A \stackrel{\text { def }}{=}[G, A]_{\mathrm{PZ}} \tag{3.7}
\end{equation*}
$$

Consequently, $\tilde{G}^{n} A=\left[G,\left[G, \ldots[G, A]_{\mathrm{PZ}} \ldots\right]_{\mathrm{PZ}}\right]_{\mathrm{PZ}}$. Therefore, to each analytic function $F(G)$ of $G$ the supervariable $\tilde{F}(G)$ may be associated using Taylor expansion of $F$. Par-
ticularly, the exponential function

$$
\begin{equation*}
\mathrm{e}^{\gamma \tilde{G}} \stackrel{\text { def }}{=} \sum_{n=0}^{\infty} \frac{\gamma^{n}}{n!} \tilde{G}^{n} \tag{3.8}
\end{equation*}
$$

defines one-parameter group of the transformations associated to $G$. Therefore, $\tilde{G}$ is the Lie-algebra generator of this group.

The Galilean transformations are in the variable space represented by the following supervariables:

$$
\begin{align*}
\tilde{D}\left(g(\boldsymbol{\phi}, 0,0,0)=\mathrm{e}^{-\boldsymbol{\phi} \cdot \tilde{\boldsymbol{l}}},\right. & \tilde{D}\left(g(0, \boldsymbol{a}, 0,0)=\mathrm{e}^{-\boldsymbol{a} \cdot \tilde{\boldsymbol{p}}}\right.  \tag{3.9a}\\
\tilde{D}\left(g(0,0, \boldsymbol{v}, 0)=\mathrm{e}^{-\boldsymbol{m} \cdot \tilde{\boldsymbol{q}}},\right. & \tilde{D}\left(g(0,0,0, \tau)=\mathrm{e}^{-\tau \tilde{H}}\right. \tag{3.9b}
\end{align*}
$$

This is verified by the action on the fundamental variables, i.e. in the phase space:

$$
\begin{align*}
\mathrm{e}^{-\phi \cdot \tilde{\boldsymbol{l}}} \boldsymbol{r}=R_{\phi} \boldsymbol{r}, & \mathrm{e}^{-\phi \cdot \tilde{\boldsymbol{l}}} \boldsymbol{p}=R_{\phi} \boldsymbol{p} ;  \tag{3.10a}\\
\mathrm{e}^{-a \cdot \tilde{p}} \boldsymbol{r}=\boldsymbol{r}+\boldsymbol{a}, & \mathrm{e}^{-a \cdot \tilde{p}} \boldsymbol{p}=\boldsymbol{p} ;  \tag{3.10b}\\
\mathrm{e}^{m \boldsymbol{v} \cdot \tilde{\boldsymbol{q}}} \boldsymbol{r}=\boldsymbol{r}, & \mathrm{e}^{m \boldsymbol{v} \cdot \tilde{\boldsymbol{q}} \boldsymbol{p}=\boldsymbol{p}-m \boldsymbol{v} ;}  \tag{3.10c}\\
\mathrm{e}^{\tau \tilde{H}} \boldsymbol{r}=\boldsymbol{r}+\frac{\tau}{m} \boldsymbol{p}, & \mathrm{e}^{-\tau \tilde{H}} \boldsymbol{p}=\boldsymbol{p} . \tag{3.10d}
\end{align*}
$$

In the last equations the hamiltonian of the free particle $H=\boldsymbol{p}^{2} / 2 m$ is assumed, since the particle in the empty space-time is only considered. Note that all the transformations are canonical. Therefore the Lie algebra generators are angular momentum $\boldsymbol{l}=\boldsymbol{r} \times \boldsymbol{p}$ for rotations, linear momentum $\boldsymbol{p}$ for translations, coordinates $\boldsymbol{q}$ for boosts and Hamiltonian for time translations. Note that (3.10) also defines the action on any other classical variable $A(\boldsymbol{r}, \boldsymbol{p})$ : on the functions of the fundamental variables action is defined by (3.7), which gives:

$$
\begin{equation*}
\tilde{D}(g(\boldsymbol{\phi}, \boldsymbol{a}, \boldsymbol{v}, \tau)) A(\boldsymbol{r}, \boldsymbol{p})=A(\tilde{D}(g(\boldsymbol{\phi}, \boldsymbol{a}, \boldsymbol{v}, \tau)) \boldsymbol{r}, \tilde{D}(g(\boldsymbol{\phi}, \boldsymbol{a}, \boldsymbol{v}, \tau)) \boldsymbol{p}) \tag{3.11}
\end{equation*}
$$

Exercise 3.1: Prove (3.9) in the coordinate form.
Note that although spatial translations and boosts commute, the supervariables corresponding to them do not. Therefore, the obtained set of supervariables is not a representation of $\mathcal{G}$, despite the fact that each subgroup is correctly represented. In fact, there is no faithful representation of the Galilean group. Since the faithful representations, relating transformations and supervariables bijectively are necessary, only the Euclidean subgroup and time translations can be simultaneously considered. While the pure geometrical invariance results in the conservation laws of the generators (angular and linear momentum), the time translations are manifested as the conservation of energy. On the other side, the boost invariance leads to the low of uniform motion of the centre of mass of the isolated system (since the kinetic energy is not the same in the boosted systems, the action is also changed, and the Nöther theorem is not applicable [?]).

### 3.3 Quantization: Wigner theorem

Being classical space-time symmetry, Galilean group is symmetry group of a single (nonrelativistic probe) particle, or of a centre of mass of any isolated system. Therefore, observable characteristics of such a system must be same before and after Galilean transformation, i.e. such characteristics are independent on the chosen inertial reference frame.

In the context of quantum mechanics, the observable quantities are various probabilities, which can be reduced to squares of absolute value of scalar products (recall (1.4)). In other words, if Galilean transformation $T$ transform states $|\phi\rangle$ and $|\psi\rangle$ into $\left|\phi^{\prime}\right\rangle$ and $\left|\psi^{\prime}\right\rangle$, then the space-time symmetry requires that

$$
\left.|\langle\phi \mid \psi\rangle|=\left|\left\langle\phi^{\prime} \mid \psi^{\prime}\right\rangle\right|=\left|\langle\phi| T^{\dagger} T\right| \psi\right\rangle \mid .
$$

Consequently, $T$ is in the state space represented by an operator $D(T)$ preserving absolute value of the scalar products. The famous Wigner theorem shows that such operators are either unitary or antiunitary.

Note that without absolute value in the above relation, only unitary operators would be allowed. In addition, due the absolute value together with $D(T)$ any operator $\mathrm{e}^{\mathrm{i} \chi_{T}} D(T)$ can be used equally well. This reflects the projective structure of the state space, i.e. that the collinear vectors describe the same state. One can try to use the freedom in the choice of the phases $\mathrm{e}^{\mathrm{i} \chi_{T}}$ to get the set of the unitary and antiunitary operators forming a representation of the Galilean group. However, this attempt fails, as well as in the classical physics, because Galilean group has no faithful representation. Still, all the important subgroups may be independently quantizied, and Euclidean group can be combined with time translations.

Although Wigner theorem does not specify which operators are represented by unitary or antiunitary operators, it is obvious by the continuity argument that the connected component of the identical element must be represented unitary. On the other hand, product of two antiunitary operators is unitary, meaning that either the whole group or its halving subgroup is unitary represented. Thus, at most two cosets are antiunitary.

Really, the complex conjugated Schrödinger equation (for a particle in the time independent potential) corresponds to the time reversed one with conjugated vectors, meaning that the time reversal is antiunitary in the state space. As the spatial inversion is obviously linear, it follows that purely spatial halving subgroup is linear, while time reversal and total inversion give the cosets having antilinear representations.

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### 3.4 Quantization of the Galilean group

Since the action of the Galilean transformation in the variable space is well defined in terms of Poison bracket and variables, the postulate of quantization may be directly applied to get the Galilean action in the space of the operators (superspace). Analogously to the supervariables, we define the superoperators with help of the commutator. So, to each operator $\hat{G}$ in the state space $\mathcal{H}$, its superoperator $\hat{G}$ is defined by its action on the
arbitrary operator $\hat{A}$ :

$$
\begin{equation*}
\hat{\hat{G}} \hat{A} \stackrel{\text { def }}{=}[\hat{G}, \hat{A}] . \tag{3.12}
\end{equation*}
$$

Due to the appearance of the term i $\hbar$ in the canonical quantization, the exponential function $\mathrm{e}^{\gamma \tilde{G}}$ is quantized as $\mathrm{e}^{\frac{\gamma}{\hbar}} \hat{G}$. Indeed, $\mathrm{e}^{\frac{\gamma}{\hbar}} \hat{G} \hat{A}=\sum_{n=0}^{\infty} \frac{\gamma^{n}}{n!} \frac{1}{\mathrm{i} \hbar}\left[\hat{G}, \frac{1}{\mathrm{i} \hbar}\left[\hat{G}, \ldots, \frac{1}{\mathrm{i} \hbar}[\hat{G}, \hat{A}] \ldots\right]\right]$, which is obviously the quantized exponential supervariable action.

Therefore, the superoperator representation of the Galilean group is directly obtained from (3.9):

$$
\begin{array}{cl}
\hat{\hat{D}}\left(g(\boldsymbol{\phi}, 0,0,0)=\mathrm{e}^{-\frac{1}{i \hbar} \phi \hat{\hat{\boldsymbol{l}}}},\right. & \hat{\hat{D}}\left(g(0, \boldsymbol{a}, 0,0)=\mathrm{e}^{-\frac{1}{\mathrm{i} \hbar} \boldsymbol{a} \cdot \hat{\hat{\boldsymbol{p}}}}\right. \\
\hat{\hat{D}}\left(g(0,0, \boldsymbol{v}, 0)=\mathrm{e}^{-\frac{1}{\mathrm{i} \hbar} m \boldsymbol{v} \cdot \hat{\boldsymbol{q}}},\right. & \hat{\hat{D}}\left(g(0,0,0, \tau)=\mathrm{e}^{-\frac{1}{i \hbar} \tau \hat{H}} .\right. \tag{3.13b}
\end{array}
$$

These superoperators are bijective image of the Galilean group, but the same problem as for supervariables appears: representatives of the boosts and translations do not commute as they should. Again, the Euclidean group and the time translations only can simultaneously quantized.

The superoperator representation is not really appropriate for the usual tasks of quantum mechanics, and the operator representation is looked for to realize the group action in the state space. The Baker-Hausdorff lemma solves this task. It asserts that there is bijection between the exponential superoperators and exponential operators, i.e. to each exponential action $\mathrm{e}^{\hat{\hat{G}}}$ in the operator space uniquely corresponds exponential action $\mathrm{e}^{\hat{G}}$ in the state space, such that superoperator essentially acts as the similarity transformation:

$$
\begin{equation*}
\mathrm{e}^{\hat{\hat{G}}} \hat{A}=\mathrm{e}^{\hat{G}} \hat{A} \mathrm{e}^{-\hat{G}} . \tag{3.14}
\end{equation*}
$$

Therefore in the state space, the Galilean transformations are represented by:

$$
\begin{array}{cl}
\hat{D}\left(g(\boldsymbol{\phi}, 0,0,0)=\mathrm{e}^{\frac{1}{i \hbar} \boldsymbol{\phi} \cdot \hat{\boldsymbol{l}}},\right. & \hat{D}\left(g(0, \boldsymbol{a}, 0,0)=\mathrm{e}^{\frac{1}{\hbar} \boldsymbol{a} \cdot \hat{\boldsymbol{p}}},\right. \\
\hat{D}\left(g(0,0, \boldsymbol{v}, 0)=\mathrm{e}^{\frac{1}{i \hbar} m \cdot \boldsymbol{v} \cdot \hat{\boldsymbol{q}}},\right. & \hat{D}\left(g(0,0,0, \tau)=\mathrm{e}^{\frac{1}{i \hbar} \tau \hat{H}} .\right. \tag{3.15b}
\end{array}
$$

Note that the result for the translations is known from the chapter 1. Indeed, the operators $\hat{D}(g(0, \boldsymbol{a}, 0,0)$ representing translations are direct three-dimensional generalization of the translational operator $U(a)$ used in the Theorem 1.2.

Exercise 3.2: Find $e^{\mathrm{i} \pi I} \psi(\boldsymbol{r})$, where $I$ is spatial inversion, i.e. $I \psi(\boldsymbol{r})=\psi(-\boldsymbol{r})$.

Exercise 3.3: Find the projectors $P_{ \pm}$on the subspaces spanned by even $\left(P_{+}\right)$and odd ( $P_{-}$) functions according to the spatial inversion.

Exercise 3.4: Find the coefficients $\alpha$ and $\beta$ for which the operator $U_{\alpha \beta}=e^{\alpha x^{2}+\beta p^{2}}$ becomes spatial inversion, i.e. rotation for $\pi$ in the phase space.

### 3.5 Active and passive interpretations

The Galilean transformations are defined by their action on the space-time points. In this sense, they actively map one point to another, and therefore this interpretation is called active interpretation. On the other hand, there is passive interpretation, describing the same act as the inverse transformation of the referent frame, with no action on the space-time points. Thus, to each active transformation $T$ on the space time-vectors represented in the reference frame $\left\{\boldsymbol{e}_{1}, \boldsymbol{e}_{2}, \boldsymbol{e}_{3}, \boldsymbol{e}_{t}\right\}$ by the column $(\boldsymbol{r}, t)$, corresponds the inverse transformation of the frame:

$$
\begin{equation*}
\text { Active : } T\binom{\boldsymbol{r}}{t} \rightarrow\binom{\boldsymbol{r}^{\prime}}{t^{\prime}} ; \quad \text { Passive : } T^{-1} \boldsymbol{e}_{i} \rightarrow \boldsymbol{e}_{i}^{\prime} \tag{3.16}
\end{equation*}
$$

such that in the transformed frame the initial point is described as $\left(\boldsymbol{r}^{\prime}, t^{\prime}\right)$. Obviously, some of the transformations cannot be really realized actively, e.g. time translations, and reversal transformations. Also the boosts are even defined with help of the reference systems. Due to the inversion of the transformation when one comes from active to the passive interpretation, they are anti-isomorphic. Acting on the variables, being the functions on the phase space, the supervariables (3.9) are essentially passive interpretation, and therefore the sign minus appears in the exponents.

Finally, from the action $\hat{D}(g)|\boldsymbol{q}\rangle=|g \boldsymbol{q}\rangle$, it follows that $\langle\boldsymbol{q}| \hat{D}(g)|\psi\rangle=\left\langle g^{-1} \boldsymbol{q} \mid \psi\right\rangle$. Thus, coordinate representation of the Galilean transformations is

$$
\begin{equation*}
\hat{D}(g) \psi(\boldsymbol{q})=\psi\left(g^{-1} \boldsymbol{q}\right) \tag{3.17}
\end{equation*}
$$



Figure 3.1: Active and passive interpretions of transformation. Left column (kvadrati I, III i V i izomorfizmi $\iota_{2}$ i $\iota_{6}$ ) daje elemente za pasivnu, a desna kolona (kvadrati II, IV i VI i izomorfizmi $\iota_{4}$ i $\iota_{7}$ ) za aktivnu interpretaciju grupa $E(3)$ i $T_{3}^{(v)}$. Antiizomorfizmi $\alpha_{1}$, $\alpha_{3}, \alpha_{5}$ i $\alpha_{8}$, kao i izomorfizmi $\iota_{2}, \iota_{4}, \iota_{6}$ i $\iota_{7}$ povezuju osnovne podgrupe Galilejeve grupe koje deluju u skupovima označenim sa I do VI. Specijalno: $\iota_{2} \stackrel{\text { def }}{=} \alpha_{5} \circ \alpha_{3}=\alpha_{5} \circ \iota_{4} \circ \alpha_{1}^{-1}$; $\alpha_{3}=\iota_{4} \circ \iota_{1}$ je prevoenje vienja izmeu dva opservera; $\iota_{4}$ je reprezentovanje u fiksiranom koordinatnom sistemu sa implikacijama u impulsnom prostoru.

## Chapter 4

## Rotations and Angular Momentum

### 4.1 Elementary Properties of Rotations

As explained in the Chapter 3, rotations are parameterized by the vectors $\phi$, specifying the angle $\phi$ and the axis (by the unit vector $\hat{\boldsymbol{\phi}}=\boldsymbol{\phi} / \phi$ ) of each rotation. The defining action of the rotations in the Euclidean space $\mathbb{R}^{3}$ is given by the special orthogonal matrices, and therefore the rotational group is identified with the group $\mathrm{SO}(3, \mathbb{R})$.

The obvious equality $R_{\phi}=R_{(2 \pi-\phi)(-\hat{\phi})}$ of the rotations around opposite axes shows that all the rotations are obtained taking all the unit vectors and the angles not greater than $\pi$, i.e. that the range of the vectors $\phi$ (manifold of the parameters) is $\pi$-ball, a ball of radius $\pi$. However, the opposite points of the same diameter of this ball correspond to the same rotation: $R_{\pi \boldsymbol{u}}=R_{-\pi \boldsymbol{u}}$ for arbitrary unit vector $\boldsymbol{u}$. Hence, the rotational group is compact (parameter space is bounded and closed), but double connected, as any diameter of the $\pi$-ball is a closed path not homotopic to any closed path in the $\pi$-ball's interior.

Further, the first relation in (3.15), $\hat{D}\left(g(\boldsymbol{\phi}, 0,0,0)=\mathrm{e}^{\frac{1}{i \hbar} \phi \cdot \hat{l}}\right.$, shows that rotations are represented in the orbital state space by the exponents of the operators of the orbital angular momentum. Therefore the components of the angular momentum form the Lie algebra of the rotational group. Using the definition $\boldsymbol{l}=\boldsymbol{r} \times \boldsymbol{p}$, and the canonical coordinatesmomenta commutation relations $\left[q_{i}, p_{j}\right]=\mathrm{i} \hbar \delta_{i j}$, the commutation relations of the angular momenta are easily derived:

$$
\begin{equation*}
\left[l_{i}, l_{j}\right]=\mathrm{i} \hbar \sum_{k} \epsilon_{i j k} l_{k} \tag{4.1}
\end{equation*}
$$

( $\epsilon$ is Levy-Chivita tensor). This shows that the Lie algebra is so $(3, \mathbb{R})$ (of course, this follows immediately from the general theory), the set of the traceless hermitian operators ${ }^{1}$. Now, several general (and elementary) results of the group theory can be exploited to get additional insight into basic properties of the rotations and angular momenta.

[^3]First, compactness implies that rotational group is represented by unitary operators, while the angular momenta are hermitian (precisely, equivalent to the unitary and hermitian operators, or alternatively, that there is a scalar product - selected in physics - such that the representative operators are unitary and hermitian, respectively). As a consequence, each representation is decomposable to the irreducible ones, the later being necessarily finite dimensional.

The fact that the angular momenta form three-dimensional simple Lie algebra so( $3, \mathbb{R}$ ), enables to construct irreducible representation according to the well elaborated algorithm for the semisimple algebras (Section 4.2.1). Additionally, its rank is one, which reveals that there is a single Cartan element and a single Kazimir operator: conventionally, these are the third component, $l_{3}$, and the square, $l^{2}=l_{1}^{2}+l_{2}^{2}+l_{3}^{2}$, of the angular momentum). Therefore, irreducible representations are classified by a single parameter, different eigenvalues of the Kazimir operator, while a basis within arbitrary irreducible space is counted by the different eigenvalues of the Cartan element as another parameter.

Doubly connected parameter space implies that there is a double (universal) covering, locally isomorphic to $\mathrm{SO}(3, \mathbb{R})$, i.e. a simply connected Lie group $\widetilde{\mathrm{SO}(3, \mathbb{R})}$, with two-toone homomorphism onto $\mathrm{SO}(3, \mathbb{R})$, but with the same Lie algebra. As it will be shown latte, this is the group $\operatorname{SU}(2)$, with algebra $\mathrm{su}(2)=\mathrm{so}(3, \mathbb{R})$. Accordingly, the obtained representations of the algebra of the angular momenta do not really produce (as exponentials) the representations of the rotational but of the covering group. Only half of them represent rotations, while the rest are the "double valued" representations, associating a pair of the operators to each rotation; the characteristic feature of the double valued representation, differing them from the ordinary ones, is that rotations for $2 \pi$ reverse all the vectors (of the representative space), and only rotations for $4 \pi$ are associated to identity operators. A final consequence is that the notion of the rotations should be reconsidered: without any concrete physical content specified, a general concept is effectively defined by the commutation relations (4.1) only. This essentially introduces the covering group as that underlying the angular momenta, and not the rotations. The existence of the fermionic spin degrees of freedom (Section 4.5), which can be explained only by these double-valued representations, proves that this is not just a mathematical extrapolation.

### 4.2 Algebra of angular momentum

To stress out that from now on the notion of the angular momentum is as general as possible, free of any specific physical situation, the angular momenta are defined as a triple of the hermitian operators $K_{i}(i=1,2,3)$ satisfying the commutation relations:

$$
\begin{equation*}
\left[K_{i}, K_{j}\right]=\mathrm{i} \hbar \sum_{k} \epsilon_{i j k} K_{k} . \tag{4.2}
\end{equation*}
$$

In this way only the Lie algebra properties are extracted from the initial (prototypic) example, the orbital angular momentum. Indeed, only this is really necessary for the following construction of the representation (neatly following the prescription developed for the semisimple complex Lie algebras).

### 4.2.1 Irreducible representations of angular momentum

Construction of the irreducible representations starts by choice of a suitable basis (in mathematical language, this corresponds to the standard form of the complexified algebra). Mutually adjoint complex linear combinations

$$
\begin{equation*}
K_{ \pm}=K_{1} \pm \mathrm{i} K_{2} \tag{4.3}
\end{equation*}
$$

of $K_{1}$ and $K_{2}$, together with $K_{3}$ form standard components of angular momentum. They span the same space as the three momenta, since (4.3) is solvable in the Cartesian components: $K_{1}=\frac{1}{2}\left(K_{+}+K_{-}\right)$and $K_{2}=\frac{1}{2 i}\left(K_{+}-K_{-}\right)$. Therefore, space is irreducible for the Cartesian components if and only if it is irreducible for the standard ones.

Lema 4.1 The standard components satisfy the following commutation relations:

$$
\begin{align*}
& {\left[K_{3}, K_{ \pm}\right]= \pm \hbar K_{ \pm}}  \tag{4.4a}\\
& {\left[K_{+}, K_{-}\right]=2 \hbar K_{3} ;}  \tag{4.4b}\\
& {\left[K_{3}, K_{+} K_{-}\right]=\left[K_{3}, K_{-} K_{+}\right]=\left[K_{+} K_{-}, K_{-} K_{+}\right]=0} \tag{4.4c}
\end{align*}
$$

-Proof: (a) and (b) are obtained applying (4.3) and then (4.2).
(c) $\left[K_{3}, K_{ \pm} K_{\mp}\right]=\left[K_{3}, K_{ \pm}\right] K_{\mp}+K_{ \pm}\left[K_{3}, K_{\mp}\right]= \pm \hbar K_{ \pm} K_{\text {甲 }} \mp \hbar K_{ \pm} K_{\text {甲 }}=0$;
$\left[K_{+} K_{-}, K_{-} K_{+}\right]=K_{-}\left[K_{+} K_{-}, K_{+}\right]+\left[K_{+} K_{-}, K_{-}\right] K_{+}=$
$K_{-} K_{+}\left[K_{-}, K_{+}\right]+K_{-}\left[K_{+}, K_{+}\right] K_{-}+K_{+}\left[K_{-}, K_{-}\right] K_{+}+\left[K_{+}, K_{-}\right] K_{-} K_{+}=$ $K_{-} K_{+}\left(-2 \hbar K_{3}\right)+2 \hbar K_{3} K_{-} K_{+}=0 . \quad$ I

The operators $K_{ \pm} K_{\mp}$ are hermitian (an positive), and together with $K_{3}$ make a set of compatible observables (as justified by (4.4c)). Thus, they have a common orthonormal eigenbasis. Number $m$ is used to denote eigenvalues $m \hbar$ of $K_{3}$ (the units of angular momentum and action are same). In the possibly degenerated eigensubspace $\mathcal{V}_{m}$ of $K_{3}$, a common (for the whole triple) eigenbasis is denoted by $|m \lambda\rangle\left(\lambda=1, \ldots,\left|\mathcal{V}_{m}\right|\right)$ :

$$
\begin{equation*}
K_{3}|m \lambda\rangle=m \hbar|m \lambda\rangle, \quad K_{ \pm} K_{\mp}|m \lambda\rangle=\alpha_{m}^{ \pm}(\lambda)|m \lambda\rangle, \tag{4.5}
\end{equation*}
$$

with $\alpha_{m}^{ \pm}(\lambda)$ being the eigenvalues of $K_{ \pm} K_{\mp}$.
Lema 4.2 For each $|m \lambda\rangle$ the vectors $K_{ \pm}|m \lambda\rangle$ are mutually orthogonal for different $\lambda$, and satisfy

$$
\begin{align*}
& K_{3}\left(K_{ \pm}|m \lambda\rangle\right)=(m \pm 1) \hbar\left(K_{ \pm}|m \lambda\rangle\right),  \tag{4.6a}\\
& K_{ \pm} K_{\mp}\left(K_{ \pm}|m \lambda\rangle\right)=\alpha_{m}^{\mp}(\lambda)\left(K_{ \pm}|m \lambda\rangle\right),  \tag{4.6b}\\
& K_{ \pm} K_{\mp}\left(K_{\mp}|m \lambda\rangle\right)=\left(\alpha_{m}^{ \pm}(\lambda) \pm 2 \hbar^{2}(m \mp 1)\right)\left(K_{\mp}|m \lambda\rangle\right) . \tag{4.6c}
\end{align*}
$$

-Proof: All the relations follows from (4.5). To show orthogonality and (4.6b) one regroups the terms:

$$
\begin{gathered}
\left(\langle m \lambda| K_{ \pm}^{\dagger}\right)\left(K_{ \pm}\left|m \lambda^{\prime}\right\rangle\right)=\langle m \lambda|\left(K_{\mp} K_{ \pm}\left|m \lambda^{\prime}\right\rangle\right)=\alpha_{m}^{\mp}(\lambda)\left\langle m \lambda \mid m \lambda^{\prime}\right\rangle=\alpha_{m}^{\mp}(\lambda) \delta_{\lambda \lambda^{\prime}} ; \\
\left(K_{ \pm} K_{\mp}\right)\left(K_{ \pm}|m \lambda\rangle\right)=K_{ \pm}\left(K_{\mp} K_{ \pm}\right)|m \lambda\rangle=\alpha_{m}^{\mp}(\lambda)\left(K_{ \pm}|m \lambda\rangle\right) .
\end{gathered}
$$

For (4.6a) and (4.6c) it suffices to apply (4.4a) and (4.4c), respectively

$$
K_{3}\left(K_{ \pm}|m \lambda\rangle\right)=\left(\left[K_{3}, K_{ \pm}\right]+K_{ \pm} K_{3}\right)|m \lambda\rangle=( \pm 1+m) \hbar\left(K_{ \pm}|m \lambda\rangle\right) ;
$$

$$
\begin{aligned}
K_{ \pm} K_{\mp}\left(K_{\mp}|m \lambda\rangle\right) & =\left(\left[K_{ \pm}, K_{\mp}\right]+K_{\mp} K_{ \pm}\right)\left(K_{\mp}|m \lambda\rangle\right)= \\
& =\left( \pm 2 \hbar K_{3}+K_{\mp} K_{ \pm}\right)\left(K_{\mp}|m \lambda\rangle\right)=\left( \pm 2 \hbar^{2}(m \mp 1)+\alpha_{m}^{ \pm}(\lambda)\right)\left(K_{\mp}|m \lambda\rangle\right) .
\end{aligned}
$$

(in the last equation (4.6a) is used).
The lema enlightens that the vectors of the basis $|m \lambda\rangle$ of $\mathcal{V}_{m}$ are transferred by $K_{ \pm}$ into $\mathcal{V}_{m \pm 1}$, but remain eigenvectors of $K_{ \pm} K_{\mp}$ for the fixed $\lambda$; still, some of them vanish, since the eigenequations (4.6) are trivially solved by the zero vector. As the rotation group is compact, all its irreducible representations are finite dimensional, and there are maximal $k^{+}$and minimal $k^{-}$values of $m$ for which $K_{ \pm}\left|k^{ \pm} \lambda\right\rangle=0$. Thus, the subspace $\mathcal{V}_{\lambda}^{\left(k^{+}, k^{-}\right)}$spanned by the vectors $\left\{|m \lambda\rangle \mid m=k^{-}, \ldots, k^{+}\right\}$is invariant and irreducible for $K_{z}$ and $K_{ \pm}$, i.e. for the angular momentum.

Theorem 4.1 All the irreducible representations of angular momentum are $D^{(k)}(\boldsymbol{K})$ with the maximal weight

$$
\begin{equation*}
k=k^{+}=-k^{-}=0, \frac{1}{2}, 1, \frac{3}{2}, \ldots \tag{4.7}
\end{equation*}
$$

They are $2 k+1$-dimensional, and in the standard basis (4.5), denoted as $|k m\rangle$ with $m=$ $-k, \ldots, k$, of the irreducible space $\mathcal{V}^{(k)}$, the matrices $D^{(k)}(\boldsymbol{K})$ are determined by:

$$
\begin{align*}
& K_{3}|k m\rangle=m \hbar|k m\rangle, \quad K_{ \pm}|k m\rangle=c_{m}^{(k)^{ \pm}} \hbar|k, m \pm 1\rangle  \tag{4.8a}\\
& c_{m}^{(k)^{ \pm}}=\sqrt{k(k+1)-m(m \pm 1)} \tag{4.8b}
\end{align*}
$$

- Proof: According to the previous results, in the single irreducible subspace $K_{3}$ is nondegenerate, making $\lambda$ superfluous, and the standard basis is $|m\rangle$, with $m=k^{-}, \ldots, k^{+}$. A direct consequence of (4.6a) are equations

$$
K_{3}|m\rangle=m \hbar|m\rangle, \quad K_{ \pm}|m\rangle=c_{m}^{ \pm} \hbar|m \pm 1\rangle,
$$

having the form of (4.8a). The coefficients $c_{m}^{ \pm}$are to be still determined in accordance with the previously assumed orthonormality. Applying $K_{+}=K_{-}^{\dagger}$ to act by $K_{+}$on the left and on the right in $\langle m+1| K_{+}|m\rangle$, we get $c_{m+1}^{-*}=c_{m}^{+}$. Further, we already proved that there are minimal, $k^{-}$and maximal, $k^{+}$, value of $m$, for which $K_{ \pm}\left|k^{ \pm}\right\rangle=0$, allowing to assume $c_{k^{ \pm}}^{ \pm}=0$. Then we use (4.4a), to find

$$
2 \hbar\langle m| K_{3}|m\rangle=\langle m|\left[K_{+}, K_{-}\right]|m\rangle=\langle m| K_{+} K_{-}-K_{-} K_{+}|m\rangle=\| K_{-}|m\rangle\left\|^{2}-\right\| K_{+}|m\rangle \|^{2} .
$$

In the view of (4.8a), this reads:

$$
\begin{equation*}
2 m=\left|c_{m}^{-}\right|^{2}-\left|c_{m}^{+}\right|^{2} \tag{4.9}
\end{equation*}
$$

Now we sum these equations for all values of $m$, and find:

$$
2 \sum_{m=k^{-}}^{k^{+}} m=\sum_{m=k^{-}}^{m=k^{+}}\left(\left|c_{m}^{-}\right|^{2}-\left|c_{m}^{+}\right|^{2}\right) .
$$

When we substitute the equalities $\left|c_{m+1}^{-}\right|=\left|c_{m}^{+}\right|$and $c_{k^{ \pm}}^{ \pm}=0$ found previously, the right hand side vanishes (Fig. 4.1), and we get

$$
\sum_{m=0}^{k^{+}-k^{-}}\left(k^{-}+m\right)=\frac{\left(k^{+}+k^{-}\right)\left(k^{+}-k^{-}+1\right)}{2}=0 .
$$

As $k^{+} \geq k^{-}$, this implies $k^{-}=-k^{+}$. Finally, the difference $k^{+}-k^{-}=2 k^{+}$is an integer, meaning that $k=k^{+}=0, \frac{1}{2}, \ldots$, and $\left|\mathcal{V}^{(k)}\right|=2 k+1$. Analogously, summing (4.9) from $m$ to $k=k^{+}$, we find:

$$
\left|c_{m}^{-}\right|^{2}=\left|c_{m-1}^{-}\right|^{2}=2 \sum_{s=m}^{k} s=2\left(\sum_{s=1}^{k} s-\sum_{s=1}^{m-1} s\right)=k(k+1)-m(m-1),
$$

which is, with a convention on the (common) phase, the value in the theorem.

$$
\begin{array}{rlcrl}
\left|k^{+}\right\rangle & \bullet & 2 k^{+} & =\left|c_{k+1}^{-}\right|^{2}-\left|c_{k+}^{+}\right|^{2} \\
\left|k^{+}-1\right\rangle & \circ & 2\left(k^{+}-1\right) & =\left|c_{k^{+-1}}^{-}\right|^{2}-\left|c_{k^{+-1}}^{+}\right|^{2} \\
\left|k^{+}-2\right\rangle & \circ & 2\left(k^{+}-2\right) & =\left|c_{k^{+}-2}^{-}\right|^{2}-\left|c_{k^{+-2}}^{+}\right|^{2} \\
& \vdots & & \\
|m+1\rangle & \circ & 2(m+1) & =\left|c_{m+1}^{-}\right|^{2}-\left|c_{m+1}^{+}\right|^{2} \\
|m\rangle & \circ & 2 m & =\left|c_{m}^{-}\right|^{2}-\left|c_{m}^{+}\right|^{2} \\
& \vdots & & \\
\left|k^{-}+1\right\rangle & \circ & 2\left(k^{-}+1\right) & =\left|c_{k^{-}+1}^{-}\right|^{2}-\left|c_{k^{-}+1}^{+}\right|^{2} \\
\left|k^{-}\right\rangle & \bullet & 2 k^{-} & =\left|c_{k}^{-}\right|^{2}-\left|c_{k^{-}}^{+}\right|^{2}
\end{array}
$$

Figure 4.1: Standard basis of an irreducible representation of the angular momenta. For each standard vector $|m\rangle$ (denoted by $\circ$, or • for maximal and minimal weight) the corresponding equation (4.9) is written, and the terms which are canceled in the course of summation are singled out by the gray lines.

Exercise 4.1: Show that the matrices of the angular momenta in the standard basis are:

$$
D^{(k)}\left(K_{3}\right)=\hbar\left[\begin{array}{ccccc}
-k & 0 & 0 & 0 & 0  \tag{4.10}\\
0 & -k+1 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & \frac{0}{0} & 0 \\
0 & 0 & 0 & 0 & k
\end{array}\right], D^{(k)}\left(K_{+}\right)=\hbar\left[\begin{array}{cccccc}
0 & 0 & 0 & 0 & 0 & 0 \\
\sqrt{2 k} & 0 & 0 & 0 & 0 & 0 \\
0 & \sqrt{4 k-2} & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & & 0 \\
0 & 0 & 0 & 0 & 0
\end{array}\right], D^{(k)}\left(K_{-}\right)=D^{(k)^{\dagger}}\left(K_{+}\right) .
$$

The matrices representing rotations can be now easily find from the first equation in (3.15). In particular, a rotation around $z$-axis is represented as

$$
D^{(k)}\left(R_{\phi \boldsymbol{e}_{z}}\right)=\mathrm{e}^{-\frac{\mathrm{i}}{\hbar} \phi D^{(k)}\left(K_{3}\right)}=\operatorname{diag}\left[\mathrm{e}^{\mathrm{i} k \phi}, \mathrm{e}^{\mathrm{i}(k-1) \phi}, \ldots, \mathrm{e}^{-\mathrm{i} k \phi}\right] .
$$

It is important to note that $D^{(k)}\left(R_{2 \pi e_{z}}\right)=( \pm)^{2 k} I_{2 k+1}$, i.e. only for an integer $k$ this is the identity, while this is the case for the rotations for multiples of $4 \pi$ in the case of half-integral $k$. As mentioned in Section 4.1, this manifests the fact that the presented construction gives the representation of the Lie algebra of of rotational group, and not of the group itself. Therefore the exponential map of these representations coincides with the representations of the universal covering group $\mathrm{SU}(2)$ of the rotational group. Essentially, this is the consequence of the nontrivial fundamental group of the rotational group $\mathrm{SO}(3)$. However, this seemingly artifact of the construction, is a substantial physical property of the particles, as it will be seen in the discussion of the interior degrees of freedom, like spin (Section 4.5).

### 4.2.2 Square of the angular momentum

Square of angular momentum is positive observable

$$
\begin{equation*}
K^{2}=K_{1}^{2}+K_{2}^{2}+K_{3}^{2} . \tag{4.11}
\end{equation*}
$$

It is easy to show that $K^{2}$ commutes with all the components $K_{i}$ of angular momentum. Therefore, eigensubspaces of $K^{2}$, being invariant under all $K_{i}$, contain complete irreducible subspaces. Thus each irreducible subspace $\mathcal{V}^{(k)}$ is a subspace of an eigensubspace of $K^{2}$, meaning that in $\mathcal{V}^{(k)}$ operator $K^{2}$ acts as a multiplication constant. To find out this constant, we simply act on an arbitrary vector $|k m, \lambda\rangle$ by $K^{2}$, written in the more convenient form:

$$
\begin{equation*}
K^{2}=\frac{1}{2}\left(K_{+} K_{-}+K_{-} K_{+}\right)+K_{3}^{2} \tag{4.12}
\end{equation*}
$$

Then, using the results of the Theorem 4.1, we find

$$
\begin{equation*}
K^{2}|k m, \lambda\rangle=k(k+1) \hbar^{2}|k m, \lambda\rangle . \tag{4.13}
\end{equation*}
$$

Exercise 4.2: Show (4.13).
Thus the eigensubspaces of $K^{2}$ are multiple irreducible subspaces, i.e. the sum of the irreducible subspaces of the same maximal weight $k$. This fact enables to find standard basis in any space where the operators of the angular momentum are defined as the common eigenvectors of $K_{3}$ and $K^{2}$, i.e. of the system of eigenproblems:

$$
\begin{equation*}
K^{2}|k m, \lambda\rangle=k(k+1) \hbar^{2}|k m, \lambda\rangle, \quad K_{3}|k m, \lambda\rangle=m \hbar|k m, \lambda\rangle . \tag{4.14}
\end{equation*}
$$

This is the most efficient way to determine standard basis in the state space, which will be systematically used later on.

### 4.2.3 Discussion

Operators of the angular momentum are related to the rotations by the relations (3.15). Essentially, this means that the angular momentum components span the Lie algebra of the rotational group. Therefore, the method used in the construction of the irreducible representations is the usual prescription in the theory of semisimple Lie algebras. In particular, the first two equation of (4.4) are the so called standard form of such an algebra, with $K_{3}$ being the only Cartan vector, while $K_{ \pm}$are roots. Analogously, (4.6a) manifests the well known relation between the roots and the weights. Finally, as the algebra is of rank 1 , it has a single independent Kazimir operator, which is $K^{2}$.

It is useful to illustrate some differences in the physical contents of the classical and quantum angular momentum. As usual, the source of them is commutativity of the classical physical variables, in contrast to their quantum analogues.

From the uncertainty relations (4.2) one directly gets:

$$
\begin{equation*}
\Delta K_{i} \Delta K_{j} \geq \frac{\hbar}{2} \sum_{k} \epsilon_{i j k}\left|\left\langle K_{k}\right\rangle\right| . \tag{4.15}
\end{equation*}
$$

Therefore, unless the average of one component of the angular momentum in the given state vanishes, the dispersions of the other components are nonzero.

Second, if the state $|x\rangle$ is non-dispersive for $K_{i}$, i.e. $\Delta K_{i}=0$ meaning that it is an eigenstate of $K_{i}$, then from (4.15) follows that averages of both other components of angular momentum vanish.

Finally, in the classical mechanics only the component of the angular momentum along the axis of rotation is nonzero; those in the perpendicular directions vanish, as well as their squares. On the contrary, in quantum mechanics the averages $\left\langle K_{i}^{2}\right\rangle$ and $\left\langle K_{j}^{2}\right\rangle$ vanish only in the state with vanishing total angular momentum, $|k=0, m=0, \lambda\rangle$. This shows that the mentioned criterion to determine rotational axis is not applicable in quantum mechanics.
Exercise 4.3: Show that $\left\langle K_{i}^{2}\right\rangle_{|x\rangle}=\left\langle K_{j}^{2}\right\rangle_{|x\rangle}=0$ if and only if for all the components $K_{p}|x\rangle=0$, i.e. $|x\rangle=|k=0, m=0, \lambda\rangle$.

### 4.3 Orbital Angular Momentum

### 4.3.1 Coordinate representation

After general considerations we turn back to the original definition of angular momentum, i.e. to the orbital angular momentum $\boldsymbol{l}=\boldsymbol{r} \times \boldsymbol{p}$. Its Cartesian components $l_{i}=\sum_{j k} \epsilon_{i j k} r_{i} p_{j}$ in the coordinate representation, when momentum components $p_{i}=-\mathrm{i} \hbar \frac{\partial}{\partial r_{i}}$ are substituted, become:
$l_{1}=-\mathrm{i} \hbar\left(r_{2} \frac{\partial}{\partial r_{3}}-r_{3} \frac{\partial}{\partial r_{2}}\right), \quad l_{2}=-\mathrm{i} \hbar\left(r_{3} \frac{\partial}{\partial r_{1}}-r_{1} \frac{\partial}{\partial r_{3}}\right), \quad l_{3}=-\mathrm{i} \hbar\left(r_{1} \frac{\partial}{\partial r_{2}}-r_{2} \frac{\partial}{\partial r_{1}}\right)$.
The forthcoming considerations are most efficient performed in spherical coordinates. Also, angular momentum is particularly important (as conserved quantity) for the systems with spherical symmetry, for which spherical coordinates are natural choice. To find the spherical components of the angular momentum we use Cartesian components and apply coordinate transformations: $x=r \sin \theta \cos \varphi, y=r \sin \theta \sin \varphi, z=r \sin \theta \cos \varphi$. The Jacobian of this transformation relates the coordinate partial derivatives:

$$
\left(\begin{array}{l}
\frac{\partial}{\partial r}  \tag{4.17}\\
\frac{\partial}{\partial \theta} \\
\frac{\partial}{\partial \varphi}
\end{array}\right)=\left(\begin{array}{ccc}
\sin \theta \cos \varphi & \sin \theta \sin \varphi & \cos \theta \\
r \cos \theta \cos \varphi & r \cos \theta \sin \varphi & -r \sin \theta \\
-r \sin \theta \sin \varphi & r \sin \theta \cos \varphi & 0
\end{array}\right)\left(\begin{array}{l}
\frac{\partial}{\partial x} \\
\frac{\partial}{\partial x} \\
\frac{\partial}{\partial z}
\end{array}\right) .
$$

Its inverse is used to express the Cartesian derivatives in terms of spherical coordinates:

$$
\begin{align*}
\frac{\partial}{\partial x} & =\sin \theta \cos \varphi \frac{\partial}{\partial r}+\frac{1}{r} \cos \theta \cos \varphi \frac{\partial}{\partial \theta}-\frac{1}{r} \frac{\sin \varphi}{\sin \theta} \frac{\partial}{\partial \varphi} \\
\frac{\partial}{\partial y} & =\sin \theta \sin \varphi \frac{\partial}{\partial r}+\frac{1}{r} \cos \theta \sin \varphi \frac{\partial}{\partial \theta}+\frac{1}{r} \frac{\cos \varphi}{\sin \theta} \frac{\partial}{\partial \varphi}  \tag{4.18}\\
\frac{\partial}{\partial z} & =\cos \theta \frac{\partial}{\partial r}-\frac{1}{r} \sin \theta \frac{\partial}{\partial \theta}
\end{align*}
$$

Knowing this, Cartesian components are easily expressed in terms of spherical coordinates:

$$
\begin{align*}
& l_{x}=\mathrm{i} \hbar\left(\sin \varphi \frac{\partial}{\partial \theta}+\cot \theta \frac{\partial}{\partial \varphi}\right) \\
& l_{y}=-\mathrm{i} \hbar\left(\cos \varphi \frac{\partial}{\partial \theta}-\cot \theta \frac{\partial}{\partial \varphi}\right)  \tag{4.19}\\
& l_{z}=-\mathrm{i} \hbar \frac{\partial}{\partial \varphi}
\end{align*}
$$

We also find $l_{ \pm}$and $l^{2}$ as general theory emphasized their importance:

$$
\begin{equation*}
l_{ \pm}=\hbar \mathrm{e}^{ \pm \mathrm{i} \varphi}\left( \pm \frac{\partial}{\partial \theta}+\mathrm{i} \cot \theta \frac{\partial}{\partial \varphi}\right), \quad l^{2}=-\hbar^{2}\left[\frac{1}{\sin ^{2} \theta} \frac{\partial^{2}}{\partial \varphi^{2}}+\frac{1}{\sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial}{\partial \theta}\right)\right] \tag{4.20}
\end{equation*}
$$

All the considered operators act in the Lebesgue space $\mathcal{L}\left(\mathbb{R}^{3}\right)$ of the single three dimensional particle, which is the product of the three Lebesgue spaces over Cartesian coordinates: $\mathcal{L}\left(\mathbb{R}^{3}\right)=\mathcal{L}\left(\mathbb{R}_{x}\right) \otimes \mathcal{L}\left(\mathbb{R}_{y}\right) \otimes \mathcal{L}\left(\mathbb{R}_{z}\right)$. On the other side, this space is seemingly product of the Lebesgue spaces over the spherical coordinates. It would be very convenient to retain this product structure in spherical coordinate system, although due to the singularities of this system this is not completely correct (e.g., functions $f(r, \theta, \varphi)$ are acceptable only if for $r=0$ they are independent of $\theta$ and $\varphi$, i.e. if $f(0, \theta, \varphi)$ is constant; otherwise the value in coordinate origin is undefined). So, these deficiencies are cured by extra conditions on the functions $f(r, \theta, \varphi)$ :

$$
\mathcal{L}\left(\mathbb{R}^{3}\right)=\left\{\left.f(r, \theta, \varphi) \in \mathcal{L}\left((0, \infty)_{r}\right) \otimes \mathcal{L}\left((0, \pi)_{\theta}\right) \otimes \mathcal{L}\left((0,2 \pi)_{\varphi}\right)\right|_{\begin{array}{c}
f(0, \theta, \varphi)=\text { const, }  \tag{4.21}\\
f(r, \theta, \theta, \pi, \varphi)=f(r)=f(r, \theta, \varphi+2 \pi)
\end{array}} ^{f(r)}\right\}
$$

Note that none of the angular momentum operators (4.19) and (4.20) are $r$ dependent, and therefore they act as identity in the factor space $\mathcal{L}\left((0, \infty)_{r}\right)$. This reflects that the angular moment generates rotations, which do not affect the radius. A consequence is that the orbital angular momentum is effectively defined in the angular Lebesgue space of the functions over the sphere:

$$
\begin{equation*}
\mathcal{L}\left(S^{2}\right)=\left\{\left.f(\theta, \varphi) \in \mathcal{L}\left((0, \pi)_{\theta}\right) \otimes \mathcal{L}\left((0,2 \pi)_{\varphi}\right)\right|_{\substack{f(\theta, \varphi)=f(\theta, \varphi+2 \pi)}} ^{f(\theta=0, \pi, \varphi)=\text { const, }}\right\} . \tag{4.22}
\end{equation*}
$$

The last condition explicates $2 \pi$ periodicity of all the functions in the Lebesgue space. As a consequence, the orbital angular momentum has only integer irreducible components, as it will be explicated soon.

### 4.3.2 Standard Basis: Spherical Harmonics

According to the general theory, the standard basis is the common eigenbasis of $l_{z}$ and $l^{2}$. The ineffectiveness of these operators in the radial factor space allows to apply the theorem on the separation of variables (Appendix A.1), i.e. to consider this eigenproblem in $\mathcal{L}\left(S^{2}\right)$ only: the obtained eigenbasis $Y_{l}^{m}(\theta, \varphi)=\langle\theta, \varphi \mid l m\rangle$ in this space may be multiplied by any basis $f_{n}(r)$ of $\mathcal{L}\left((0, \infty)_{r}\right)$, to get basis $f_{n}(r) Y_{l}^{m}(\theta, \varphi)=\langle r, \theta, \varphi \mid n l m\rangle$ in the total state space.

In the view of (4.14) functions $Y_{l}^{m}(\theta, \varphi)$ are solution of the system of partial differential equations:

$$
\begin{gather*}
-\mathrm{i} \frac{\partial}{\partial \varphi} Y_{l}^{m}(\theta, \varphi)=m Y_{l}^{m}(\theta, \varphi),  \tag{4.23a}\\
-\left[\frac{1}{\sin ^{2} \theta} \frac{\partial^{2}}{\partial \varphi^{2}}+\frac{1}{\sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial}{\partial \theta}\right)\right] Y_{l}^{m}(\theta, \varphi)=l(l+1) Y_{l}^{m}(\theta, \varphi) . \tag{4.23b}
\end{gather*}
$$

Since $l_{z}$ effectively acts in the factor space $\mathcal{L}\left((0,2 \pi)_{\varphi}\right)$, one again applies theorem on the separation of variables to solve it. Obviously, the first of these equations is solved by the
functions $g_{l}^{m}(\theta) \mathrm{e}^{\mathrm{i} m \varphi}$, with $m$ restricted to the integer values by the last condition in (4.21). In the second eigenproblem we again use separation of variables: $l^{2}=A_{\theta} \otimes l_{z}+B_{\theta} \otimes \mathbf{1}_{\varphi}$; the factors in $\mathcal{L}\left((0,2 \pi)_{\varphi}\right)$ commute, the second one being identity, while the first one is function of $l_{z}$. Thus, the eigenfunctions have form $g_{l}^{m}(\theta) \mathrm{e}^{\mathrm{i} m \varphi}$; substituting this, the second eigenproblem becomes equation in $g_{l}^{m}(\theta)$ :

$$
\begin{equation*}
\left[\frac{m^{2}}{\sin ^{2} \theta}-\frac{1}{\sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial}{\partial \theta}\right)-l(l+1)\right] g_{l}^{m}(\theta)=0 \tag{4.24}
\end{equation*}
$$

After transition to new variable $\chi=\cos \theta$ (consequently $\frac{\partial}{\partial \chi}=-\frac{1}{\sin \theta} \frac{\partial}{\partial \theta}$ ), this becomes general hypergeometric equation:

$$
\begin{equation*}
\left[\frac{\partial^{2}}{\partial \chi^{2}}-\frac{2 \chi}{1-\chi^{2}} \frac{\partial}{\partial \chi}+\frac{l(l+1)\left(1-\chi^{2}\right)-m^{2}}{\left(1-\chi^{2}\right)^{2}}\right] g_{l}^{m}(\chi)=0 . \tag{4.25}
\end{equation*}
$$

The first condition in (4.22) implies that for $m \neq 0$ only functions such that $g_{l}^{m}(\chi= \pm 1)=$ 0 should be considered. Straightforward application of the general method (Appendix A.2) shows that the for each $l=0,1, \ldots$ and $m=0, \pm 1, \ldots, \pm l$ there is a single solution of (4.25) in the form of the Legendre polynomial $g_{l}^{m}(\chi)=\left(1-\chi^{2}\right)^{\frac{m}{2}} P_{l-|m|}^{(|m|,|m|)}(\chi)$. Finally, including the factors providing normalization

$$
\begin{equation*}
\int_{0}^{2 \pi} \mathrm{~d} \varphi \int_{0}^{\pi} \sin \theta \mathrm{d} \theta Y_{l}^{m^{*}}(\theta, \varphi) Y_{l^{\prime}}^{m^{\prime}}(\theta, \varphi)=\delta_{l l^{\prime}} \delta_{m m^{\prime}} \tag{4.26}
\end{equation*}
$$

one gets the spherical harmonics as the standard basis:

$$
\begin{equation*}
Y_{l}^{m}(\theta, \varphi)=\frac{(-)^{l-m}}{2^{l} l!\sqrt{2 \pi}} \sqrt{\frac{2 l+1}{2} \frac{(l+m)!}{(l-m)!}} \frac{1}{\sin ^{|m|} \theta} \mathrm{e}^{\mathrm{i} m \varphi} \frac{\mathrm{~d}^{l-|m|}}{\mathrm{d}(\cos \theta)^{l-|m|}} \sin ^{2 l} \theta . \tag{4.27}
\end{equation*}
$$

Exercise 4.4: Using the method explained in Appendix A. 2 prove that (4.27) are solutions of (4.23).

### 4.4 Central potentials

Quantum numbers $l$ and $m$ are conserved whenever orbital angular momentum commutes with hamiltonian. In such cases $l^{2}$ and $l_{z}$ may be used together with hamiltonian (and if necessary some other observables) to form a complete set of compatible observables. This way not only solving of the Schrödinger equation is simplified, but also its solution are clearly interpreted. Since the kinetic part of the hamiltonian is reduced to the square of momentum, it always commutes with angular momentum components. Hence, compatibility of $H$ and angular momentum is determined by the potential. The condition $\left[V(\boldsymbol{r}), l_{i}\right]=0$ immediately means that the potential commutes with all rotations: $\left[V, D\left(R_{\phi}\right)\right]=0$, i.e. $D^{-1}\left(R_{\phi}\right) V D\left(R_{\phi}\right)=V$. Assuming that the potential depends only on the coordinates, the last condition reads that the potential must have spherical symmetry:

$$
\begin{equation*}
D^{-1}\left(R_{\phi}\right) V(\boldsymbol{r}) D\left(R_{\phi}\right)=V\left(R_{\phi} \boldsymbol{r}\right) . \tag{4.28}
\end{equation*}
$$

This analysis emphasizes that in the problems with spherical symmetry only, the orbital angular momentum has particularly important role, with its components being conserved quantities. Technically this fact is efficiently exploited if $l^{2}$ and $l_{z}$ are included in CSCO together with $H$.

Obviously, in the spherical coordinates (4.28) shows that the potential depends only on the distance $r$ from the coordinate origin. Indeed, acting on some chosen point $\boldsymbol{r}=$ $(r, \theta, \varphi)$, the set of all rotations generates sphere with the radius $r$, and along this sphere $V(\boldsymbol{r})$ is constant, meaning that $V$ depends on $r$ only, i.e. (4.28) is satisfied if and only if the potential is central $V(\boldsymbol{r})=V(r)$.

According to the results of Section 4.3, in such problems the time independent Schrödinger equation is solved by the wave functions of the form $R_{n l}(r) Y_{l}^{m}(\theta, \varphi)$. In fact, as in the angular factor space (4.22) of the total space (4.21) $l^{2}$ and $l_{z}$ are CSCO, the spherical harmonics are necessarily angular part of any CSCO in the total space. Therefore, only the radial functions $R_{n l}(r)$ are to be determined for any specific potential. Really, since the hamiltonian of the particle in the central potential is:

$$
\begin{equation*}
H=\frac{p^{2}}{2 m}+V(r)=-\frac{\hbar}{2 m} \triangle+V(r) \tag{4.29}
\end{equation*}
$$

Further, according to (4.20) the laplacian has the form $\triangle=\frac{1}{r^{2}} \frac{\partial}{\partial r} r^{2} \frac{\partial}{\partial r}-\frac{1}{\hbar^{2} r^{2}} l^{2}$. Hence, using the theorem on the separation of variables, the time independent Schrödinger equation reduces to the radial part only:

$$
\begin{equation*}
\left[\frac{1}{r^{2}} \frac{\mathrm{~d}}{\mathrm{~d} r} r^{2} \frac{\mathrm{~d}}{\mathrm{~d} r}+\left(\frac{2 m}{\hbar^{2}}\left(E_{n l}-V(r)\right)-\frac{l(l+1)}{r^{2}}\right)\right] R_{n l}(r)=0 . \tag{4.30}
\end{equation*}
$$

Recall that the radial eigenbasis is to be found solving such an equation for each $l=$ $0,1, \ldots$ Particularly, as only for $l=0$ spherical harmonic is constant, for $l>0$ the first condition in (4.21) implies that $R_{n l}(r=0)=0$. The finite norm of the solutions means that $\int_{0}^{\infty} r^{2} \mathrm{~d} r\left|R_{n l}(r)\right|^{2}<\infty$.

In some problems it is convenient to use $\tilde{R}_{n l}(r)=r R_{n l}(r)$, simplifying (4.30) to

$$
\begin{equation*}
\left[\frac{\mathrm{d}^{2}}{\mathrm{~d} r^{2}}+\frac{2 m}{\hbar^{2}}\left(E_{n l}-V(r)\right)-\frac{l(l+1)}{r^{2}}\right] \tilde{R}_{n l}(r)=0 \tag{4.31}
\end{equation*}
$$

Again, for $l>0$ the first condition in (4.21) implies that for $r \rightarrow 0$ the limiting value of the function is $\tilde{R}_{n l}(r) / r \rightarrow 0$; the normalization is enabled by $\int_{0}^{\infty} \mathrm{d} r\left|\tilde{R}_{n l}(r)\right|^{2}<\infty$.

### 4.4.1 Free particle

It has been shown that the plane waves are solutions of the time independent Schrödinger equation of the free particle, $V(r)=0$. These are found taking momentum components as the CSCO compatible with hamiltonian. Recall that the eigenenergies $E_{k}=E_{k}=$ $\hbar^{2} k^{2} / 2 m$ are degenerate along the spheres of the radius $k$.

However, the constant zero potential possesses spherical symmetry, and therefore CSCO can be formed by $H, l^{2}$ and $l_{z}$. Accordingly, there is alternative eigenbasis of
$H$, subdued to this choice of CSCO. The radial equation (4.31) becomes

$$
\left[\frac{\mathrm{d}^{2}}{\mathrm{~d} r^{2}}+\frac{2 m}{\hbar^{2}} E-\frac{l(l+1)}{r^{2}}\right] \tilde{R}_{n l}(r)=0
$$

Denoting $\frac{2 m}{\hbar^{2}} E$ by $k^{2}$ (in accordance with known solution) and introducing variable $x=k r$, one gets

$$
\begin{equation*}
\left[\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+1-\frac{l(l+1)}{x^{2}}\right] \tilde{R}_{k l}(x)=0 \tag{4.32}
\end{equation*}
$$

Note that instead of $n$, the energies are counted by $k \geq 0$. However, as the equation itself is $k$ independent, the solvability gives no restriction on $k$, the spectrum is continual (as expected), and for each $k$ the same set of solutions (the functions of $x$, not of $r!$ ) must be obtained. The equation can be solved iteratively. In fact, for $l=0$, it reduces to $\left(\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+1\right) \tilde{R}_{k 0}(x)=0$, solved by $\tilde{R}_{k 0}(x)=2 \sin x$. On the other hand, the solution can be looked for in the form $\tilde{R}_{k l}(x)=x^{l+1} \chi_{k l}(x)$, when (4.32) becomes equation in $\chi$.

$$
\begin{equation*}
\left[\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+\frac{2(l+1)}{x} \frac{\mathrm{~d}}{\mathrm{~d} x}+1\right] \chi_{k l}(x)=0 . \tag{4.33}
\end{equation*}
$$

To solve it, one differentiate the whole equation, getting the equation in $\chi_{k l}^{\prime}(x)$. When auxiliary function $\phi_{k l}(x)$ is introduced by $\chi_{k l}^{\prime}(x)=x \phi_{k l}(x)$, it turns out that the obtained equation in $\phi_{k l}(x)$

$$
\left[\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+\frac{2(l+2)}{x} \frac{\mathrm{~d}}{\mathrm{~d} x}+1\right] \phi_{k l}(x)=0
$$

is same as (4.33), but for $l+1$. Therefore, $\chi_{k, l+1}(x)=\frac{1}{x} \frac{\mathrm{~d}}{\mathrm{~d} x} \chi_{k l}(x)$, meaning that $\chi_{k, l}(x)=$ $\left(\frac{1}{x} \frac{\mathrm{~d}}{\mathrm{~d} x}\right)^{l} \chi_{k 0}(x)$. This conclusion, together with the found result for $\tilde{R}_{k 0}(x)$ straightforwardly gives: $\tilde{R}_{k l}(x)=x^{l+1}\left(\frac{1}{x} \frac{\mathrm{~d}}{\mathrm{~d} x}\right)^{l} 2 \frac{\sin x}{x}$. Finally, the solutions of the radial part are:

$$
\begin{equation*}
R_{k l}(r)=2(-)^{l} \frac{r^{l}}{k^{l}}\left(\frac{1}{r} \frac{\mathrm{~d}}{\mathrm{~d} r}\right)^{l} 2 \frac{\sin k r}{r}=2 k j_{l}(k r) \tag{4.34}
\end{equation*}
$$

Here, $j_{l}(k r)$ are Bessel's functions of the first kind (integer, or spherical Bessel's functions). Altogether, the generalized eigenbasis for CSCO with angular momentum is

$$
\begin{equation*}
\psi_{k l m}(\boldsymbol{r})=\langle r \theta \varphi \mid k l m\rangle=2 k j_{l}(k r) Y_{l}^{m}(\theta, \varphi) . \tag{4.35}
\end{equation*}
$$

### 4.4.2 Coulomb potential and Hydrogen like atoms

Hamiltonian of the relative particle of the hydrogen like atom with $Z$ protons in the nucleus is

$$
\begin{equation*}
H=\frac{p^{2}}{2 m}-\frac{Z e^{2}}{r} \tag{4.36}
\end{equation*}
$$

where $m=m_{\mathrm{e}} m_{\mathrm{N}} /\left(m_{\mathrm{e}}+m_{\mathrm{N}}\right)$ and $r=\left|\boldsymbol{R}_{\mathrm{N}}-\boldsymbol{r}_{\mathrm{e}}\right|$; note that $m \approx m_{\mathrm{e}}$ as $m_{\mathrm{e}} \ll m_{\mathrm{N}}$. The radial Schrödinger equation becomes:

$$
\begin{equation*}
\left[\frac{\mathrm{d}^{2}}{\mathrm{~d} r^{2}}+\frac{2 m}{\hbar^{2}}\left(E_{n l}+\frac{Z e^{2}}{r}\right)-\frac{l(l+1)}{r^{2}}\right] \tilde{R}_{n l}(r)=0 . \tag{4.37}
\end{equation*}
$$

This equation takes more standard form if the involved quantities are reparameterized. Using Bohr radius $a_{0}=\hbar^{2} / m_{\mathrm{e}} e^{2}=0.529 \AA$, and atomic unit of energy (one Henry equal to two Rydbergs) $e^{2} / a_{0}=4.3610^{-18} \mathrm{~J}=27.21 \mathrm{eV}$, the energy and radius are substituted by the dimensionless $\varepsilon_{n l}=a_{0} E_{n l} / e^{2}$ and $x=r / a_{0}$, to get the generalized hypergeometric equation:

$$
\begin{equation*}
\left[\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+\frac{2 \varepsilon_{n l} x^{2}+2 Z x-l(l+1)}{x^{2}}\right] \tilde{R}_{n l}(x)=0 . \tag{4.38}
\end{equation*}
$$

The standard method of solving of such equation (Appendix A.2) straightforwardly gives that the energy spectrum and radial functions are for each $n=1,2, \ldots$ and $l=0, \ldots, n-1$ :

$$
\begin{equation*}
E_{n}=-\frac{m e^{4}}{2 \hbar^{2}} \frac{Z}{n^{2}}, \quad R_{n l}(r)=C_{n l} \mathrm{e}^{-\frac{2 Z}{n} \frac{m e^{2}}{\hbar^{2}} r}\left(\frac{m e^{2}}{\hbar^{2}} r\right)^{l+1} L_{n-l-1}^{2 l+1}\left(\frac{2 Z}{n} \frac{m e^{2}}{\hbar^{2}} r\right), \tag{4.39}
\end{equation*}
$$

with normalizing factor $C_{n l}=\sqrt{\frac{Z(n-l-1)!}{n!(n+l)!}}$, and $L_{q}^{p}(x)$ being generalized Laguerre polynomials. Note that the energies are degenerate in $l$, due to the hidden symmetry of another group isomorphic to the rotational one (but not related to the spatial rotations; this is characteristic of the Coulomb potential). This enlarges degeneracy of energy levels to $\sum_{l=0}^{n-1}(2 l+1)=n^{2}$.

### 4.5 Spin

There are several experiments showing that the orbital space $\mathcal{S}_{\text {o }}$ which we are dealing with in the previous sections is not sufficient to describe completely the single particle states.

### 4.5.1 Zeeman's Effect

Hamiltonian of the particle with the mass $m$ and charge $q$ in the electromagnetic field defined by the scalar and vector potential $\Phi$ and $\boldsymbol{A}$ is

$$
\begin{equation*}
H=\frac{1}{2 m}\left(\boldsymbol{p}-\frac{q}{c} \boldsymbol{A}\right)^{2}+q \Phi \tag{4.40}
\end{equation*}
$$

( $c$ is the speed of light). Particularly, the potentials of the homogeneous constant magnetic field $\boldsymbol{B}=B \boldsymbol{e}_{z}$ can be in the Coulomb gauge $(\operatorname{div} \boldsymbol{A}=0)$ taken in the form $\boldsymbol{A}=\frac{1}{2} \boldsymbol{B} \times \boldsymbol{r}$ and $\Phi=0$, giving
$\left(\boldsymbol{p}-\frac{q}{c} \boldsymbol{A}\right)^{2}=p^{2}-\frac{q}{2 c}[\boldsymbol{p} \cdot(\boldsymbol{B} \times \boldsymbol{r})+(\boldsymbol{B} \times \boldsymbol{r}) \cdot \boldsymbol{p}]+\frac{q^{2}}{4 c^{2}}(\boldsymbol{B} \times \boldsymbol{r})^{2}=p^{2}-\frac{q}{c} B l_{z}+\frac{q^{2}}{4 c^{2}} B^{2} r_{\perp}^{2}$.
Here, $r_{\perp}$ is the radius vector component perpendicular to the field.
Consequently, when a hydrogen like atom is exposed to such a magnetic field, the hamiltonian of the electron becomes:

$$
\begin{equation*}
H=-\frac{\hbar^{2}}{2 m} \Delta-\frac{Z e^{2}}{r}+\frac{\mu_{\mathrm{B}}}{\hbar} B l_{z}+\frac{e^{2} B^{2}}{8 m c^{2}} r_{\perp}^{2}, \tag{4.41}
\end{equation*}
$$

where $\mu_{\mathrm{B}}=\frac{e \hbar}{2 m c}=0.92710^{-23} \mathrm{~J} / \mathrm{T}$ is the Bohr's magneton. The terms involving magnetic field can be roughly estimated by the mean values for the lowest hydrogen like atom states: $\left\langle r_{\perp}^{2}\right\rangle \sim a_{0}^{2},\left\langle l_{z}\right\rangle \sim \hbar$, one finds that $\left\langle\mu_{\mathrm{B}} B l_{z}\right\rangle \sim 0.92710^{-23} B \mathrm{~J} / \mathrm{T},\left\langle\frac{e^{2} B^{2}}{8 m c^{2}} r_{\perp}^{2}\right\rangle \sim$ $410^{-26} B^{2} \mathrm{~J} / \mathrm{T}^{2}$. The maximal experimental magnetic fields are of the order 1 T , making therefore the last term by thousands less then the third term. In addition, assuming that the typical energies of the lowest hydrogen like atom states are of the order of $1 \mathrm{eV}=1.610^{-19} \mathrm{~J}$, the third term is by thousands less than the energy of the isolated atom. Therefore the last term may be neglected, while the rest of the hamiltonian may be treated perturbatively, using the isolated atom hamiltonian as the nonperturbed one, while the third therm is perturbation. Accordingly, using the expression (6.5) for the first correction of the degenerate level, and noticing that the hydrogen like atom eigenstates $|n l m\rangle$ are also eigenstates for the perturbation, one finds the perturbed energies

$$
\begin{equation*}
E_{n m}=E_{n}+\mu_{\mathrm{B}} B m \tag{4.42}
\end{equation*}
$$

Thus, quite generally, the energy level is split in the $2 l+1$ equidistant sublevels. This is called Zeeman's effect. Particularly, for the hydrogen like atoms, with energy levels $E_{n}$ being degenerate in angular momentum values $l=0, \ldots n-1$, this result implies that each level is in the magnetic field split into $2 l_{\max }+1$ equidistant sublevels $E_{n m}$ for $m=-l, \ldots, l$ and $l_{\text {max }}=n-1$.

However, in the experiments with hydrogen like atoms, it appears that there are even number of sublevels, and this is called anomalous Zeeman's effect. Obviously, it cannot be described within introduced formalism.

### 4.5.2 Interior degrees of freedom

Further evidence on the incompleteness of the formalism was given by the Stern-Gerlach experiment: passing through the inhomogeneous magnetic field, well collimated monochromatic beam of atoms with a single electron is split into parts. Analysis analogous to that of the Zeeman's effect shows again that the number of obtained beams should be odd, e.g. no splitting can occur for the single $s$-state electron. But in such cases even number of split beams is found experimentally.

Still, there is a way to explain the described effects within quantum mechanical formalism. The Postulate 4 emphasizes that the total state space is the direct product of the state spaces for each degree of freedom. In the view of this, the results of the previous experiments shows that there is another degree of freedom, being not of the orbital type: the possible values are not parameterized by the continual variable (e.g. coordinate or momentum), but take a finite number of values. Such degrees are called interior degrees of freedom.

## POSTULATE 6. - INTERIOR DEGREES

Besides orbital, there are also interior degrees of freedom, contributing to the total state space $\mathcal{S}$ by the finite dimensional interior factor space:

$$
\mathcal{S}=\mathcal{S}_{\text {orbital }} \otimes \mathcal{S}_{\text {interior }} .
$$

The construction of the total space as the product of the orbital and interior ones is in fact required by the previous Postulate 4. Of course, the same argument shows that if there are several interior degrees of freedom, the interior space is constructed as the direct product of the corresponding interior spaces: $\mathcal{S}_{\text {interior }}=\mathcal{S}_{\text {interior }}^{1} \otimes \mathcal{S}_{\text {interior }}^{2} \otimes \ldots$ Only the finiteness of the interior spaces is different with respect to the previous spaces based on the particle position coordinates.

Finite dimension of the space of function over some set $X$ means that the set itself is finite. Indeed, as any function over finite (meaning also discrete) set is continuous, such a space is spanned by the functions $f_{x}$ being zero everywhere but in a single $x$ from $X$, e.g. $f_{x}(y) \stackrel{\text { def }}{=} \delta_{x y}$ (Kronecker step function). Obviously, the dimension of the spanned space $\mathcal{L}^{2}(X)$ is equal to the number of points in $X$. Consequently, the last postulate means that the interior degrees of freedom are finite sets.

This fact may be considered as an empirical one, although relativistic theory gives a profound proof for it through the analysis of the possible extensions of the Poincare group: Coleman-Mandula theorem asserts that any group having the Poincare group as a subgroup, and describing the space-time symmetries in the physically correct way, must be a direct product of the Poincare group and some compact groups. Therefore, its irreducible representations are products of the representations of the Poincare group with the finite-dimensional representations of the compact factor. So, while the Poincare group leads to infinite dimensional orbital space (its unitary irreducible representations are infinite dimensional), the compact factors introduce finite-dimensional interior spaces.

### 4.5.3 Formalism of the spin $s=\frac{1}{2}$

The simplest nontrivial case of the angular momentum is the representation for $k=1 / 2$. According to the Theorem 4.1, the dimension of the space $\mathcal{S}^{\left(\frac{1}{2}\right)}$ of this representation is $\left|\mathcal{S}^{\left(\frac{1}{2}\right)}\right|=2 k+1=2$. A suitable hermitian basis of the corresponding operator space (four dimensional) is formed by $\sigma_{0}=\left(\begin{array}{ll}1 & 0 \\ 0 & 1\end{array}\right)$ and three Pauli matrices:

$$
\sigma_{1}=\left(\begin{array}{ll}
0 & 1  \tag{4.43}\\
1 & 0
\end{array}\right), \quad \sigma_{2}=\left(\begin{array}{cc}
0 & -\mathrm{i} \\
\mathrm{i} & 0
\end{array}\right), \quad \sigma_{3}=\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right) .
$$

For the Pauli matrices the following properties are easily verified:

$$
\begin{equation*}
\sigma_{i}^{2}=I, \quad \operatorname{Tr} \sigma_{i}=0, \quad \operatorname{Det} \sigma_{i}=-1, \quad \sigma_{i} \sigma_{j}+\sigma_{j} \sigma_{i}=0, \quad \sigma_{1} \sigma_{2} \sigma_{3}=\mathrm{i} \sigma_{0} \tag{4.44}
\end{equation*}
$$

It follows that the basis $\sigma_{i}(i=0, \ldots, 3)$ is orthogonal with respect to the operators scalar product: $\operatorname{Tr} \sigma_{i} \sigma_{j}=2 \delta_{i j}$. Therefore, since each operator in $\mathcal{S}^{\left(\frac{1}{2}\right)}$ can be expanded in the form $A=\sum_{i=0}^{4} a_{i} \sigma_{i}$, with the complex coefficients $a_{i}=\frac{1}{2} \operatorname{Tr} \sigma_{i} A$. The last three coefficients can be grouped into the vector $\boldsymbol{a}$, to get $A=a_{0} \sigma_{0}+\boldsymbol{a} \cdot \boldsymbol{\sigma}$ (here, like before, we use vector notation for the triple of Pauli matrices). The second part $\boldsymbol{a} \cdot \boldsymbol{\sigma}$ gives the traceless part of $A$, while the first one gives the scalar part with trace $2 a_{0}$. For the hermitian operators, the coefficients $a_{i}$ are real.

Pauli matrices are used to define the spin matrices by

$$
\begin{equation*}
S_{i}=\frac{1}{2} \hbar \sigma_{i} \quad(i=1,2,3) . \tag{4.45}
\end{equation*}
$$

It is easy to verify that

$$
\begin{equation*}
\left[S_{i}, S_{j}\right]=\mathrm{i} \hbar \sum_{k} \epsilon_{i j k} S_{k} \tag{4.46}
\end{equation*}
$$

showing that these operators represent the angular momentum components in $\mathcal{S}^{\left(\frac{1}{2}\right)}$. Accordingly, the spin square eigenvalue is $s(s+1) \hbar^{2}=\frac{3}{4} \hbar^{2}$, which is also follows from $S^{2}=\sum_{i} S_{i}^{2}=\frac{3}{4} \hbar^{2} \sigma_{0}$.

Using the general properties of angular momentum (or directly solving the eigenproblem of $S_{3}$ ), one finds that the standard basis in $\mathcal{S}^{\left(\frac{1}{2}\right)}$ is $\left\{|+\rangle=\left|m=\frac{1}{2}\right\rangle,|-\rangle=\left|m=-\frac{1}{2}\right\rangle\right\}$, for the eigenvalues $\pm \frac{1}{2} \hbar$ of $S_{3}$.

The rotations in the spin space are represented by the operators $D^{\left(\frac{1}{2}\right)}\left(R_{\phi \boldsymbol{u}}\right)=\mathrm{e}^{-\frac{1}{2} \phi \boldsymbol{u} \cdot \boldsymbol{\sigma}}$. Particularly, in the standard basis to the rotation for $\phi$ around $z$ axis corresponds the matrix $\left(\begin{array}{cc}\exp \left(-\frac{i}{2} \phi\right) & 0 \\ 0 & \exp \left(\frac{i}{2} \phi\right)\end{array}\right)$. Obviously, for $\phi=2 \pi$ the matrix $-I_{2}$ is obtained, while the identity transformation is achieved only for $\phi=4 \pi$. This shows that matrices for spin $\frac{1}{2}$ represent group $\mathrm{SU}(2)$.

### 4.6 Addition of angular momenta

Let us consider state space being product

$$
\begin{equation*}
\mathcal{S}=\mathcal{S}_{1} \otimes \mathcal{S}_{2} \tag{4.47}
\end{equation*}
$$

with angular momenta $\boldsymbol{K}_{1}$ and $\boldsymbol{K}_{2}$ in the factors spaces. Such situation is very frequent: e.g. orbital and spin momenta of a single electron, or two particles with angular momenta of each of them. From the definition $\left[K_{p i}, K_{p j}\right]=\mathrm{i} \hbar \sum_{k} K_{p k}$, it is easily found that the operators

$$
\begin{equation*}
K_{i}=K_{1 i} \otimes I_{2}+I_{1} \otimes K_{2 i}, \quad i=1,2,3, \tag{4.48a}
\end{equation*}
$$

defined in the whole space $\mathcal{S}$, satisfy the same commutation relations; in other words they also form angular momentum $\boldsymbol{K}$ being sum of the momenta $\boldsymbol{K}_{1}$ and $\boldsymbol{K}_{2}$ :

$$
\begin{equation*}
\boldsymbol{K}=\boldsymbol{K}_{1}+\boldsymbol{K}_{2} . \tag{4.48b}
\end{equation*}
$$

### 4.6.1 Irreducible subspaces and standard basis

For various purposes it is useful to know irreducible components and standard basis of the composed momentum $\boldsymbol{K}$. It is assumed that irreducible components of the momenta $\boldsymbol{K}_{1}$ and $\boldsymbol{K}_{2}$ are known, i.e. that the decompositions

$$
\begin{equation*}
\mathcal{S}_{i}=\oplus_{k_{i}, \lambda_{i}} \mathcal{S}_{i}^{\left(k_{i} \lambda_{i}\right)}, \quad i=1,2 \tag{4.49}
\end{equation*}
$$

are given. Here, $\boldsymbol{S}_{i}^{\left(k_{i} \lambda_{i}\right)}$ are irreducible subspaces of the momentum $\boldsymbol{K}_{i}$, where $k_{i}\left(k_{i}+1\right) \hbar$ is eigenvalue of $\boldsymbol{K}_{i}^{2}$ and $\lambda_{i}$ counts different irreducible subspaces of the same type when frequency number of $k_{i}$ is greater than 1 .

## Irreducible components and Clebsch-Gordan series

The product structure (4.47) of $\mathcal{S}$, together with decomposition (4.49) gives:

$$
\begin{equation*}
\mathcal{S}=\oplus_{k_{1}, k_{2}} \oplus_{\lambda_{1}, \lambda_{2}} \mathcal{S}_{1}^{\left(k_{1} \lambda_{1}\right)} \otimes \mathcal{S}_{2}^{\left(k_{2} \lambda_{2}\right)} \tag{4.50}
\end{equation*}
$$

Therefore, to reveal the irreducible components of angular momenta in $\mathcal{S}$ it is sufficient to find out the irreducible components in the product $\mathcal{V}^{\left(k_{1}\right)} \otimes \mathcal{V}^{\left(k_{2}\right)}$ of irreducible subspaces, i.e. Clebs-Gordan series.

Theorem 4.2 (Clebsch-Gordan series of angular momentum) The product of the irreducible spaces of the weight $k_{1}$ and $k_{2}$ of angular momentum contains exactly once each irreducible subspace with weight $k=\left|k_{1}-k_{2}\right|, \ldots, k_{1}+k_{2}$ :

$$
\begin{equation*}
\mathcal{V}^{\left(k_{1}\right)} \otimes \mathcal{V}^{\left(k_{2}\right)}=\oplus_{k=\left|k_{1}-k_{2}\right|}^{k_{1}+k_{2}} \mathcal{V}^{(k)} \tag{4.51}
\end{equation*}
$$

-Proof: This is one of the theorems easily derived as a special case of the Weyl's theorem on the semisimple Lie algebras. Here we recall that in each irreducible space $\left|k_{i} m_{i}\right\rangle\left(m_{i}=-k_{i}, \ldots, k_{i}\right)$ is standard basis. One basis of the direct product space is $\left|k_{1} m_{1} ; k_{2} m_{2}\right\rangle=\left|k_{1} m_{1}\right\rangle \otimes\left|k_{2} m_{2}\right\rangle$, and this basis is represented by the points ( $m_{1}, m_{2}$ ) in Fig. 4.2. Vectors of this basis are eigenvectors of $K_{3}$ :

$$
K_{3}\left|k_{1} m_{1} ; k_{2} m_{2}\right\rangle=\left(m_{1}+m_{2}\right) \hbar\left|k_{1} m_{1} ; k_{2} m_{2}\right\rangle .
$$

Therefore, vectors with the same sum $m=m_{1}+m_{2}$ span eigensubspaces of $K_{3}$. These subspaces are denoted by the inclined lines connected vectors with the same $m$; below each line the corresponding $m_{1}+m_{2}$ and number $d_{m}$ of vectors on it are written. Obviously, the maximal sum is $k_{1}+k_{2}$ with a single vector $\left|k_{1} k_{1} ; k_{2} k_{2}\right\rangle$ on it. The sum takes on all the values $k_{1}+k_{2}, k_{1}+k_{2}-1, \ldots,-\left(k_{1}+k_{2}\right)$, and the degeneracy of these values increases by one for $m=k_{1}+k_{2}, k_{1}+k_{2}-1, \ldots, k_{1}-k_{2}$ (we assumed that $k_{1} \geq k_{2}$ ), then remains constant for $m=k_{1}-k_{2}, k_{1}-k_{2}-1, \ldots, k_{2}-k_{1}$, and finally decreases by one for $m=k_{2}-k_{1}, k_{2}-k_{1}-1, \ldots,-k_{1}-k_{2}$. Thus, in the decomposition appears single irreducible subspace corresponding to the maximal value $k_{1}+k_{2}$; it must contain a single eigenvector of $K_{3}$ for all eigenvalues $k_{1}+k_{2}, k_{1}+k_{2}-1, \ldots,-\left(k_{1}+k_{2}\right)$. When this representation is taken into account, all degeneracies should be diminished by 1 . Therefore, among the remaining $m$ values, the maximal $m_{1}+m_{2}-1$ is now nondegenerate. We repeat procedure, collecting all less eigenvalues corresponding to this representation, and subtracting from the diagram. So, all the representations from $k_{1}+k_{2}$ to $k_{1}-k_{2}$ appear once, and after this the diagram is exhausted (as the degeneracies are constant).

This way the total space reduction (4.50) becomes

$$
\begin{equation*}
\mathcal{S}=\oplus_{k_{1}, k_{2}} \oplus_{k=\left|k_{1}-k_{2}\right|}^{k_{1}+k_{2}} \oplus_{\lambda_{1}, \lambda_{2}} \mathcal{V}_{\lambda_{1} \lambda_{2}}^{(k)} . \tag{4.52}
\end{equation*}
$$

## Standard basis and Clebsch-Gordan coefficients

After the Clebsch-Gordan series are found, we can proceed to determine the standard basis in the product space. According to the last decomposition, one can use standard basis of the direct product of the irreducible spaces. Namely, if $\left|k_{i}, m_{i}\right\rangle(i=1,2)$ aree standard bases in the irreducible spaces $\mathcal{V}^{\left(k_{1}\right)}$ and $\mathcal{V}^{\left(k_{2}\right)}$, then in the product space, since each irreducible component occurs once, the standard basis $\left|k_{1} k_{2} ; k m\right\rangle$ is unique (up to the common phase factor). Therefore the non-correlated $\left|k_{1}, m_{1} ; k_{2}, m_{2}\right\rangle=\left|k_{1} m_{1}\right\rangle \otimes\left|k_{2} m_{2}\right\rangle$


Figure 4.2: Decomposition of the product of irreducible subspaces of angular momentum. Case $k_{1}=3$ and $k_{2}=3 / 2$ is illustrated. At the end of each inclined line the sum $m_{1}+m_{2}$ for the point ( $m_{1}, m_{2}$ ) on it is depicted (concrete example is given above in gray) and the number of these points (degeneracy of the sum value) is below. Values of $m_{1}+m_{2}$ included in the irreducible subspace with fixed maximal weight (written on the right side of diagram) are the lines crossed by the shaded region corresponding to maximal weight.
and standard basis are related by the unitary transition matrix, with elements being Clebsch-Gordan coefficients $\left\langle k_{1}, m_{1} ; k_{2}, m_{2} \mid k_{1} k_{2} ; k m\right\rangle$ :

$$
\begin{equation*}
\left|k_{1} k_{2} ; k m\right\rangle=\sum_{m_{1}=-k_{1}}^{k_{1}} \sum_{m_{2}=-k_{2}}^{k_{2}}\left\langle k_{1}, m_{1} ; k_{2}, m_{2} \mid k_{1} k_{2} ; k m\right\rangle\left|k_{1}, m_{1} ; k_{2}, m_{2}\right\rangle . \tag{4.53}
\end{equation*}
$$

The Clebsch-Gordan coefficients as completely mathematical (i.e. independent of physical context) constants are calculated and tabulated ${ }^{2}$ in various forms (including computer programs). Still, it is interesting to exploit the eigensubspaces of $K_{3}$ used in the proof of the Theorem 4.2. In the first step we found that the standard vector of the maximal weight is noncorrelated: $\left|k_{1} k_{2} ; k=k_{1}+k_{2}, m=k_{1}+k_{2}\right\rangle=\left|k 1, m_{1}=k_{1} ; k_{2}, m_{2}=k_{2}\right\rangle$. Then, using the operator $K_{-}=K_{1-} \otimes I_{2}+I_{1} \otimes K_{2-}$ we find (see Theorem 4.1):

$$
K_{-}\left|k_{1} k_{2} ; k_{1}+k_{2}, k_{1}+k_{2}\right\rangle=c_{m}^{(k)-}\left|k_{1} k_{2} k_{1}+k_{2}, k_{1}+k_{2}-1\right\rangle .
$$

This vector belongs to the eigenspace of $K_{3}$ for eigenvalue $k_{1}+k_{2}-1$. This eigenspace (Fig. 4.2, the second inclined line) is double degenerate. Thus, only one vector orthogonal to $\left|k_{1} k_{2} ; k_{1}+k_{2}, k_{1}+k_{2}\right\rangle$ can be found, and it must be $\left|k_{1} k_{2} ; k_{1}+k_{2}-1, k_{1}+k_{2}-1\right\rangle$. Then, in the triply degenerate subspace of the eigenvalue $m=k_{1}+k_{2}-2$ two basis vectors are proportional to $K_{-}^{2}\left|k_{1} k_{2} ; k_{1}+k_{2}, k_{1}+k_{2}\right\rangle$ and $K_{-}\left|k_{1} k_{2} ; k_{1}+k_{2}-1, k_{1}+k_{2}-1\right\rangle$, and the vector orthogonal to these two is (up to a phase) $\left|k_{1} k_{2} ; k_{1}+k_{2}-2, k_{1}+k_{2}-2\right\rangle$. Repeating procedure all the standard vectors and the corresponding Clebsch-Gordan coefficients are derived.

[^4]It is important to stress out that though the decomposition (4.50) involves noncorrelated bases differing by indices $\lambda_{i}$, the Clebsch-Gordan coefficient do not depend on them, giving that the total standard basis is:

$$
\begin{equation*}
\left|k_{1} k_{2} ; k m ; \lambda_{1} \lambda_{2}\right\rangle=\sum_{\lambda_{1}, \lambda_{2}} \sum_{m_{1}=-k_{1}}^{k_{1}} \sum_{m_{2}=-k_{2}}^{k_{2}}\left\langle k_{1}, m_{1} ; k_{2}, m_{2} \mid k_{1} k_{2} ; k m\right\rangle\left|k_{1}, m_{1}, \lambda_{1} ; k_{2}, m_{2}, \lambda_{2}\right\rangle \tag{4.54}
\end{equation*}
$$

Note that $\left\{K_{1}^{2}, K_{13}, A_{1}, K_{2}^{2}, K_{23}, A_{2}\right\}$ is CSCO with non-correlated basis as common eigenbases, while CSCO for the standard basis (4.54) is $\left\{K_{1}^{2}, A_{1}, K_{2}^{2}, A_{2}, K^{2}, K_{3}\right\}$.

## Wigner-Eckart theorem and selection rules

In the state space with defined angular momentum $\boldsymbol{K}$ there is also a representation $D(S U(2))$. Standard basis $|k m \lambda\rangle$, defined by equations (4.14) corresponds to the decomposition of the space to the irreducible subspaces of $\mathrm{SU}(2): \mathcal{S}=\oplus_{k \lambda} \nu_{\lambda}^{(k)}$, meaning that (4.14) are manifestation of the general requirement for standard basis $D(U)|k m \lambda\rangle=$ $\sum_{m^{\prime}=-k}^{k} D_{m^{\prime} m}^{(k)}(U)\left|k m^{\prime} \lambda\right\rangle$.

However, the transformations of $\mathrm{SU}(2)$ act also on the operators in $\mathcal{S}$ according to the rule: $\hat{U} A=D(U) A D^{\dagger}(U)$. This generates a representation in the operator space, and the standard basis among the operators can be defined:

$$
\begin{equation*}
\hat{\hat{U}} A_{m \lambda}^{(k)}=\sum_{m^{\prime}=-k}^{k} D_{m^{\prime} m}^{(k)}(U) A_{m^{\prime} \lambda}^{(k)} \tag{4.55}
\end{equation*}
$$

Usually, we are not interested in the whole basis, but only in the particular subset corresponding to an irreducible component

$$
\begin{equation*}
\mathrm{A}^{(k)}=\left\{A_{m}^{(k)} \mid m=-k, \ldots, k\right\} \tag{4.56}
\end{equation*}
$$

called irreducible tensor (operators). Therefore, index $\lambda$ is not used any more. Again, using Lie algebra, i.e. angular momentum, (4.55) becomes:

$$
\begin{equation*}
\left[K_{3}, A_{m}^{(k)}\right]=m \hbar A_{m}^{(k)}, \quad\left[K_{ \pm}, A_{m}^{(k)}\right]=\sqrt{(k \mp m)(k \pm m+1)} \hbar A_{m \pm 1}^{(k)} \tag{4.57}
\end{equation*}
$$

This way symmetry adapted bases are imposed both in the space of operators and in the state space. It is interesting to find the matrix elements of the irreducible tensors between standard basis elements.

Theorem 4.3 (Wigner-Eckart theorem) Matrix element of an irreducible tensor component $A_{m}^{(k)}$ between two standard vectors is factorized to Clebsch-Gordan coefficient (independent on $\lambda_{1}$ and $\lambda$ ) and the reduced matrix element independent on $m, m_{1}$ and $m_{2}$ :

$$
\begin{equation*}
\langle k m \lambda| A_{m_{2}}^{\left(k_{2}\right)}\left|k_{1} m_{1} \lambda_{1}\right\rangle=\left\langle k_{1} m_{1} ; k_{2} m_{2} \mid k_{1} k_{2} k m\right\rangle\langle k \lambda|\left|A^{\left(k_{2}\right)}\right|\left|k_{1} \lambda_{1}\right\rangle . \tag{4.58}
\end{equation*}
$$

Note that the first term, Clebsch-Gordan coefficient, is strictly symmetry based and $a$ priori known, while physical context is given only by the reduced matrix element, which is constant for the whole multiplets of $m, m_{1}$ and $m_{2}$. Therefore it can be calculated (or experimentally found) for one particular choice of $m$ 's, and then used in all of the $(2 k+1)\left(2 k_{1}+1\right)\left(2 k_{2}+1\right)$ cases with fixed $k, k_{1}$ and $k_{2}$.

This conclusion is very important, since the matrix elements are basic ingredient in the calculations of the transition probabilities, i.e. in study of any physical process. Due to the Wigner-Eckart theorem, one can immediately find selection rules, i.e. list of the pairs of states between which the transition in the studied process is not possible: whenever Clebsch-Gordan coefficient vanishes the corresponding transition is forbidden.

## Addition of more than two angular momenta

In the case that there are more than two angular momenta the procedure of their addition is the same as for two, only the question of the order of addition appears.

For the three angular momenta two possibilities are usually considered: $\boldsymbol{K}=\left(\boldsymbol{K}_{1}+\right.$ $\left.\boldsymbol{K}_{2}\right)+\boldsymbol{K}_{3}=\boldsymbol{K}_{1}+\left(\boldsymbol{K}_{2}+\boldsymbol{K}_{3}\right)$. The difference is not in the irreducible components, but in the corresponding standarrd base. In the first case the standard basis is common eigenbasis for CSCO $\left\{K_{1}^{2}, K_{2}^{2}, K_{3}^{2}, K_{12}^{2}, K_{123}^{2}, K_{123,3}, A_{1}, A_{2}, A_{3}\right\}$, while in the second case CSCO is $\left\{K_{1}^{2}, K_{2}^{2}, K_{3}^{2}, K_{23}^{2}, K_{123}^{2}, K_{123,3}, A_{1}, A_{2}, A_{3}\right\}$; here $\boldsymbol{K}_{i j}=\boldsymbol{K}_{i}+\boldsymbol{K}_{j}$.

Considering only irreducible subspaces, i.e. the product $\mathcal{V}^{\left(k_{1}\right)} \otimes \mathcal{V}^{\left(k_{2}\right)} \otimes \mathcal{V}^{\left(k_{3}\right)}$, this results in two standard bases $\left|\left(k_{1} k_{2}\right) k_{3} k_{12} k m\right\rangle$ and $\left|k_{1}\left(k_{2} k_{3}\right) k_{23} k m\right\rangle$. These bases are related by the unitary matrix of $6 j$-symbols.

Analogously, for addition of four angular momenta we define $9 j$-symbols, etc.

## Superselection rule of total momentum

Note that each particle has spin corresponding to an irreducible angular momentum, i.e. being $s=0, \frac{1}{2}, 1, \ldots$, and that orbital momentum is always integer. Therefore, according to the theorem on the addition of angular momenta, total angular momentum of a particle is either half integer or integer. Further, the same rule for addition gives that total momentum of many particles is again either integer or half integer. This is known as the:

Theorem 4.4 (Superselection rule of momentum) State combining half integer and integer angular momentum cannot be realized in the nature.

To distinguish between states with integer and half integer angular momentum it is useful to introduce superselection observable $R=R_{2 \pi \boldsymbol{u}}=\mathrm{e}^{-\frac{\mathrm{i}}{\hbar} \boldsymbol{u} \cdot \boldsymbol{K}}$. This observable acts as identity on the vectors corresponding to integer $k$, and reverses the vectors with half integer $k$. Thus, if there is a state space combing integer and half integer total angular momentum, the corresponding vectors would span the eigensubspaces of $R$ for the eigenvalues 1 and -1 . The superselection principle asserts that in each state space only one eigenvalue is realized.

In particular, the space of the operators (superspace) is always with integer angular momentum.

### 4.6.2 Examples and applications

Addition of two spins $s=1 / 2$
In this case the dimension of $\mathcal{S}$ is four. While noncorrelated basis is $\left|s_{1} m_{s 1} ; s_{2} m_{s 2}\right\rangle$ for $m_{s_{i}}= \pm 1 / 2$, or in short $|++\rangle=(1,0,0,0)^{T},|+-\rangle=(0,1,0,0)^{T},|-+\rangle=(0,0,1,0)^{T}$,
 with the standard basis $\left|\frac{1}{2} \frac{1}{2} 00\right\rangle,\left|\frac{1}{2} \frac{1}{2} 1,-1\right\rangle,\left|\frac{1}{2} \frac{1}{2} 10\right\rangle,\left|\frac{1}{2} \frac{1}{2} 11\right\rangle$. While the maximal weight vectors is $\left|\frac{1}{2} \frac{1}{2} 11\right\rangle=|++\rangle$, all other are found with help of $S_{-}=\hbar\left(\begin{array}{cccc}0 & 0 & 0 & 0 \\ 1 & 0 & 0 & 0 \\ 1 & 0 & 0 \\ 0 & 1 & 1 & 0\end{array}\right)$ :

$$
\left|\frac{1}{2} \frac{1}{2} 10\right\rangle=(|+-\rangle+|-+\rangle) / \sqrt{2}, \quad\left|\frac{1}{2} \frac{1}{2} 1-1\right\rangle=|--\rangle, \quad\left|\frac{1}{2} \frac{1}{2} 00\right\rangle=(|+-\rangle-|-+\rangle) / \sqrt{2} .
$$

## Spin harmonics of electron

It is known that orbital angular momentum is integer, and contains all integer irreducible components: $\mathcal{S}_{O}=\oplus_{l=0}^{\infty} \mathcal{V}^{(l)}$ Taking spin of electron into account, one gets addition in the form $\boldsymbol{j}=\boldsymbol{l}+\boldsymbol{s}$ and the total space $\mathcal{S}=\mathcal{S}_{O} \otimes \mathcal{S}_{s}=\oplus_{l=0}^{\infty} \mathcal{V}^{(l)} \otimes \mathcal{V}^{\left(\frac{1}{2}\right)}$ is decomposed according to (4.51) into irreducible subspaces:

$$
\begin{equation*}
\mathcal{S}=\mathcal{V}_{l=0}^{\left(j=\frac{1}{2}\right)}+\oplus_{l=1}^{\infty}\left(\mathcal{V}_{l}^{\left(j=l-\frac{1}{2}\right)} \oplus \mathcal{V}_{l}^{\left(j=l+\frac{1}{2}\right)}\right)=\oplus_{j=\frac{1}{2}}^{\infty}\left(\mathcal{V}_{l=j-\frac{1}{2}}^{(j)} \oplus \mathcal{V}_{l=j+\frac{1}{2}}^{(j)}\right) \tag{4.59}
\end{equation*}
$$

Hence, each half integer total angular momentum occurs twice, due to coupling of the


Figure 4.3: Total angular momentum of electron.
spin to the orbital momentum $l=j-\frac{1}{2}$ and for to the orbital momentum $l=j+\frac{1}{2}$. In each of the irreducible subspaces $\mathcal{V}_{l}^{(j)}$ the standard basis $\left|l \frac{1}{2} ; j m_{j}\right\rangle$ is with help of the Clebsch-Gordan coefficients expressed as the combination of the product basis:

$$
\left|l \frac{1}{2} ; j m_{j}\right\rangle=\sum_{m=-l}^{l} \sum_{m_{s}=-\frac{1}{2}}^{\frac{1}{2}}\left\langle l m ; \left.\frac{1}{2} m_{s} \right\rvert\, l \frac{1}{2} ; j m_{j}\right\rangle\left|l m ; \frac{1}{2} m_{s}\right\rangle .
$$

As only the Clebsch-Gordan coefficients satisfying $m_{j}=m+m_{s}$ may not vanish, taking into account that $m_{s}$ takes only values $\pm \frac{1}{2}$, this sum (for $l>0$ ) reduces to:

$$
\begin{equation*}
\left|l \frac{1}{2} ; j m_{j}\right\rangle=c_{+}\left|l m_{j}-\frac{1}{2} ; \frac{1}{2} \frac{1}{2}\right\rangle+c_{-}\left|l m_{j}+\frac{1}{2} ; \frac{1}{2},-\frac{1}{2}\right\rangle, \quad c_{ \pm}=\left\langle l m_{j} \mp \frac{1}{2} ; \frac{1}{2}, \left. \pm \frac{1}{2} \right\rvert\, l \frac{1}{2} ; j m_{j}\right\rangle . \tag{4.60}
\end{equation*}
$$

The coordinate representation (only angular coordinates are considered, and in spin space vectors $\left\langle\frac{1}{2}, \frac{1}{2}\right\rangle=\binom{1}{0}$ and $\left.\left|\frac{1}{2},-\frac{1}{2}\right\rangle=\binom{0}{1}\right)$ becomes:

$$
\begin{equation*}
\left\langle\theta \varphi \left\lvert\, l \frac{1}{2}\right. ; j m_{j}\right\rangle=\binom{c_{+} Y_{l}^{m_{j}-\frac{1}{2}}(\theta, \varphi)}{c_{-} Y_{l}^{m_{j}+\frac{1}{2}}(\theta, \varphi)} \tag{4.61}
\end{equation*}
$$

This spinor function is called spin (spherical) harmonic.

## Ground state of the helium atom

Helium atom besides nucleus contains two electrons. Therefore the electronic hamiltonian besides noninteracting par $H_{0}$ contains interaction:

$$
\begin{equation*}
H=H_{0}+\frac{e^{2}}{r_{12}}, \quad H_{0}=h_{1}+h_{2}, \quad h_{i}=T_{i}-\frac{2 e}{r_{i}} . \tag{4.62}
\end{equation*}
$$

Here, $h_{i}$ are single particle hamiltonians, with eigenenergies $\epsilon_{n}=-\frac{2 m e^{2}}{n^{2} \hbar^{2}}$ and eigenfunctions $\langle r, \theta, \varphi \mid n l m\rangle=R_{n l}(r) Y_{l}^{m}(\theta, \varphi)$. We consider mutual interaction as perturbation, and $H_{0}$ as the nonperturbed part.

The ground single electron states are $n=1, l=m=0$. This gives total state with $L=0$; as it is known that the ground state is with $S=0$ (changing sign when two particles are interchanged, i.e. antisymmetrized state Sec. 5.1), thus state $\left|\frac{1}{2} \frac{1}{2} 00\right\rangle$ in spin space, and therefore $J=0$. Thus, orbital ground state is symmetrized. This gives

$$
\Psi\left(\boldsymbol{r}_{1}, \boldsymbol{r}_{2}\right)=\frac{8}{\sqrt{2} \pi a_{0}^{3}} \mathrm{e}^{-2 \frac{r_{1}+r_{2}}{a_{0}}}(|+-\rangle-|-+\rangle)
$$

The single electron energy is $\epsilon_{1}=-54.4 \mathrm{eV}$, i.e. nonperturbed energy of the ground state is $E_{0}^{0}=2 \epsilon_{1}=-108.8 \mathrm{eV}$. Taking into account integral $\int \frac{\mathrm{e}^{-\left(a r_{1}+b r_{2}\right)}}{\left|r_{1}-r_{2}\right|} \mathrm{d} \boldsymbol{r}_{1} \mathrm{~d} \boldsymbol{r}_{2}=16 \pi^{2} \frac{2 a^{2}+2 b^{2}+6 a b}{a^{2} b^{2}(a+b)^{2}}$, one easily finds that the first correction $E_{0}=E_{0}^{0}+E_{0}^{(1)}=-74.8 \mathrm{eV}$. Note that the experimental value is $E_{0}^{\exp }=-78.6 \mathrm{eV}$.

## Stark effect of diatomic molecule

We assume that in the diatomic molecule distance between atoms is fixed, i.e. that the vibrational energies are too high to be excitet by the electrical field. Therefore, only rotational degrees of freedom remain, i.e. in the homogeneous electrical field $E$ along $z$-axis the effective hamiltonian is $H=H_{0}+H^{\prime}$, where $H_{0}=\boldsymbol{l}^{2} / 2 m r_{0}^{2}$ and perturbation is $H^{\prime}=-d E \cos \theta$; here $\boldsymbol{d}=q_{1} \boldsymbol{r}_{1}+q_{2} \boldsymbol{r}_{2}$ is dipole moment of the molecule, and $\theta$ the angle between the molecule and $z$-axis. The nonperturbed term is kinetic rotational energy with eigenvalues and eigenfunctions $E_{l}^{0}=l(l+1) \hbar^{2} / 2 m r_{0}^{2}$ and $Y_{l}^{m}(\theta, \varphi)$, respectively. Perturbation is $z$-component of the vector operator, and since $Y_{1}^{0}(\theta, \varphi) \sim \cos \theta$ we can easily calculate matrix elements:

$$
\langle l m| H^{\prime}\left|l^{\prime} m^{\prime}\right\rangle \sim \int_{0}^{\pi} \int_{0}^{2 \pi} \sin \theta \mathrm{~d} \theta \mathrm{~d} \varphi Y_{l}^{m *}(\theta, \varphi) Y_{1}^{0}(\theta, \varphi) Y_{l^{\prime}}^{m^{\prime}}(\theta, \varphi)
$$

We see that nonzero matrix elements appear only when selection rules $\Delta m=m-m^{\prime}=0$ and $\Delta l=l-l^{\prime}=0, \pm 1$ are satisfied: the first one is obvious conservation of the $z$ component of angular momentum, while for the second one note that $D^{(l)}$ must be a component of $D^{(1)} \otimes D^{\left(l^{\prime}\right)}=D^{\left(l^{\prime}-1\right)} \oplus D^{\left(l^{\prime}\right)} \oplus D^{\left(l^{\prime}+1\right)}$.

Analogous considerations for spatial inversion exclude the term $l^{\prime}$, as on the left side the parity is $\left.(-1)^{( } l^{\prime}+1\right)$. This immediately means that the first order correction vanishes. The second order perturbation is

$$
E_{l}^{(2)}=-\sum_{l>0} \frac{\left.\sum_{m=-l}^{l}\left|\langle l=0, m=0| H^{\prime}\right| l m\right\rangle\left.\right|^{2}}{l(l+1) \hbar^{2} / m r_{0}^{2}}=-\frac{\left.\left|\langle l=0, m=0| H^{\prime}\right| l=1, m=0\right\rangle\left.\right|^{2}}{l(l+1) \hbar^{2} / m r_{0}^{2}}
$$

## Stark effect in hydrogen atom

The hamiltonian of the electron in the homogeneous electric field along $z$-axis is $H=$ $H_{0}+H^{\prime}$, where nonperturbed part is usual hamiltonian of the hydrogen atom electron, while perturbation is $H^{\prime}=-e E z$. The state space includes spin, though the hamiltonian is spin independent, and results only in additional degeneracy.

As the perturbation quantum numbers are $l=1$ and $m=0$, while parities of bra, berturbation and ket are $(-1)^{l},-1$, and $(-1)^{l^{\prime}}$, the selection rules give:

$$
\left\langle n l m m_{s}\right| H^{\prime}\left|n^{\prime} l^{\prime} m^{\prime} m_{s}^{\prime}\right\rangle=\delta_{m_{s}, m_{s}^{\prime}} \delta_{m m^{\prime}}\left(\delta_{l, l^{\prime}+1} V_{n n^{\prime}}^{+}+\delta_{l, l^{\prime}-1} V_{n n^{\prime}}^{-}\right) .
$$

Let us consider level $n=2$, which is fourfold degenerate (nonperturbed) with basis $|11\rangle,|10\rangle,|1,-1\rangle$ and $|00\rangle$. The perturbation matrix must have vanishing elements when $m \neq m^{\prime}$, as well as when $l=l^{\prime}$. This means that the linear effect can be found only in the subspace spanned by the $m=0$ vectors $|210\rangle$ and $|200\rangle$; reduced matrix is $H^{\prime}=e E\left(\begin{array}{c}\langle 210| z|210\rangle=0 \\ \langle 210| z|200\rangle \\ \langle 200| z|z| 200\rangle=0\end{array}\right)$, and the eigenvalues are $\left.E^{(1)}= \pm|\langle 210| z| 200\right\rangle||e| E=$ $\pm 3 e E a_{0}$.

### 4.6.3 $\quad L-S$ and $j-j$ coupling

The electronic system in the atom with $Z$ electrons is described by the hamiltonian:

$$
\begin{equation*}
H=\sum_{i=1}^{Z}\left(\frac{p_{i}^{2}}{2 m}-\frac{Z e^{2}}{r_{i}}+v_{i}^{j j}\right)+\sum_{i<j} \frac{e^{2}}{\left|\boldsymbol{r}_{i}-\boldsymbol{r}_{j}\right|} \tag{4.63}
\end{equation*}
$$

In the single-particle part besides usual kinetic and Coulomb terms there is spin-orbit interaction $v_{i}^{j j}=\boldsymbol{l}_{i} \cdot \boldsymbol{s}_{i} f\left(r_{i}\right)$, the term commuting with $\boldsymbol{j}=\boldsymbol{l}+\boldsymbol{s}$. Intuitively it appears as the magnetic dipole of an electron (proportional to its spin) moves in the electric field of the nucleus. Given form can be justified by the perturbation theory within relativistic approach. Depending on the order of the perturbation theory, the form of the function $f(r)$ can be specified (e.g. $f(r)=\frac{m^{2} c^{2}}{r} \frac{\mathrm{~d}}{\mathrm{~d} r} \frac{-Z e^{2}}{r}$ ), and for the present purpose it is important only that it is spherically symmetric.

The last term includes many-electron problem, and we assumed that only two-particle interactions exist. Such terms, related to many-body systems, prevent exact solution
of the problem, and this is probably the most important difficulty in the contemporary physics. Usual approximate approach is to extract some average single electron potential, i.e. to write the last term in the form

$$
\begin{equation*}
\sum_{i<j} \frac{e^{2}}{\left|\boldsymbol{r}_{i}-\boldsymbol{r}_{j}\right|}=\sum_{i} W\left(r_{i}\right)+\sum_{i<j} V_{i j}^{\mathrm{rez}} \tag{4.64}
\end{equation*}
$$

The residual potential $V^{L S}=\sum_{i<j} V_{i j}^{\text {rez }}$ is the remaining part of the interaction which has not been included by the averaged potential $W$ (more detailed analysis of this approach will be given in the Section 6.3.2). It is assumed to be spin independent.

Adding the averaged potential $W$ to the Coulomb electron-nucleus term, we write the hamiltonian as the sum of the single-particle central hamiltonian $H_{0}$, two-particle term $V^{L S}$, and spin-orbit (still single particle) term $V^{j j}$ :

$$
\begin{equation*}
H=H_{0}+V^{L S}+V^{j j} . \tag{4.65}
\end{equation*}
$$

Of course, the total hamiltonian commutes with total angular momentum $\boldsymbol{J}=\boldsymbol{L}+\boldsymbol{S}$. However, it is assumed that many-body term depends only on the orbital variables, and commutes both with total orbital and spin angular momenta $\boldsymbol{L}$ and $\boldsymbol{S}$. On the other hand, coupling electronic orbital and spin degrees of freedom of each electron, $V^{j j}$ commutes with $\boldsymbol{j}$. Therefore, further analysis depends on the relation between spin-orbit and manybody part. It turns out that many-body term is weakly dependent on the atomic number, while spin-orbit term is proportional to it. Therefore, $V^{L S}$ is the dominant for the light atoms, while $V^{j j}$ overwhelms in the case in the second half of periodic system.

Consequently, as the first perturbation to $H_{0}$ for light atoms we take $V^{L S}$ term, getting Russel-Saunders or $L S$-coupling hamiltonian $H^{L S}=H_{0}+V^{L S}$; only and afterward the last term $V^{j j}$ is included as the perturbation. Corresponding decomposition of the total space is

$$
\mathcal{S}=\mathcal{S}_{O} \otimes \mathcal{S}_{S}=\left(\mathcal{S}_{O 1} \otimes \cdots \otimes \mathcal{S}_{O Z}\right) \otimes\left(\mathcal{S}_{S 1} \otimes \cdots \otimes \mathcal{S}_{S Z}\right)
$$

with momenta $\boldsymbol{L}=\boldsymbol{l}_{1}+\ldots \boldsymbol{l}_{Z}$ and $\boldsymbol{S}=\boldsymbol{s}_{1}+\ldots \boldsymbol{s}_{Z}$ acting in $\mathcal{S}_{O}$ and $\mathcal{S}_{S}$. There is an eigen-basis $\left|L M ; S M_{S} ; \lambda\right\rangle$ for $H^{L S}$ being standard both for $\boldsymbol{L}$ and $\boldsymbol{S}$, with $(2 L+1)(2 S+1)$ degenerate eigenspaces. When we add $V^{j j}$, the perturbation theory predicts that each of these levels $E_{L S \lambda}$ will split: for each $J=|L-S|, \ldots, L+S$ an sublevel $E_{L S J \lambda}$ appears.

For heavy atoms $V^{j j}$ interaction prevails, and the procedure is reversed. Now we start with the $j j$-coupling, i.e. with $H_{j j}=H_{0}+V^{j j}$. As this describes non-interacting particles in the external field, for each electron the eigen-basis $\left|n l j m_{j}\right\rangle$ can be found, and then the total eigenbasis can be adapted to the set of the compatible observables $J^{2}, J_{3}, j_{1}^{2}, j_{2}^{2}, \ldots$, corresponding to decomposition :

$$
\mathcal{S}=\mathcal{S}_{1} \otimes \cdots \otimes \mathcal{S}_{Z}, \quad \mathcal{S}_{i}=\mathcal{S}_{O i} \otimes \mathcal{S}_{S i}
$$

In each space $\mathcal{S}_{i}$ acts total electronic angular momentum $\boldsymbol{j}_{i}=\boldsymbol{l}_{i}+\boldsymbol{s}_{i}$. The degeneracy of the eigenenergies depends on $Z$, and it is in general high. When $V^{L S}$ is included, only $J^{2}$ and $J_{3}$ remain conserved, and this leads to the splitting of the obtained levels to the $2 J+1$ degenerate ones.

## Periodical system of elements

Classification of the states of the hydrogen atom and derived rules for addition of the angular momentum give insight to the systematization of the atoms with more electrons according to angular momenta (orbital, spin and total). To this end we anticipate that state of many-electron system is antisymmetric (i.e. changes sign when any two of electrons are permuted), and consequently Pauli principle (to be derived in the next Chapter) forbidding two electrons to occupy the same state.

At first we make inventory of the involved angular momenta and introduce conventional notation. Each electron has orbital, spin and total angular, denoted as $\boldsymbol{l}, \boldsymbol{s}$ and $\boldsymbol{j}=\boldsymbol{l}+\boldsymbol{s}$. For the complete electronic system total orbital, spin and angular momentum, $\boldsymbol{L}=\sum_{i} \boldsymbol{l}_{i}, \boldsymbol{S}=\sum_{i} \boldsymbol{s}_{i}$ and $\boldsymbol{J}=\sum_{i} \boldsymbol{j}_{i}$ are defined. For the states with orbital momentum quantum number $l$ equal to $0,1,2,3$, etc. we use symbols $s, p d, f$ etc. (and analogously $S, P, D, F, \ldots$ for $L)$.

For each main quantum number $n=1,2, \ldots$, possible orbital momenta are $l=$ $0, \ldots, n-1$; in other words, shell $n$ contains subshells $n l$ with these values of $l$. Each of them contains $2(2 l+1)$ states with $m=0, \pm 1, \ldots, \pm l$ and $m_{s}= \pm \frac{1}{2}$. Starting with known eigenstates and eigenenergies, we begin to fill the states with electrons, starting from the lowest levels.

Thus, for $n=1$, only $l=m=0$ is allowed, and the ground state of hydrogen atom is $\left|n=1, l=0, m=0 ; s=1 / 2, m_{s}\right\rangle$, with total momenta $L=l=0, S=s=1 / 2, J=$ $j=1 / 2$. Spectroscopic notation of these angular momenta and electronic configuration is ${ }^{2 S+1} L_{J}={ }^{2} S_{1 / 2}$ and $(n l)=(1 s)$. For the next atoms the higher energy levels are occupied, and angular momenta are given in the table (note that $L S$-coupling is assumed, in agreement with previous conclusions). Generally, for pairs of electrons the orbital part of state is allowed to be the same, and then the spin part is antisymmetric.

Note that this algorithm is not valid for all elements. For example $d$-shell is open before $3 p$ is filled, etc.

Table 4.1: Beginning of the periodic system. Horizontal lines denote fulfilled subshells, while the shells are separated by double lines; both have all angular momenta zero. Quantum numbers $n, l$ and $m_{s}$ are given only for the electrons out of the last filled (sub)shell. For He there was another possible total spin $S=1$, which was rejected by Pauli principle. On the other hand, $J=1 / 2$ is preferred for B by energy with respect to another possibility $J=3 / 2$.

| $Z$ | Symbol | $n$ | $l$ | $m_{s}$ | $L$ | $S$ | $J$ | ${ }^{2 S+1} L_{J}$ | Configuration |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| 1 | H | 1 | 0 | $\pm \frac{1}{2}$ | 0 | $\frac{1}{2}$ | $\frac{1}{2}$ | ${ }^{2} S_{\frac{1}{2}}$ | $(1 s)$ |
| 2 | He | 1,1 | 0,0 | $\frac{1}{2},-\frac{1}{2}$ | 0 | 0 | 0 | ${ }^{1} S_{0}$ | $(1 s)^{2}$ |
| 3 | Li | 2 | 0 | $\pm \frac{1}{2}$ | 0 | $\frac{1}{2}$ | $\frac{1}{2}$ | ${ }^{2} S_{\frac{1}{2}}$ | $(\mathrm{He})(2 s)$ |
| 4 | Be | 2,2 | 0,0 | $\frac{1}{2},-\frac{1}{2}$ | 0 | 0 | 0 | ${ }^{1} S_{0}$ | $(\mathrm{He})(2 s)^{2}$ |
| 5 | B | 2 | 1 | $\pm \frac{1}{2}$ | 1 | $\frac{1}{2}$ | $\frac{1}{2}$ | ${ }^{2} P_{\frac{1}{2}}$ | $(\mathrm{He})(2 s)^{2}(2 p)$ |

## Chapter 5

## Identical Particles

### 5.1 Quantum Formalism

### 5.1.1 Permutational Indistinguishability and Symmetrization

According to the conclusion of the subsection 1.3.2, the state space of $N$ particles is the product of their single particle states. Thus, in the case of $N$ identical particles, when all the particles are described by the same single particle state space $\mathcal{S}$, the total state space is a priori constructed as $\mathcal{S}^{N}=\mathcal{S} \otimes \cdots \otimes \mathcal{S}$. On the other hand, the intuitive notion of the identity of the particles, means that permutations among them cannot be observable. In other words, no measurement can distinguish between initial and permuted state. As for the pure states this means that the initial and the permuted state vectors are collinear, differing at most by a phase. However, in the space $\mathcal{S}^{N}$ there are vectors not obeying this requirement. For example, consider non-correlated states, with $p^{\text {th }}$ particle being in the state $\left|\psi_{p}\right\rangle$. The composite state is shortly denoted as

$$
\begin{equation*}
\left|\psi_{1}, \ldots, \psi_{N}\right\rangle=\left|\psi_{1}\right\rangle_{1} \otimes \cdots \otimes\left|\psi_{N}\right\rangle_{N} . \tag{5.1}
\end{equation*}
$$

Note that the $p$ th particle is in the state being at the $p$ th place, and therefore the second index (outside ket) may be omitted. Any permutation $\pi$ of the particles transform this state into

$$
\begin{equation*}
\Delta(\pi)\left|\psi_{1}, \ldots, \psi_{N}\right\rangle \stackrel{\text { def }}{=} \cdots \otimes\left|\psi_{1}\right\rangle_{\pi 1} \otimes \cdots \otimes\left|\psi_{N}\right\rangle_{\pi N} \otimes \cdots=\left|\psi_{\pi^{-1} 1}, \ldots, \psi_{\pi^{-1} N}\right\rangle \tag{5.2}
\end{equation*}
$$

Now, if $\pi$ nontrivially acts on the particles being in the orthogonal single particle states, the initial and permuted states are mutually orthogonal, and obviously not collinear.

This shows that the construction of the state space of identical particles must be further suited to the intuitive requirement of indistinguishability. To this end we emphasize that (5.2) effectively introduces the operators $\Delta(\pi)$ in $\mathcal{S}^{N}$ associated to the permutation $\pi$. Indeed, any basis $|i\rangle(i=1, \ldots,|\mathcal{S}|)$ of $\mathcal{S}$ induces the non-correlated basis $\left|i_{1}, \ldots, i_{N}\right\rangle$ ( $i_{p}=1, \ldots,|\mathcal{S}|$ ), and (5.2) is applicable for all the vectors of the basis. This suffices to define linear operators $\Delta(\pi)$. In addition, all the basis vectors are by $\Delta(\pi)$ mapped to the vectors of the same basis, i.e. $\Delta(\pi)$ permutes the basis; this means that $\Delta(\pi)$ is a unitary representation (orthonormal single particle basis is assumed) of the symmetric group $\boldsymbol{S}_{N}$ in $\mathcal{S}^{N}$.

The basic intuitive requirement of indistinguishability of the initial $|\Psi\rangle$ and the permuted state $\Delta(\pi)|\Psi\rangle$ (conveniently, for $N$ particle state capital letters are used) can be now precisely expressed as:

$$
\begin{equation*}
\Delta(\pi)|\Psi\rangle=\mathrm{e}^{\mathrm{i} \varphi(\pi, \Psi)}|\Psi\rangle \tag{5.3}
\end{equation*}
$$

with the phase depending both on permutation and state. In other words, each physically allowable state $|\Psi\rangle$ spans one dimensional subspace of $\mathcal{S}^{N}$, in which the representation $\Delta\left(\boldsymbol{S}_{N}\right)$ is reduced to the phases $\mathrm{e}^{\mathrm{i} \varphi(\pi, \Psi)}$. Of course, these sub-representations are one dimensional, thus irreducible ones. However, the permutation group $\boldsymbol{S}_{N}$ for $N>1$ has exactly two one-dimensional irreducible representations (Exercise 5.1), symmetric (identical) $D^{(+)}$and antisymmetric (alternating) $D^{(-)}$one:

$$
\begin{equation*}
D^{( \pm)}(\pi)=( \pm)^{\tilde{\pi}} \tag{5.4}
\end{equation*}
$$

Here, $\tilde{\pi}$ denotes the parity of the permutation $\pi$ : each permutation is a product of the transpositions, and although this decomposition is not unique, as well as the number of the involved transpositions, this number is uniquely either even, $\tilde{\pi}=0$ or odd $\tilde{\pi}=1$. Particularly, $D^{( \pm)}\left(\tau_{p p^{\prime}}\right)= \pm 1$ for the transposition $\tau_{p p^{\prime}}$ of any two particles $p$ and $p^{\prime}$.
Exercise 5.1: Show that the representations (5.4) are the only one-dimensional irreducible representations of $\boldsymbol{S}_{N}$ for $N=2,3, \ldots$.

According to the above analysis we conclude that the permutation group $\boldsymbol{S}_{N}$ is in $\mathcal{S}^{N}$ represented by $\Delta\left(\boldsymbol{S}_{N}\right)$, being reducible, with a number of irreducible components. However, physically allowable states form multiple irreducible subspaces $\mathcal{S}_{ \pm}^{N}$ of the two one dimensional representations $D^{( \pm)}\left(\boldsymbol{S}_{N}\right)$, and therefore these subspaces are singled out as the ranges of the corresponding group projectors:

$$
\begin{equation*}
P^{( \pm)}=\frac{1}{N!} \sum_{\pi}( \pm)^{\tilde{\pi}} \Delta(\pi) . \tag{5.5}
\end{equation*}
$$

These projectors onto symmetric and antisymmetric subspaces $\mathcal{S}_{ \pm}^{N}$ of $\mathcal{S}^{N}$ are called symmetrizer and antisymmetrizer. The only physically allowable states are within these subspaces, and for $|\Psi\rangle \in \mathcal{S}_{+}^{N}$ the phase in (5.3) is $\mathrm{e}^{\mathrm{i} \varphi(\pi, \Psi)}=1$, while for $|\Psi\rangle \in \mathcal{S}_{-}^{N}$ one gets $\mathrm{e}^{\mathrm{i} \varphi(\pi, \Psi)}=(-)^{\tilde{\pi}}$. However, any nontrivial linear combination $\alpha_{+}|+\rangle+\alpha_{-}|-\rangle$of
 is not collinear to the initial vector. To retain the superposition principle, the system of identical particles must be described by one of these two subspaces only.

These conclusions are summarized in:

## POSTULATE 7. - IDENTICAL PARTICLES

The state space of a system of $N$ identical particles with the single particle state space $\mathcal{S}$ is either symmetric $\mathcal{S}_{+}^{N}$ or antisymmetric $\mathcal{S}_{-}^{N}$ subspace of the direct product space $\mathcal{S}^{N}$. The choice depends only on the type of the particles: bosons, particles with the integer spin, have the symmetrized $N$ particle states, while for the fermions, the particles with the half-integer spin, states are antisymmetrized.

The last part of the Postulate, relating spin to the type of the transformation of the particle states under permutations can be proved within relativistic quantum mechanics,
using the locality of interaction [?]. This has deep influence not only to the kinematical, but also to the dynamical properties of the particles: as a consequence of the different permutational symmetry for the systems of fermions and bosons the different statistics, Bose-Einstein or Fermi-Dirac, are applied.

### 5.1.2 Structure of the state space - occupation numbers

The permutational symmetry of the physical states gives further insight to the structure of the boson and fermion state spaces. At first, it should be noticed that (5.2) shows that the operators $\Delta\left(\boldsymbol{S}_{N}\right)$ in $\mathcal{S}^{N}$ permute the non-correlated basis vectors $\left|i_{1}, \ldots, i_{N}\right\rangle$. Indeed, their action consists in the rearrangement of the single particle states quantum numbers $i_{1}, \ldots, i_{N}$. This inspires introduction of the occupation numbers: for the $N$ particle state $\left|i_{1}, \ldots, i_{N}\right\rangle$ one counts the number $n_{i}\left(\left|i_{1}, \ldots, i_{N}\right\rangle\right)$ of the particles occupying the single particle state $|i\rangle$. Thus, each basis state defines $|\mathcal{S}|$ occupation numbers $\boldsymbol{n}\left(\left|i_{1}, \ldots, i_{N}\right\rangle\right)=$ $\left(n_{1}, \ldots, n_{|\mathcal{S}|}\right)$, with the obvious relation $\sum_{i=1}^{|\mathcal{S}|}=N$. More than one basis vector may have same array $\boldsymbol{n}$. Precisely, all the permutations acting by the representation $\Delta\left(\boldsymbol{S}_{N}\right)$ on the vector $\left|i_{1}, \ldots, i_{N}\right\rangle$ give the vectors with the same occupation numbers. Hence, the non-correlated basis is by the action of $\boldsymbol{S}_{N}$ partitioned into the orbits, each of them characterized by the unique $\boldsymbol{n}$. In other words, each $|\mathcal{S}|$-tuple $\boldsymbol{n}$ of the non-negative integers $n_{s}$ such that $\sum_{s=1}^{|\mathcal{S}|} n_{s}=N$ defines an orbit of the non-correlated basis.

The permutations interchanging particles being in the same single particle states leave the basis vector invariant (e.g. $\left.\Delta\left(\tau_{12}\right)\left|i_{1}, i_{2}=i_{1}, i_{3}, \ldots, i_{N}\right\rangle=\left|i_{1}, i_{2}=i_{1}, i_{3}, \ldots, i_{N}\right\rangle\right)$. Taking from each orbit $\boldsymbol{n}$ a representative basis vector $\left|i_{1}, \ldots, i_{N}\right\rangle$, it is easy to find its stabilizer: it is the group $\boldsymbol{S}_{\boldsymbol{n}}=\boldsymbol{S}_{n_{1}} \otimes \cdots \otimes \boldsymbol{S}_{n_{|\mathcal{S}|}} ; \boldsymbol{S}_{n_{s}}$ contains only the permutations interchanging the particles in the same state $\left|i_{s}\right\rangle$ and leaving invariant all other particles (therefore different factors permute disjoint subsets of the particles and commute mutually). Obviously, the order of the stabilizer is $\left|\boldsymbol{S}_{\boldsymbol{n}}\right|=\boldsymbol{n}!\stackrel{\text { def }}{=} \prod_{s=1}^{|\mathcal{S}|} n_{s}$ !, showing that the number of the basis vectors in the same orbit of $\boldsymbol{S}_{N}$ is

$$
\begin{equation*}
|\boldsymbol{n}|=\frac{\left|\boldsymbol{S}_{N}\right|}{\left|\boldsymbol{S}_{\boldsymbol{n}}\right|}=\frac{N!}{n_{1}!\cdots n_{|\mathcal{S}|}!}=\frac{N!}{\boldsymbol{n}!} . \tag{5.6}
\end{equation*}
$$

Since in each factor $\boldsymbol{S}_{n_{s}}$ for $n_{s}>1$ one half of the permutations are even and the other half odd, the same is valid for the whole stabilizer $\boldsymbol{S}_{\boldsymbol{n}}$ : one half of its permutations are even and the rest are odd; only when all $n_{s}$ are equal to one, only the identity element, which is even, is in the stabilizer.

Now, recall that each orbit $\boldsymbol{n}$ is a subset of the non-correlated basis. Therefore, it spans a $|\boldsymbol{n}|$-dimensional subspace $\mathcal{S}_{\boldsymbol{n}}$ of $\mathcal{S}^{N}$, and

$$
\begin{equation*}
\mathcal{S}^{N}=\oplus_{n} \mathcal{S}_{n} \tag{5.7}
\end{equation*}
$$

As $\mathcal{S}^{N}$ is spanned by the whole non-correlated basis, one finds $\sum|\boldsymbol{n}|=|\mathcal{S}|^{N}$. Each of these subspaces $\boldsymbol{S}_{\boldsymbol{n}}$ is invariant under the permutation operators $\Delta$, showing that (5.7) is a decomposition of the total space onto the invariant subspaces of the representation $\Delta\left(\boldsymbol{S}_{N}\right)$. Hence, this representation reduces in the occupation number spaces $\mathcal{S}_{n}$ onto the sub-representations $\Delta_{n}\left(\boldsymbol{S}_{N}\right)$ geting the reduced form $\Delta\left(\boldsymbol{S}_{N}\right)=\oplus_{n} \Delta_{n}\left(\boldsymbol{S}_{N}\right)$.

Finally, we perform (anti)symmetrization, as requested by the Postulate VII. The symmetric/antisymmetric space $\mathcal{S}_{ \pm}^{N}$ is found as the range of the (anti)symmetrizer (5.5). The invariance of $\mathcal{S}_{n}^{N}$, resulting in decomposition of $\Delta\left(\boldsymbol{S}_{N}\right)$, implies that the group projectors also reduce in the occupation number subspaces:

$$
\begin{equation*}
P^{( \pm)}=\oplus_{n} P_{n}^{( \pm)}, \quad P_{n}^{( \pm)}=\frac{1}{N!} \sum_{\pi}( \pm)^{\tilde{\pi}} \Delta_{\boldsymbol{n}}(\pi) . \tag{5.8}
\end{equation*}
$$

Due to the orthogonality of the subspaces $\mathcal{S}_{n}^{N}$, the operators $P_{n}^{( \pm)}$are projectors, and their ranges $\mathcal{S}_{n \pm}^{N}$ are orthogonally summed into the range of $P^{( \pm)}$. The first conclusion emerges: the (anti)symmetrized space $\mathcal{S}_{ \pm}^{N}$ is orthogonal sum of the (anti)symmetrized subspaces $\mathcal{S}_{n \pm}^{N}$ of the occupation number spaces $\mathcal{S}_{n}^{N}$.

To find these ranges we chose for each orbit the orbit representative. We sort the orbit elements in the increasing order with respect to the single particle state indices, and chose the first of them for the representative; effectively, we choose the non-correlated vector $|\boldsymbol{n}\rangle=\left|i_{1} \leq i_{2} \leq \cdots \leq i_{n}\right\rangle$ for which from $i_{p} \leq i_{p^{\prime}}$ follows that $p \leq p^{\prime}$, i.e. the single particle state indices do not decrease from left to right (e.g. for $\mathcal{S}=7$ and $N=6,|5,2,5,3,2,5\rangle$ is in the orbit $\boldsymbol{n}=(0,2,1,0,3,0,0)$ represented by $|\boldsymbol{n}\rangle=|2,2,3,5,5,5\rangle$; the stabilizer of this state is $\boldsymbol{S}_{0} \otimes \boldsymbol{S}_{2} \otimes \boldsymbol{S}_{1} \otimes \boldsymbol{S}_{0} \boldsymbol{S}_{3} \otimes \boldsymbol{S}_{0} \boldsymbol{S}_{0}=\boldsymbol{S}_{2} \otimes \boldsymbol{S}_{3}$, with $|\boldsymbol{n}|=12$ elements). According to the general group theoretical theorem, $\boldsymbol{S}_{\boldsymbol{N}}$ acts such that exactly $\left|\boldsymbol{S}_{\boldsymbol{n}}\right|$ permutations map $|\boldsymbol{n}\rangle$ to each element of the orbit $\boldsymbol{n}$. Therefore, only the stabilizer elements of the sum in $P_{\boldsymbol{n}}^{( \pm)}$contribute to $\langle\boldsymbol{n}| P_{\boldsymbol{n}}^{( \pm)}|\boldsymbol{n}\rangle$, giving $\operatorname{Tr} P_{\boldsymbol{n}}^{( \pm)}=\frac{1}{\boldsymbol{n}!} \sum_{\pi}( \pm)^{\tilde{\pi}}\langle\boldsymbol{n}| P_{\boldsymbol{n}}^{( \pm)}|\boldsymbol{n}\rangle$.

## Bosons

In the case of $P_{n}^{(+)}$the contributions of all the stabilizer elements are equal to $1 / N!$; thus $\langle\boldsymbol{n}| P_{\boldsymbol{n}}^{(+)}|\boldsymbol{n}\rangle=\boldsymbol{n}!/ N!$. This holds for all $|\boldsymbol{n}|$ vectors of the orbit $\boldsymbol{n}$, and the trace of $P_{\boldsymbol{n}}^{(+)}$ is $|\boldsymbol{n}|\langle\boldsymbol{n}| P_{\boldsymbol{n}}^{(+)}|\boldsymbol{n}\rangle=1$. Thus, in the case of bosons, each occupation number subspace $\mathcal{S}_{\boldsymbol{n}}^{N}$ contains exactly one dimensional symmetric subspace spanned by the normalized symmetric vector $\left|\boldsymbol{n}^{+}\right\rangle$obtained by the action of the symmetrizer onto $|\boldsymbol{n}\rangle=\left|i_{1} \leq \cdots \leq i_{N}\right\rangle$ :

$$
\begin{equation*}
\left|\boldsymbol{n}^{+}\right\rangle=\sqrt{\frac{N!}{n_{1}!\cdots n_{|\mathcal{S}|}}} P_{\boldsymbol{n}}^{(+)}|\boldsymbol{n}\rangle=\sqrt{\frac{1}{n_{1}!\cdots n_{|\mathcal{S}|}!N!}} \sum_{\pi}\left|i_{\pi^{-1} 1}, \ldots, i_{\pi^{-1} N}\right\rangle . \tag{5.9}
\end{equation*}
$$

The normalization factor is easily found, taking into account that the whole orbit with $|\boldsymbol{n}|$ orthonormal vectors appear after the action of the permutations, and each of them exactly $\boldsymbol{n}$ ! times.

Thus $\mathcal{S}_{+}^{N}$ is orthogonal sum of the one-dimensional symmetrized occupation number subspaces:

$$
\begin{equation*}
\mathcal{S}_{+}^{N}=\sum_{n} \mathcal{S}_{n+}^{N}, \quad \mathcal{S}_{n+}^{N}=\operatorname{Span}\left(\left|\boldsymbol{n}^{+}\right\rangle\right) . \tag{5.10}
\end{equation*}
$$

It is obvious that each occupation number $\boldsymbol{n}$ uniquely defines the symmetrized vector $\left|\boldsymbol{n}^{+}\right\rangle$, and in this sense the occupation numbers are basis in the bosonic space.

## Fermions

Analogously to the bosons, the space of fermions is to be found with help of $P_{n}^{(-)}$. However, if any $n_{s}$ is greater than one, only the even half of the stabilizer elements contribute to $\langle\boldsymbol{n}| P_{\boldsymbol{n}}^{(-)}|\boldsymbol{n}\rangle$ by $1 / N$ !, and the odd part by $-1 / N$ !. Thus, if any occupation number is greater than one, the matrix element vanishes, meaning that the trace of $P_{n}^{(-)}$is zero, and such $\mathcal{S}_{\boldsymbol{n}}$ has no antisymmetric physical states. Consequently, only the spaces with all $n_{s}=0,1$, i.e. with $\boldsymbol{n}!=1$, contribute to the fermionic state space $\mathcal{S}_{-}^{N}$, by the one dimensional subspace space spanned by the antisymmetric vector

$$
\begin{equation*}
\left|\boldsymbol{n}^{-}\right\rangle=\sqrt{N!} P_{\boldsymbol{n}}^{(-)}|\boldsymbol{n}\rangle=\sqrt{\frac{1}{N!}} \sum_{\pi}(-)^{\tilde{\pi}}\left|i_{\pi^{-1} 1}<\cdots<i_{\pi^{-1} N}\right\rangle . \tag{5.11}
\end{equation*}
$$

Again, $\mathcal{S}_{-}^{N}$ is the orthogonal sum of the one-dimensional symmetrized occupation number subspaces, but the sum includes only the terms with $\boldsymbol{n}$ having components $n_{s}=$ 0,1 :

$$
\begin{equation*}
\mathcal{S}_{-}^{N}=\sum_{n\left(n_{s}<2\right)} \mathcal{S}_{n^{-}}^{N}, \quad \mathcal{S}_{n^{-}}^{N}=\operatorname{Span}\left(\left|\boldsymbol{n}^{-}\right\rangle\right) . \tag{5.12}
\end{equation*}
$$

Clearly, occupation number $\boldsymbol{n}$ satisfying $n_{s}<2$ uniquely defines the antisymmetrized vector $\left|\boldsymbol{n}^{-}\right\rangle$, and such occupation numbers are basis in the fermionic space.

### 5.1.3 States

## Pure states

We have shown that the substantial difference between bosons and fermions was the type of symmetrization. This has far reaching consequences on the behavior of the two type of particles.

The most important one is the already emphasized constraint $n_{i}=0,1$ to the fermionic occupation numbers. In fact, this is the famous Pauli exclusion principle: two fermions cannot be in the same state. The most frequent example is that at most two electrons may have the same orbital wave function $\psi(\boldsymbol{r})$ : then their states $\left|\psi, m_{s}\right\rangle$ differ in the spin space by the quantum number $m_{s}=\frac{1}{2}$ and $m_{s}=-\frac{1}{2}$. This fact is the corner stone of the periodic system of elements.

As there is no analogous requirement for bosons, the different statistics are obeyed by the ensembles of the identical fermions and bosons. This can be easily understood even on the simplest systems. E.g., we can consider the system of two particles in the states $a$ and $b$. Classically, there are four states of the pair: $a_{1} a_{2}, a_{1} b_{2} b_{1} a_{1}$ and $b_{1} b_{2}$ (the subscript enumerates the particles). So, assuming equal probabilities of these states, one can find both particles in the state $a$ with probability $1 / 4$, the same for the state $b$, and the probability to find different states of the particles is $1 / 2$. However, in the quantum mechanical treatment, these states are combined to the three symmetric states $|a a\rangle,|a b\rangle+|b a\rangle,|b b\rangle$ and one antisymmetric state $|a b\rangle-|b a\rangle$. Hence, in the case of bosons, the probability is $1 / 3$ to find both the particles in the state $|a\rangle$, same for $|b\rangle$, and same for the particles in different states. Finally, one finds the fermions in the different states with certainty. While for the fermions the coincident states are completely forbidden, for
bosons these are even more probable then for the classical particles. In fact, the total energy of non-interacting bosons is sum of all the energies of the occupied single particle states; the ground state of such a system is obviously one with all the particles in the same single particle state of the minimal energy. This effect is called Bose condensation. It is important important in understanding of various phenomena, and probably its most striking manifestation is superfluidity and superconductivity.

The same example may illustrate another striking difference between two types of the particles. The antisymmetrized vector (5.11) is correlated for any $\boldsymbol{n}$, while when all bosons are in the same state $|j\rangle$, i.e. when $n_{i}=N \delta_{i j}$, state $\left|\boldsymbol{n}^{+}\right\rangle$is non-correlated. In other words, fermions are always correlated: Pauli principle imposes, through the antisymmetrization, at least the exchange or Pauli correlation. Of course, general physical states are nontrivial superposition of the basis vectors $\left|\boldsymbol{n}^{-}\right\rangle$, being in a sense more correlated than $\left|\boldsymbol{n}^{-}\right\rangle$. In the same sense, Pauli correlation is the minimal one, and being not related to dynamics, it is purely kinematical. On the other hand, depending on the particular dynamics (including external fields and interaction between particles), the non-correlated states may be even preferred for bosons.

Despite these, and many other physical differences between bosons and fermions, many related expressions are formally analogous, and to a far extent it is possible to treat them in a unified way. Basically this is due to the unified form

$$
\begin{equation*}
\left|\boldsymbol{n}^{ \pm}\right\rangle=\sqrt{\frac{1}{\boldsymbol{n}!N!}} \sum_{\pi}( \pm)^{\tilde{\pi}} \Delta(\pi)\left|i_{1} \leq \cdots \leq i_{N}\right\rangle \tag{5.13}
\end{equation*}
$$

of the occupation number basis vectors, where for fermions $\boldsymbol{n}!=1$ is taken into account for fermions to get (5.11), and otherwise (5.9) is obtained.

For further reasons we introduce convention that if any occupation number is negative the corresponding vector vanishes: $n_{i}<0$ implies $\left|n^{ \pm}\right\rangle=0$. Also, we use $\boldsymbol{n}<\boldsymbol{n}^{\prime}$ to denote that $n_{s} \leq n_{s}^{\prime}$ for each $s$, i.e. that the single particle states involved in $\left|\boldsymbol{n}^{ \pm}\right\rangle$are subset of those appearing in $\left|\boldsymbol{n}^{\prime \pm}\right\rangle$ (this tacitly assumes different number of particles described by $\left|\boldsymbol{n}^{ \pm}\right\rangle$and $\left|\boldsymbol{n}^{\prime \pm}\right\rangle$ ).

As for the fermions, there is a well known convenient way to perform the antisymmetrization on the left of (5.13). In fact, this vector is obtained automatically if one calculates Slater determinant $N!^{-1 / 2}\left(\left|i_{p}\right\rangle_{q}\right)$, using the direct product as the multiplication of the matrix elements, which also means that the factors in the product are not commutative. The expansion rules for determinants provide the correct signs of the factors. However, if the signs are always taken to be positive, the definition of permanent is obtained, allowing one to write (5.13) in the form:

$$
\left|\boldsymbol{n}^{ \pm}\right\rangle=\sqrt{\frac{1}{\boldsymbol{n}!N!}}\left|\begin{array}{cccc}
\left|i_{1}\right\rangle & \left|i_{1}\right\rangle & \cdots & \left|i_{1}\right\rangle  \tag{5.14}\\
\left|i_{2}\right\rangle & \left|i_{2}\right\rangle & \cdots & \left|i_{2}\right\rangle \\
\cdots & \cdots & \cdots & \cdots \\
\left|i_{N}\right\rangle & \left|i_{N}\right\rangle & \cdots & \left|i_{N}\right\rangle
\end{array}\right|_{ \pm}
$$

Here, sign + stands for permanent and - for determinant. Alas, the algebra of permanents is much more cumbersome than that of determinants, making this notation widely used for fermions only. An illustration is easy proof (Exercise 5.2) that the $\left|n^{-}\right\rangle$is biuniquelly
related to the subspace $\mathcal{S}_{\boldsymbol{n}}=\operatorname{Span}\left(\left|i_{1}\right\rangle, \ldots,\left|i_{N}\right\rangle\right)$ spanned by the occupied single particle states (this subspace of $\mathcal{S}$ should be distinguished from $\mathcal{S}_{n}^{N}$ of the $N$-particle space). Further important application is the coordinate representation of the states $\left|\boldsymbol{n}^{ \pm}\right\rangle$. Indeed, with $\psi_{i}\left(\boldsymbol{r}_{p}\right)=\left\langle\boldsymbol{r}_{p} \mid i\right\rangle$, from (5.14) it is obvious that the corresponding wave function $\Psi_{\boldsymbol{n}^{ \pm}}\left(\boldsymbol{r}_{1}, \ldots, \boldsymbol{r}_{N}\right)=\left\langle\boldsymbol{r}_{1}, \ldots, \boldsymbol{r}_{N} \mid \boldsymbol{n}^{ \pm}\right\rangle$is:

$$
\Psi_{\boldsymbol{n}^{ \pm}}\left(\boldsymbol{r}_{1}, \ldots, \boldsymbol{r}_{N}\right)=\sqrt{\frac{1}{\boldsymbol{n}!N!}}\left|\begin{array}{cccc}
\psi_{i_{1}}\left(\boldsymbol{r}_{1}\right) & \psi_{i_{1}}\left(\boldsymbol{r}_{2}\right) & \cdots & \psi_{i_{1}}\left(\boldsymbol{r}_{N}\right)  \tag{5.15}\\
\psi_{i_{2}}\left(\boldsymbol{r}_{1}\right) & \psi_{i_{2}}\left(\boldsymbol{r}_{2}\right) & \cdots & \psi_{i_{2}}\left(\boldsymbol{r}_{N}\right) \\
\cdots & \cdots & \cdots & \cdots \\
\psi_{i_{N}}\left(\boldsymbol{r}_{1}\right) & \psi_{i_{N}}\left(\boldsymbol{r}_{2}\right) & \cdots & \psi_{i_{N}}\left(\boldsymbol{r}_{N}\right)
\end{array}\right|_{ \pm} .
$$

Exercise 5.2: Show that arbitrary basis $\left|j_{p}\right\rangle$ of $\left.\operatorname{Span}\left(\left|i_{p}\right\rangle|p=1, \ldots,| i_{N}\right\rangle\right)$ gives by (5.13) the same physical state as the basis $|i\rangle$ (i.e. the vector differing from $\left|\boldsymbol{n}^{-}\right\rangle$at most by a phase).

Being the basis of the state space $\mathcal{S}_{ \pm}^{N}$, the states $\left|\boldsymbol{n}^{ \pm}\right\rangle$do not exhaust it. The superposition principle inevitably includes their linear combinations. Although some of these linear combinations are again purely kinematically correlated states, made with the different bases of the single particle space, there are (anti)symmetrized states which cannot be written in the form (5.13) for any possible choice of the single particle states. Usually the eigenstates of the real N -particle systems are just such states, and then the additional correlation in comparison to the exchange one is naturally called dynamical correlation.

## Mixed States

The methods of the statistical physics are quite important in many body problems. They are based on the mixed states, and we consider some of their elementary properties. A statistical operator in $\mathcal{S}_{ \pm}^{N}$ can be expanded in the form

$$
\begin{equation*}
\rho=\sum_{\boldsymbol{n}, \boldsymbol{n}^{\prime}} c_{\boldsymbol{n} \boldsymbol{n}^{\prime}}\left|\boldsymbol{n}^{ \pm}\right\rangle\left\langle\boldsymbol{n}^{\prime \pm}\right| . \tag{5.16}
\end{equation*}
$$

From this form it is obvious that these operators commute with permutations. Indeed, $\Delta^{-1}(\pi) \rho \Delta(\pi)=\rho$, both for bosons and fermions since the parity of $\pi$ appears twice. Such operators are called symmetric operators, and only these are relevant (by the same argument as for $\rho$ ) for the systems of identical particles.

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## Subsystems

The reduced states of the clusters of $L \leq N$ particles are suitable to describe measurements on these natural subsystems. Applying previously introduced techniques (Subsection 1.3.4) to the subsystems of $L$ and the remaining $N-L$ particles, the reduced state of the particles $p_{1}, \ldots, p_{L}$ is the partial trace over other particles:

$$
\begin{equation*}
\rho_{p_{1} \ldots p_{L}}=\operatorname{Tr}_{\hat{p}_{1} \cdots \hat{p}_{L}} \rho . \tag{5.17}
\end{equation*}
$$

Here, hat denotes the particles omitted in the trace. That this is a statistical operator in $\mathcal{S}^{L}$ follows from the general theory. However, it should be checked that it is really
a state of $L$ identical particles, i.e. a statistical operator in $\mathcal{S}_{ \pm}^{L}$ having the form (5.16) with $L$-particle symmetrized states $\left|\boldsymbol{n}_{L}^{ \pm}\right\rangle$and $\left|\boldsymbol{n}_{L}^{\prime \pm}\right\rangle\left(\sum_{i} n_{L i}=\sum_{i} n_{L i}^{\prime}=L\right)$. To this end we use the fact (see Exercise 5.3) that the partial scalar product $\left\langle i_{L+1}, \ldots, i_{N} \mid \boldsymbol{n}^{ \pm}\right\rangle$is up to the constant just the (anti)symmetrized state of the first $L$ particles (if the set of the quantum numbers in the bra is not a subset of those in the ket, it vanishes). As in the partial trace appear only such terms, we see that the reduced state is really a state of $L$ identical particles.
Exercise 5.3: Prove that the partial scalar product of the $N$-particle state $\left|\boldsymbol{n}^{ \pm}\right\rangle$with non-correlated last $N-L$ particles vector $\left|j_{L+1}, \ldots, j_{N}\right\rangle$ with the occupation numbers $\boldsymbol{n}_{N-L}$ is $L$-particle state:

$$
\begin{equation*}
\left\langle j_{L+1}, \ldots, j_{N} \mid \boldsymbol{n}^{ \pm}\right\rangle=\sqrt{\frac{\boldsymbol{n}!L!}{\boldsymbol{n}_{L}!N!}}( \pm)^{\tilde{\sigma}}\left|\boldsymbol{n}_{L}^{ \pm}\right\rangle, \quad \boldsymbol{n}_{L}=\boldsymbol{n}-\boldsymbol{n}_{N-L} \tag{5.18a}
\end{equation*}
$$

(recall that vector vanishes if any $n_{L i}$ is negative); $\sigma$ is a permutation (unique in the case of fermions and irrelevant for bosons) such that

$$
\Delta(\sigma)\left|i_{1} \leq \cdots \leq i_{N}\right\rangle=\left|i_{\sigma^{-1}} \leq \cdots \leq i_{\sigma^{-1} L}, j_{L+1}, \ldots, j_{N}\right\rangle
$$

Analogously, for the first $L$-particle vector $\left|j_{1}, \ldots, j_{L}\right\rangle$ show

$$
\begin{equation*}
\left\langle j_{1}, \ldots, j_{L} \mid \boldsymbol{n}^{ \pm}\right\rangle=\sqrt{\frac{\boldsymbol{n}!(N-L)!}{\boldsymbol{n}_{N-L}!N!}}( \pm)^{\tilde{\sigma}}\left|\boldsymbol{n}_{N-L}^{ \pm}\right\rangle, \quad \boldsymbol{n}_{N-L}=\boldsymbol{n}-\boldsymbol{n}\left(j_{1}, \ldots, j_{L}\right) \tag{5.18b}
\end{equation*}
$$

with $\sigma$ defined by $\Delta(\sigma)\left|i_{1} \leq \cdots \leq i_{N}\right\rangle=\left|j_{1}, \ldots, j_{L}, i_{\sigma^{-1}(L+1)} \leq \cdots \leq i_{\sigma^{-1} N}\right\rangle$. For $L=1$ derive:

$$
\begin{equation*}
\left\langle j \mid \boldsymbol{n}^{ \pm}\right\rangle=\sqrt{\frac{n_{j}}{N}}( \pm)^{l_{j}}\left|\boldsymbol{n}_{N-1}^{ \pm}\right\rangle, \quad n_{N-1, i}=n_{i}-\delta_{i j}, l_{j}=\sum_{i<j} n_{i} \tag{5.18c}
\end{equation*}
$$

Exercise 5.4: Show the equality (here $\binom{\boldsymbol{n}}{\boldsymbol{n}_{L}} \stackrel{\text { def }}{=} \boldsymbol{n}!/ \boldsymbol{n}_{L}!\boldsymbol{n}_{N-L}$ ! with $\boldsymbol{n}=\boldsymbol{n}_{L}+\boldsymbol{n}_{N-L}$ ):

$$
\begin{equation*}
\left\langle\boldsymbol{n}_{N-L}^{ \pm} \mid \boldsymbol{n}^{ \pm}\right\rangle=\sqrt{\binom{\boldsymbol{n}}{\boldsymbol{n}_{L}} /\binom{N}{L}}( \pm)^{\tilde{\lambda}}\left|\boldsymbol{n}_{L}^{ \pm}\right\rangle \tag{5.19}
\end{equation*}
$$

It is intuitively clear that due to the indistinguishability of particles, i.e. due to the symmetrization of the state $\rho$, each of the $L$-particle subsystem states should be described by essentially the same reduced statistical operator:

$$
\begin{equation*}
\rho_{1, \ldots, L}=\rho_{p_{1}, \ldots, p_{L}}, \quad \forall\left\{p_{1}, \ldots, p_{L}\right\} . \tag{5.20}
\end{equation*}
$$

To see this formally, we find the matrix elements of the reduced operator $\rho_{1, \ldots, L}$ :

$$
\begin{array}{r}
\left\langle i_{1}, \ldots, i_{L}\right| \rho_{1, \ldots, L}\left|i_{1}^{\prime}, \ldots, i_{L}^{\prime}\right\rangle= \\
\sum_{j_{L+1}, \ldots, j_{N}}\left\langle i_{1}, \ldots, i_{L}\right|\left\langle j_{L+1}, \ldots, j_{N}\right| \rho\left|i_{1}^{\prime}, \ldots, i_{L}^{\prime}\right\rangle\left|j_{L+1}, \ldots, j_{N}\right\rangle
\end{array}
$$

As $\rho$ is symmetric, it can be replaced by $\Delta^{-1}(\pi) \rho \Delta(\pi)$ for any $\pi$. We chose $\pi$ to map $\{1, \ldots, L\}$ into $\left\{p_{1}, \ldots, p_{L}\right\}$. Acting on the bras and kets, this permutation changes the particles omitted in the partial trace to $p_{1}, \ldots, p_{L}$, giving $\rho_{p_{1}, \ldots, p_{L}}$. On the other hand,
the commutation of $\rho$ with the permutations of the first $L$ particles, shows that $\rho_{1, \ldots, L}$ is symmetric.

It has been discussed that the pure states are among the mixed ones singled out as the one dimensional projectors. As the purely kinematically correlated states $\left|\boldsymbol{n}^{ \pm}\right\rangle$span the state space, it is useful to find the reduced $L$-particle states of $\rho=\left|\boldsymbol{n}^{ \pm}\right\rangle\left\langle\boldsymbol{n}^{ \pm}\right|$. Using (5.18), one straightforwardly shows (Exercise 5.5):

$$
\begin{equation*}
\rho_{1, \ldots, L}=\binom{N}{L}^{-1} \sum_{\boldsymbol{n}_{L} \leq \boldsymbol{n}}\binom{\boldsymbol{n}}{\boldsymbol{n}_{L}}\left|\boldsymbol{n}_{L}^{ \pm}\right\rangle\left\langle\boldsymbol{n}_{L}^{ \pm}\right|, \quad \text { with }\binom{\boldsymbol{n}}{\boldsymbol{n}_{L}} \stackrel{\text { def }}{=} \prod_{s=1}^{|\mathcal{S}|}\binom{n_{s}}{n_{L s}} . \tag{5.21a}
\end{equation*}
$$

An important consequence is the form of the single particle reduced states $\rho_{1}$. The possible states $\left|\boldsymbol{n}_{1}\right\rangle$ are in fact the single particle states: for each $s$ with $n_{s}>0$ there is $\boldsymbol{n}_{1}$ such that $\boldsymbol{n}_{1 i}=\delta_{i s}$, making $\left|\boldsymbol{n}_{1}\right\rangle=|s\rangle$. Then $\boldsymbol{n}_{1}!=1$, and (5.21a) becomes:

$$
\begin{equation*}
\rho_{1}=N^{-1} \sum_{s=1}^{|\mathcal{S}|} n_{s}|s\rangle\langle s| . \tag{5.21b}
\end{equation*}
$$

Exercise 5.5: Prove (5.21a).
Particularly interesting, and for further purposes important, is the structure of these reduced states for fermions. As the occupation number $\boldsymbol{n}$ satisfies $n_{i}=0,1,(5.21 \mathrm{~b})$ becomes $\rho_{1}=P_{n} / N$, where $P_{\boldsymbol{n}}$ is the projector onto the previously introduced singleparticle subspace $\mathcal{S}_{n}=\operatorname{Span}\left(\left|i_{1}\right\rangle, \ldots,\left|i_{N}\right\rangle\right)$. Also, the general form of the reduced states is simple (Exercise 5.6):

$$
\begin{equation*}
\rho_{1, \ldots, L}=\frac{1}{\binom{N}{L}} \underbrace{P_{n} \otimes \cdots \otimes P_{n}}_{L} P_{L}^{(-)} \tag{5.22a}
\end{equation*}
$$

The most significant special cases of this result are for $L=1,2, N$ :

$$
\begin{equation*}
\rho_{1}=\frac{P_{\boldsymbol{n}}}{N}, \quad \rho_{12}=\frac{2}{N(N-1)}\left(P_{\boldsymbol{n}} \otimes P_{\boldsymbol{n}}\right) P_{2}^{(-)}, \quad \rho=\underbrace{P_{\boldsymbol{n}} \otimes \cdots \otimes P_{\boldsymbol{n}}}_{N} P^{(-)} . \tag{5.22b}
\end{equation*}
$$

The last expression explains how the subspace of the single particle occupied state uniquelly determines $N$-fermion state, as it is anticipated earlier (Exercise 5.2).

Exercise 5.6: Prove (5.22).

Therefore, each $N$-dimensional single-particle subspace completely determines one antisymmetric physical state; different choices of bases give Slater determinants differing by phase factors. Vice versa, each $N$ particle Slater determinant uniquely determines $N$ dimensional subspace in the single particle state $\mathcal{S}$. In this sense, Slater determinants and single particle subspaces are biuniquelly related.

### 5.1.4 Operators

## Symmetric Operators

Indistinguishability of the particles is reflected also on the operators. As the statistical operators, any $N$-particle observable must be invariant under the particle permutations, i.e. $\Delta^{-1}(\pi) A \Delta(\pi)=A$. Such symmetric operators in $\mathcal{S}^{N}$ are singled out as the range of the group $\boldsymbol{S}_{N}$ super-projector

$$
\begin{equation*}
\hat{\hat{P}}^{(+)} A=A, \quad \hat{\hat{P}}^{(+)} \stackrel{\text { def }}{=} \frac{1}{N!} \sum_{\pi} \Delta^{-1}(\pi) \ldots \Delta(\pi) . \tag{5.23}
\end{equation*}
$$

In other words, the physical observables commute with permutations, and therefore they transform according to the identical representation of $\boldsymbol{S}_{N}$.

Note that a symmetric operator acts within physical state spaces, i.e. it maps physical states into other ones, which is also a necessary physical property. Indeed, this is actually a requirement that $\mathcal{S}_{ \pm}^{N}$ is an invariant subspace for any physically relevant $N$ particle operator $A$, or equivalently, that $A$ commutes with the symmetrizer and the antisymmetrizer,

$$
\begin{equation*}
\left[\left(\hat{\hat{P}}^{(+)} A\right), P^{( \pm)}\right]=0, \tag{5.24}
\end{equation*}
$$

which is easily verified (Exercise 5.7). Thus, although the symmetric operators are frequently written in $\mathcal{S}^{N}$, their physically relevant part is the reduced operator $A P^{( \pm)}=$ $P^{( \pm)} A P^{( \pm)}$in $\mathcal{S}_{ \pm}^{N}$, having the form (5.16). Hence, only symmetric operators will be further considered.

Exercise 5.7: Prove (5.24).
The symmetric $N$-particle operators are usually constructed from one- or two-particle ones. More generally, an $L$-particle operator $A_{p_{1} \cdots p_{L}}$ essentially couples $L$ particles $p_{1}, \ldots, p_{L}$ : it acts nontrivially in the product of the spaces of these particles, while it is identity in other particle spaces. Using suitable permutations, one straightforwardly obtains equivalent operators for the other $L$ particle clusters: $\Delta^{-1}(\pi) A_{p_{1} \cdots p_{L}} \Delta(\pi)=A_{p_{\pi 1} \cdots p_{\pi L}}$. In this way, by the symmetrizer (5.23) the symmetric (in $N$ particle permutations) $L$-particle operator $A^{[L]}=c \hat{\hat{P}}^{(+)} A_{p_{1} \cdots p_{L}}$ is constructed, where the constant $c$ is defined according to the particular physical content of the operator (there is no universal normalization convention as for the states). When $A_{p_{1} \cdots p_{L}}$ is symmetric itself, i.e. when it commutes with $L$ particle permutations one gets:

$$
\begin{equation*}
A^{[L]}=c \sum_{p_{1}<\cdots<p_{L}}^{N} A_{p_{1}, \ldots, p_{L}} \tag{5.25}
\end{equation*}
$$

Otherwise the sum includes various orderings of the same particles; these essentially give the symmetrized (by the super-symmetrizer (5.23)) operator. Thus, any symmetric operator $A$ in $\mathcal{S}^{N}$ may be expanded into the series $A=\sum_{L=0}^{N} A^{[L]}$.

Typical examples of the symmetric operators are additive quantities like energy (then $c=1$ in (5.25)): an one-particle operator is just the sum $A^{[1]}=\sum_{p} A_{p}$, while $A^{[2]}=$
$\sum_{p_{1}<p_{2}} A_{p_{1} p_{2}}$ is a two particle operator. Another type are the multiplicative operators, like symmetry transformations: if $D_{1}$ is a single particle operator, then its $L$-th tensor power $D_{1} \otimes \cdots \otimes D_{1}$ is a symmetric operator in $\mathcal{S}^{L}$. Also, each $L$-particle (anti)symmetrizer $P_{L}^{ \pm}$is a symmetric operator.

As the indistinguishability of the particles is now completely incorporated in the quantum formalism, few examples of the symmetric operators discussed below may indicate its profound and far pervading consequences.

It has been mentioned that the Pauli principle allows at most two electrons to be in the same orbital state $|\psi\rangle$; then their total single-particle states are $\left|\psi, m_{s}=\frac{1}{2}\right\rangle$ and $\left|\psi, m_{s}=-\frac{1}{2}\right\rangle$, yielding the two particle state (in shorten notation) $\frac{1}{\sqrt{2}}|\psi\rangle(|+-\rangle-|-+\rangle)$. Note that the total spin $S^{2}=s_{1}^{2}+s_{2}^{2}$ of this state is $S=0$. In fact, this is the simplest manifestation of the quite general result on the total angular momentum of the closed shell fermions (Exercise 5.8): if $2 k+1$ fermions occupy the shell of the angular momentum $k$ (this is the set of states $|Q ; k m\rangle, m=-k, \ldots, k$, with fixed all other quantum numbers $Q$, including $k$ in particular), then their total angular momentum vanishes: $K=0$. This refers to any angular momentum appearing in the concrete problem. Here we slightly generalize the notion of the shell introduced in the physics of atoms in analogy to the hydrogen like model: besides the main quantum number $n$ there appears the orbital angular momentum $l$, as well as the spin $s$. Then, a closed shell is formed by the $2(2 l+1)$ electrons filling the states with the fixed $n$ and $l$. It follows that all the total angular momenta, $L, S$ and $J$ of these electrons vanish.

Exercise 5.8: Prove that the total angular momentum of the fermions in a closed shell is zero.
Symmetry transformations may serve as examples of symmetric multiplicative operators. To begin with, recall that an arbitrary single particle operator $D_{1}$ generates the multiplicative $L$-particle operators $\left[D^{\otimes L}\right]^{ \pm}=\hat{\hat{P}}^{(+)} \otimes_{p=1}^{L} D_{p}$ are (anti)symmetrized tensor powers of $D_{1}$. Recall that if there are $|\mathcal{S}|$ identical fermions, their total state space $\mathcal{S}_{-}^{N}$ is one-dimensional. This also means that $\left[D^{\otimes|\mathcal{S}|}\right]^{-}$is an one-dimensional operator; actually, this is the determinant of $D$. If $D(g)$ are operators forming a representation of a group $\boldsymbol{G}$, then the (anti)symmetrized operators $\left[D^{\otimes|\mathcal{S}|}\right]^{ \pm}$are also representations ((anti)symmetrized powers). In particular, in the space of the $|\mathcal{S}|$ fermions, this representation is $\operatorname{det} D(g)$. Therefore, the general conclusion is that if in a single particle space there is a representation $D(\boldsymbol{G})$, in the fulfilled case (of $|D|=|\mathcal{S}|$ fermions) acts the representation $\operatorname{det} D(\boldsymbol{G})$. Note that the previous example is a special case of this statement. Considering the rotational group (precisely $\mathrm{SU}(2)$ ), and its irreducible representations $D^{(k)}(S U(2)$ ), in the closed shell the representation is $\operatorname{det} D^{(k)}(S U(2))$. However, the determinant is the basis independent product of the eigenvalues $\mathrm{e}^{\mathrm{i} m \phi}(m=-k, \ldots, k)$, which immediately shows that the determinants of any irreducible representation of the rotational group form the unit representation $D^{(0)}(S U(2))$, i.e. with the zero total angular momentum. A generalization of this property is elaborated in the Exercise 5.9 (in fact, the last equation of the Exercises concerns the reducible representations in the single-fermion space, corresponding to several fulfilled shells). The last conclusion also reflects the obvious equality $\otimes_{p=1}^{n} D_{p}=\mathrm{e}^{\sum_{p=1}^{n} A_{p}}$ relating some $n$-particle and single-particle symmetric operators.
Exercise 5.9: Show:

1. $\left[D^{\otimes|D|}\right]^{-}=\operatorname{det} D$.
2. If $D=D_{1} \oplus D_{2} \oplus D_{K}$, then $\left[D^{\otimes|D|}\right]^{-}=\prod_{i=1}^{K}\left(\operatorname{det} D_{i}\right)^{|D|-\left|D_{i}\right|}$.

## Mean values and reduced operators

In analogy to the reduced $L$-particle statistical operators, the $L$-particle operators in $\mathcal{S}_{p_{1}} \otimes \cdots \otimes \mathcal{S}_{p_{L}}$ are from the $N$-particle ones constructed by partial traces over the remaining particles: $A_{p_{1}, \ldots, p_{L}} \stackrel{\text { def }}{=} \operatorname{Tr}_{\hat{p}_{1}, \ldots, \hat{p}_{L}} A$. As (5.20) is implied by the fact that $\rho$ is symmetric, for all the physically relevant (i.e. symmetric) operators one gets symmetric $L$-particle operators:

$$
\begin{equation*}
A_{1, \ldots, L}=A_{p_{1}, \ldots, p_{L}}, \quad \forall\left\{p_{1}, \ldots, p_{L}\right\} \tag{5.26}
\end{equation*}
$$

Mean value of the $L$-particle symmetric operator $A^{[L]}$ in the mixed state $\rho$ is

$$
\begin{equation*}
\left\langle A^{[L]}\right\rangle_{\rho}=\operatorname{Tr} A^{[L]} \rho=\binom{N}{L} \operatorname{Tr}_{1, \ldots, L} A_{1, \ldots, L} \rho_{1, \ldots, L} \tag{5.27}
\end{equation*}
$$

To see this it suffices to use (5.25): $\operatorname{Tr} A^{[L]} \rho=\sum_{p_{1}<\cdots<p_{L}}^{N} \operatorname{Tr}_{p_{1}, \ldots, p_{L}} A_{p_{1}, \ldots, p_{L}} \operatorname{Tr}_{\hat{p}_{1}, \ldots, \hat{p}_{L}} \rho$. According to (5.26), all the terms are same giving the binomial factor. Note that $\operatorname{Tr}_{1, \ldots, L}$ is the complete trace of the expression after it, and therefore in the applications it will be denoted only by Tr.

Particularly, the energy is additive quantity, and assuming only pairwise interactions the hamiltonian of $N$-particle system is $H=H^{[1]}+H^{[2]}$. Single-particle hamiltonian is $H^{[1]}=\sum_{p=1}^{N} H_{p}$, where $H_{p}=T_{p}+U_{p}$. Analogously, pairwise interactions $V_{p q}$, contribute in the sense of (5.25) $(c=1)$, by $H^{[2]}=\frac{1}{2} \sum_{p_{1} \neq p_{2}} V_{p_{1} p_{2}}$. Therefore, the mean value of the hamiltonian is found by (5.27):

$$
\begin{equation*}
E[\rho]=N \operatorname{Tr}_{1} H_{1} \rho_{1}+\binom{N}{2} \operatorname{Tr}_{12} V_{12} \rho_{12} \tag{5.28}
\end{equation*}
$$

Suppose that the occupation numbers are defined for the eigenbasis of $H_{1}$ in $\mathcal{S}$, i.e. $H_{1}=\sum_{i} \epsilon_{i}|i\rangle\langle i|$. Then for the purely kinematically correlated states $\rho=\left|\boldsymbol{n}^{ \pm}\right\rangle\left\langle\boldsymbol{n}^{ \pm}\right|$, one gets by (5.21a):

$$
\begin{equation*}
E[\rho]=\sum_{s} n_{s} \epsilon_{s}+\sum_{\boldsymbol{n}_{2}<\boldsymbol{n}}\binom{\boldsymbol{n}}{\boldsymbol{n}_{2}} \operatorname{Tr} V_{12}\left|\boldsymbol{n}_{2}^{ \pm}\right\rangle\left\langle\boldsymbol{n}_{2}^{ \pm}\right| \tag{5.29}
\end{equation*}
$$

The first term is the energy of the noninteracting particles, and in some simple dynamical approximation only it is used. The second term has not the form of the single-particle mean, which makes the problem too hard to proceed exactly. The main obstacle is correlation. In fact, even for the single particle term, the general expression is found only for the minimally correlated states. Further, to understand that the main complexity is introduced by correlations, we consider very exceptional case of non-correlated twoparticle reduced state: $\rho_{12}=\rho_{1}^{\prime} \otimes \rho_{1}^{\prime}$ (the factors must be equal as the state is symmetrical). Then the second term in (5.28) is easy to find: $\operatorname{Tr}_{12} V_{12} \rho_{12}=\operatorname{Tr} \rho_{1}^{\prime} \operatorname{Tr}_{2} V_{12} \rho_{1}^{\prime}$. We reveal the
mean field approximation: the interaction term is the energy of the first particle in the averaged field $\operatorname{Tr}_{2} V_{12} \rho_{1}^{\prime}$ of the others (being in the states $\rho_{1}^{\prime}$ ), which reduces the problem to the single-particle one. This observation motivates the most of the approximative approaches to the many-body dynamics.

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### 5.2 Second Quantization

Until now the systems with the fixed number of particles are considered. However, it is well known that in various processes particles are created or annihilated. Although nonrelativistic quantum mechanics cannot give real insight and explanation of such processes, its formalism is sufficiently general to treat such situation. Further, in many nonrelativistic problems, usually within some approximations, the dynamics can be well described in terms of quasi-particles: it turns out that excitations of the systems have so much of the properties of the real particles (e.g. momenta, spin, etc.), that it is not only technically convenient, but heuristically very fruitful, to treat them on equal footing. Typically this occurs in the harmonic approximations, retaining only the hamiltonian terms quadratic in some generalized coordinates. Being itself an archetypical example, the harmonic oscillator problem underlies all such approaches. Finally, even when the number of particles is fixed, the previously developed formalism is deficient in the sense that most of the calculations are performed in the space $\mathcal{S}^{N}$, with cumbersome appearance of the permutations, although the physical space is only its (anti)symmetrized part. All these motivate the development of the advanced quantum mechanical formalism, called second quantization, which is indispensable tool in the contemporary many-body physics and quantum field theory (including particularly solid state and elementary particles physics).

### 5.2.1 Fock space

Although nonrelativistic quantum mechanics cannot give real insight to the processes of the transitions between particles, e.g. decays or creation of particles, the introduced formalism may treat also these situations. To this end one should firstly realize that the state space $\mathcal{S}_{ \pm}^{N}$, describing the system with $N$ identical particles space state, is insufficient framework for the systems with the varying number of particles. Obviously, such spaces with all possible $N=0,1,2 \ldots$ are to be summed, to get the correct state space, called Fock space:

$$
\begin{equation*}
\mathcal{F}_{ \pm}=\oplus_{N} \mathcal{S}_{ \pm}^{N} \tag{5.30}
\end{equation*}
$$

Recall that if $\{|i\rangle \mid i=1, \ldots, \mathcal{S}\}$ is a basis in the single particle state, than the occupation number vectors $\boldsymbol{n}=\left(n_{1}, \ldots, n_{|\mathcal{S}|}\right)$ define uniquely states $\left|\boldsymbol{n}^{ \pm}\right\rangle$. All possible choices of $n_{i}$ such that $N=\sum_{i=1}^{|\mathcal{S}|} n_{i}$ give an orthonormal basis in $\mathcal{S}_{ \pm}^{N}$. Therefore, using the states $\left|\boldsymbol{n}^{ \pm}\right\rangle$
for $N=0,1, \ldots$ one gets the occupation number orthonormal basis of the Fock space $\mathcal{F}_{ \pm}$, and any vector of this space may be expanded as $|\Psi\rangle=\sum_{n} c_{\boldsymbol{n}}\left|\boldsymbol{n}^{ \pm}\right\rangle$. Particularly, for $N=0$, one gets the one dimensional vacuum space $\mathcal{S}_{ \pm}^{0}$, spanned by the vacuum state $|0\rangle$ (of course, $\boldsymbol{n}=0$ means that $n_{1}=\cdots=n_{|\mathcal{S}|}=0$ ). Also, for $N=1$ various occupation numbers are exhausted by one particle being in a particular state $|i\rangle=|0, \ldots, 0,1,0 \ldots, 0\rangle$ ( 1 on the $i^{\text {th }}$ place), and all other states are empty; thus the initial single particle state is spanned by this basis: $\mathcal{S}_{ \pm}^{1}=\mathcal{S}$.

### 5.2.2 Creation and annihilation operators

Obviously, the change in the number of particle is manifested as the transition from one of the subspaces $\mathcal{S}_{ \pm}^{N}$ to another one $\mathcal{S}_{ \pm}^{N^{\prime}}$. Such a transition may be considered as a series of changes by one particle. Thus, it suffices to study creation or annihilation of a single particle in some particular single-particle state.

Let in the $N$-particle system being in the state $\left|\boldsymbol{n}^{ \pm}\right\rangle=\left|n_{1}, n_{2}, \ldots\right\rangle$, the particle in the state $|i\rangle$ is annihilated. The resulting state is $|\boldsymbol{n} / \pm\rangle$, describing the system of $N-1$ particles, with $n_{j}^{\prime}=n_{j}$ particles in the states $|j(\neq i)\rangle$ and $n_{i}^{\prime}=n_{i}-1$ particles in the state $|i\rangle$. This observation is used to define conveniently the annihilation operator $a_{i}$ in the $i^{\text {th }}$ state:

$$
\begin{equation*}
a_{i}\left|n_{1}, \ldots, n_{i}, \ldots\right\rangle=( \pm)^{l_{i}(\boldsymbol{n})} \sqrt{n_{i}}\left|n_{1}, \ldots, n_{i}-1, \ldots\right\rangle . \tag{5.31}
\end{equation*}
$$

Here, $l_{i}(\boldsymbol{n})=\sum_{j=1}^{i-1} n_{j}$ is the number of particles being in the states $j$ with $j<i$ (for given $\left|\boldsymbol{n}^{ \pm}\right\rangle$), i.e. the number of states on the left of the first $i$. For bosons $(+)^{l_{i}(\boldsymbol{n})}=1$ and may be omitted, but we write it to emphasize the unified treatment. The adjoint operator $a_{i}^{\dagger}$ can be found considering the matrix elements $\left\langle n_{1}^{\prime}, n_{2}^{\prime}, \ldots\right| a_{i}^{\dagger}\left|n_{1}, n_{2}, \ldots\right\rangle$. As it acts on the left as the annihilation operator, all the matrix elements vanish except for $\left\langle n_{1}, \ldots n_{i}+1, \ldots\right|$, i.e.

$$
\left\langle\boldsymbol{n}^{\prime \pm}\right| a_{i}^{\dagger}\left|\boldsymbol{n}^{ \pm}\right\rangle=\sqrt{n_{i}^{\prime \pm}}( \pm)^{l_{i}\left(\boldsymbol{n}^{\prime \pm}\right)} \delta_{n_{1}}^{n_{1}^{\prime}} \cdots \delta_{n_{i}}^{n_{i}^{\prime}-1} \cdots=\sqrt{1+n_{i}}( \pm)^{l_{i}\left(\boldsymbol{n}^{\prime}\right)} \delta_{n_{1}}^{n_{1}^{n_{1}^{\prime}}} \cdots \delta_{1+n_{i}}^{n_{i}^{\prime}} \cdots
$$

Recall that for fermions $n_{i}=0,1$, meaning that for $n_{i}=1$ the occupation number on the left should be $n_{i}^{\prime}=2$, and bra automatically vanishes, implying that the whole scalar product is always zero; as the only other case is $n_{i}=0$ the root may be for the fermions changed to $\sqrt{1-n_{i}}$. Finally, one gets the creation operator:

$$
\begin{equation*}
a_{i}^{\dagger}\left|n_{1}, \ldots, n_{i}, \ldots\right\rangle=( \pm)^{l_{i}(\boldsymbol{n})} \sqrt{1 \pm n_{i}}\left|n_{1}, \ldots, n_{i}+1, \ldots\right\rangle . \tag{5.32}
\end{equation*}
$$

Exercise 5.10: Show that $a_{s}\left|\boldsymbol{n}^{ \pm}\right\rangle=\sqrt{N}\left\langle s \mid \boldsymbol{n}^{ \pm}\right\rangle$. Derive (5.32) using this. Generalize the first result to get (5.18b) in the form:

$$
\begin{equation*}
\left\langle j_{1}, \ldots, j_{L} \mid \boldsymbol{n}^{ \pm}\right\rangle=\sqrt{\frac{(N-L)!}{N!}} a_{j_{L}} \cdots a_{j_{1}}\left|\boldsymbol{n}^{ \pm}\right\rangle . \tag{5.33}
\end{equation*}
$$

Find the form of this expression independent on the number of particles $N$ in the state $\left|\boldsymbol{n}^{ \pm}\right\rangle$.
Some important properties of the introduced operators directly follow. For any $i$ in the vacuum state

$$
\begin{equation*}
a_{i}|0\rangle=0, \quad a_{i}^{\dagger}|0\rangle=|i\rangle . \tag{5.34}
\end{equation*}
$$

While for fermions $a_{i}^{2}=a_{i}^{\dagger^{2}}=0$, for bosons $a_{i}^{k}\left|\boldsymbol{n}^{+}\right\rangle=\sqrt{\frac{n_{i}!}{\left(n_{i}-k\right)!}}\left|n_{1}, \ldots, n_{i}-k, \ldots\right\rangle$ and $a_{i}^{\dagger^{k}}\left|\boldsymbol{n}^{+}\right\rangle=\sqrt{\frac{\left(n_{i}+k\right)!}{n_{i}!}}\left|n_{1}, \ldots, k, \ldots\right\rangle$. Obviously, any state can be obtained from vacuum by creation operators:

$$
\begin{equation*}
\left|\boldsymbol{n}^{ \pm}\right\rangle=\frac{1}{\sqrt{n_{1}!\cdots n_{|\mathcal{S}|}}} a_{1}^{\dagger_{1}^{n_{1}}} \cdots a_{|\mathcal{S}|}^{\dagger^{n_{|\mathcal{S}|}}}|0\rangle \tag{5.35}
\end{equation*}
$$

Finally, note that the number of particles in the state $|i\rangle$ is easily obtained by application of occupation number operator $\hat{n}_{i} \stackrel{\text { def }}{=} a_{i}^{\dagger} a_{i}$, as its eigenvalue:

$$
\begin{equation*}
\hat{n}_{i}\left|\boldsymbol{n}^{ \pm}\right\rangle=n_{i}\left|\boldsymbol{n}^{ \pm}\right\rangle \tag{5.36}
\end{equation*}
$$

(for bosons, this directly follows from the definitions, while for fermions in addition the cases of $n_{i}=0$ and $n_{i}=1$ should be separately discussed). Accordingly, since the states with the same number of particles are comprised in the subspaces $\mathcal{S}_{ \pm}^{N}$, these are eigenspaces of the number of particles operator $\hat{N}=\sum_{i} \hat{n}_{i}=\sum_{i} a_{i}^{\dagger} a_{i}$ :

$$
\begin{equation*}
\hat{N}\left|\boldsymbol{n}^{ \pm}\right\rangle=\sum_{i} n_{i}\left|\boldsymbol{n}^{ \pm}\right\rangle=N\left|\boldsymbol{n}^{ \pm}\right\rangle \tag{5.37}
\end{equation*}
$$

The creation and annihilation operators correspond to the arbitrary chosen basis in the single particle space $\mathcal{S}$. If the considered basis is continual, then one get continual family of the creation/annihilation operators, called field operators. For example, to the coordinate basis $|x\rangle$ correspond the field operators $a^{\dagger}(x)$ creating a particle at $x:|x\rangle=a^{\dagger}(x)|0\rangle$.

### 5.2.3 Bosonic and fermionic commutation relations

It is obvious that any transition from $\left|\boldsymbol{n}^{ \pm}\right\rangle$to $\left|\boldsymbol{n}^{\prime \pm}\right\rangle$ can be obtained by action of the appropriately chosen product of various creation and annihilation operators. Therefore, it is important to manipulate easily with these operators, and the commutation relations among them enable this. In addition, the commutation relations are conceptually important, revealing in a specific way the difference between bosons and fermions, with the far reaching consequences.

At first, we look for the commutation relations of two annihilation operators, $a_{i}$ and $a_{j}$. To this end we compare the action of $a_{i} a_{j}$ and $a_{j} a_{i}$, for $i<j$ :

$$
\begin{gathered}
a_{i} a_{j}\left|\boldsymbol{n}^{ \pm}\right\rangle=( \pm)^{l_{i}(\boldsymbol{n})+l_{j}(\boldsymbol{n})} \sqrt{n_{i} n_{j}}\left|\ldots, n_{i}-1, \ldots, n_{j}-1, \ldots\right\rangle, \\
a_{j} a_{i}\left|\boldsymbol{n}^{ \pm}\right\rangle=( \pm)^{l_{i}(\boldsymbol{n})+l_{j}(\boldsymbol{n})-1} \sqrt{n_{i} n_{j}}\left|\ldots, n_{i}-1, \ldots, n_{j}-1, \ldots\right\rangle .
\end{gathered}
$$

In the bosonic case both results are equal. However, since in the second equation $a_{i}$ acts first, it diminishes by one the number of particles in the states with indices less than $j$, which is in the case of fermions manifested as the opposite sign of the result. For $i=j$, one trivially gets two equal results for bosons and zeros for fermions. Finally, for $j<i$ one should reverse the order of the equations. Hence, denoting by $[A, B]_{ \pm} \stackrel{\text { def }}{=} A B \mp B A$ the commutator and anticommutator of the operators, one gets $\left[a_{i}, a_{j}\right]_{ \pm}=0$. Quite analogously, the same results is obtained for the pair of annihilation operators.

Similarly, if one operator is annihilation and the other one creation operator the same result follows from $(i<j)$ :

$$
\begin{gathered}
a_{i} a_{j}^{\dagger}\left|\boldsymbol{n}^{ \pm}\right\rangle=( \pm)^{l_{i}(\boldsymbol{n})+l_{j}(\boldsymbol{n})} \sqrt{n_{i}\left(1 \pm n_{j}\right)}\left|\ldots, n_{i}-1, \ldots, n_{j}+1, \ldots\right\rangle, \\
a_{j}^{\dagger} a_{i}\left|\boldsymbol{n}^{ \pm}\right\rangle=( \pm)^{l_{i}(\boldsymbol{n})+l_{j}(\boldsymbol{n})-1} \sqrt{n_{i}\left(1 \pm n_{j}\right)}\left|\ldots, n_{i}-1, \ldots, n_{j}+1, \ldots\right\rangle .
\end{gathered}
$$

However, for $i=j$ there is important difference. In fact,

$$
a_{i} a_{i}^{\dagger}\left|\boldsymbol{n}^{ \pm}\right\rangle=( \pm)^{l_{i}(\boldsymbol{n})+l_{i}(\boldsymbol{n})} \sqrt{\left(1+n_{i}\right)\left(1 \pm n_{i}\right)}\left|\ldots, n_{i}, \ldots\right\rangle=\left(1 \pm n_{i}\right)\left|\boldsymbol{n}^{ \pm}\right\rangle
$$

to show this for fermions, one again notes that the whole expression a priori vanishes for $n_{i}=1$, so the change of the factor $1+n_{i}$ to $1-n_{i}$ does not matter, as well as in the case $n_{i}=0$. In the opposite order one gets the occupation number operators, i.e. $a_{i}^{\dagger} a_{i}\left|\boldsymbol{n}^{ \pm}\right\rangle=n_{i}\left|\boldsymbol{n}^{ \pm}\right\rangle$. So, all the obtained results are comprised in:

$$
\begin{equation*}
\left[a_{i}, a_{j}\right]_{ \pm}=\left[a_{i}^{\dagger}, a_{j}^{\dagger}\right]_{ \pm}=0, \quad\left[a_{i}, a_{j}^{\dagger}\right]_{ \pm}=\delta_{i j} . \tag{5.38}
\end{equation*}
$$

Note that for the field operators, the Dirac's delta function is on the right.
Exercise 5.11: Reconsider the harmonic oscillator problem in terms of the second quantization: find single particle space and Fock space. What are particles here? Are they fermions or bosons?

### 5.2.4 Representation of the second quantization

Any orthonormal basis $|i\rangle$ in the single-particle state $\mathcal{S}$ can be use to construct the set of the creation operators $a_{i}^{\dagger}$ in $\mathcal{F}_{ \pm}$, and their adjoint annihilation operators $a_{i}$. Consequently, instead of the non-correlated bases in $\mathcal{F}$ based on the single-particle basis, the corresponding creation and annihilation operators may be used to represent states and operators directly in the physical spaces $\mathcal{F}_{ \pm}$.

As for the vectors, their second quantization form is obvious. Indeed, for arbitrary state $|\Psi\rangle=\sum_{n} \Psi_{n}\left|\boldsymbol{n}^{ \pm}\right\rangle$from $\mathcal{F}_{ \pm}$, (5.35) gives $|\Psi\rangle=\sum_{n} \Psi_{n} \frac{1}{\sqrt{n!}} a_{1}^{\dagger^{n_{1}}} \cdots a_{|\mathcal{S}|}^{\dagger^{n}|\mathcal{S}|}|0\rangle$. This may be applied also for the single particle states. For example, any wave function can be written as $|\psi\rangle=\int \psi(\boldsymbol{r}) a^{\dagger}(\boldsymbol{r}) \mathrm{d} \boldsymbol{r}|0\rangle=\psi^{\dagger}|0\rangle$, where $\psi^{\dagger}=\int \psi(x) a^{\dagger}(x) \mathrm{d} x$, and $a^{\dagger}(\boldsymbol{r})$ is coordinate field. Particularly, plane waves are: $|\boldsymbol{p}\rangle=\frac{1}{(2 \pi \hbar)^{3} / 2} \int \mathrm{e}^{\frac{\mathrm{i}}{\hbar} \boldsymbol{p} \cdot \boldsymbol{r}} a^{\dagger}(\boldsymbol{r}) \mathrm{d} \boldsymbol{r}|0\rangle$.

The operators also may be expressed in terms of the creation and annihilation operators, as follows from their Dirac notation. Obviously, the operators leaving the number of particles invariant are combined from the the product of the same number of creation and annihilation operators. The additive operators among them have very convenient form. We give it for one-particle and two-particle operators $T^{[1]}=\sum_{p} T_{p}$ and $V^{[2]}=$ $\frac{1}{2} \sum_{p_{1} \neq p_{2}} V_{p_{1} p_{2}}$, with $T_{1}=T=\sum_{s s^{\prime}} T_{s^{\prime}}^{s}|s\rangle\left\langle s^{\prime}\right|$ and $V_{12}=V=\sum_{s_{1} s_{1}^{\prime} s_{2} s_{2}^{\prime}} V_{s_{1}^{s} s_{2}^{\prime}}^{s_{1} s_{2}}\left|s_{1} s_{2}\right\rangle\left\langle s_{1}^{\prime} s_{2}^{\prime}\right|$ (this notation is used to remind of the typical examples of kinetic energy and pairwise interaction, but the result is quite general):

$$
\begin{align*}
T^{[1]} & =\sum_{s s^{\prime}} T_{s^{\prime}}^{s} a_{s}^{\dagger} a_{s^{\prime}},  \tag{5.39a}\\
V^{[2]} & =\frac{1}{2} \sum_{s_{1} s_{1}^{\prime} s_{2} s_{2}^{\prime}} V_{s_{1}^{\prime} s_{2}^{\prime}}^{s_{1} s_{2}} a_{s_{1}}^{\dagger} a_{s_{2}}^{\dagger} a_{s_{2}^{\prime}} a_{s_{1}^{\prime}} . \tag{5.39b}
\end{align*}
$$

Exercise 5.12: Prove that for additive symmetric $L$-particle operator $A^{[L]}=\sum_{p_{1}<\ldots<p_{l}} A_{p_{1}, \ldots, p_{L}}$ holds

$$
\begin{equation*}
A^{[L]}=\frac{1}{L!} \sum_{\substack{s_{1}, \ldots, s_{L} \\ s_{1}, \ldots, s_{L}^{L}}} A_{s_{1}, \ldots, s_{L}^{\prime}}^{s_{1}, \ldots, s_{L}} a_{s_{1}}^{\dagger} \cdots a_{s_{L}}^{\dagger} a_{s_{L}} \cdots a_{s_{1}}, \tag{5.39c}
\end{equation*}
$$

with $A_{s_{1}^{\prime}, \ldots, s_{L}}^{s_{1}, \ldots, s_{L}}=\left\langle s_{1}, \ldots, s_{L}\right| A_{1, \ldots, L}\left|s_{1}^{\prime}, \ldots, s_{L}^{\prime}\right\rangle$.
Exercise 5.13: Express the angular momentum operators in the form (5.39a) and give the alternative proof of the closed shell property (5.8).

In the context of representations, the transitions between them should be considered, too. Taking besides $|a ; i\rangle$ the basis $|b ; i\rangle$, related to the first one by the unitary transition operator $U,|b ; q\rangle=\sum_{i q} U_{i q}|a ; i\rangle$, one gets two sets of creation (and annihilation) operators $a_{i}^{\dagger}$ and $b_{i}^{\dagger}$. Then obviously

$$
\begin{equation*}
b_{q}^{\dagger}=\sum_{i q} U_{i q} a_{i}^{\dagger}, \quad b_{q}=\sum_{i q} U_{i q}^{*} a_{i}, \tag{5.40}
\end{equation*}
$$

meaning that the creation operators transform like basis, while the annihilation ones transform like dual basis. The plane waves illustrate how the transition between bases relates the corresponding creation operators (fields): $a^{\dagger}(\boldsymbol{p})=\frac{1}{(2 \pi \hbar)^{3} / 2} \int \mathrm{e}^{\frac{\mathrm{i}}{\hbar} \boldsymbol{p} \cdot \boldsymbol{r}} a^{\dagger}(\boldsymbol{r}) \mathrm{d} \boldsymbol{r}$.

## Chapter 6

## Approximate Methods

Apart from few simple (although important) problems, Schrödinger equation cannot be solved exactly. A number of methods is developed to find the approximate solutions to it. It should be emphasized that these methods are substantially different, and suited for particular purposes.

### 6.1 Time Independent Perturbations

### 6.1.1 Perturbed and Unperturbed Hamiltonian

When the Hamiltonian $H$ can be separated into two parts, $H=H_{0}+H^{\prime}$, the first of them being dominant, it is natural to assume that the solutions (eigenvalues and eigenvectors) of the dominant part $H^{0}$ are close to those of the total hamiltonian. This intuitive fact is easily formalized. To this end the parameter $a \in[0,1]$ is introduced, such that the $a$-parameterized operator family

$$
\begin{equation*}
H(a)=H_{0}+a H^{\prime} \tag{6.1}
\end{equation*}
$$

form a continuous path from $H_{0}=H(0)$ to $H=H(1)$. We assume that along the whole path hamiltonians $H(a)$ have discrete spectrum. If $E_{n}(0)$ and $|n, \lambda, 0\rangle$ are the eigenvalues and vectors of the unperturbed hamiltonian, $H_{0}$, then the parameter $a$, which can be as small as necessary, interrelates perturbed and unperturbed eigenvalues and eigenvectors: if $E_{n}(1)$ is some eigenvalue of $H$ (i.e. in $a=1$ ), then when $a$ continuously decreases to 0 , one gets path $E_{n}(a)$ of eigenvalue, with $E_{n}(0)$ being the corresponding eigenvalue of $H_{0}$. However, several different values, say $E_{n}(a), E_{n^{\prime}}(a), \ldots$, can have the same initial unperturbed eigenvalue $E_{n}(0)$ : a degenerate level of the unperturbed hamiltonian may be split by the perturbation into several less degenerate levels ${ }^{1}$. We denote by $E_{n i}(a)$ all the perturbed levels which come together to $E_{n}^{(0)}=E_{n i}(0)$ in the unperturbed case. Then the condition for the corresponding families of the eigenvalues $E_{n i}(a)$ and eigenvectors $|n, i, \lambda, a\rangle$ is:

$$
\begin{equation*}
H(a)|n, i, \lambda, a\rangle=E_{n i}(a)|n, i, \lambda, a\rangle . \tag{6.2}
\end{equation*}
$$

[^5]In other words, index $i$ counts the splitting of the unperturbed level $E_{n}(0)$. The reminded degeneracy (if any) of the perturbed levels is counted by the index $\lambda$, such that the perturbed eigenvectors are $|n i \lambda ; a\rangle$.

### 6.1.2 Perturbative expansion

These preliminary considerations enable us to expand in $a$ all the terms in the parametric eigen problem (6.2):

$$
\begin{equation*}
\left(H_{0}+a H^{\prime}\right) \sum_{p} a^{p}|n i \lambda\rangle^{(p)}=\left(\sum_{s} a^{s} E_{n i}^{(s)}\right)\left(\sum_{p} a^{p}|n i \lambda\rangle^{(p)}\right) . \tag{6.3}
\end{equation*}
$$

Here one should take into account that $E_{n i}^{(0)}=E_{n}^{(0)}$ and $|n i \lambda\rangle^{(0)}=|n i \lambda ; 0\rangle$. While the zero order energy levels $E_{n}^{(0)}$ are assumed to be known, the zero order vectors $|n i \lambda ; 0\rangle$ are not, since the unperturbed Hamiltonian $H_{0}$ cannot single out any basis within the same eigen space. Then, separating different orders in $a$ one gets the series of equations:

$$
\begin{align*}
a^{0}: & H_{0}|n i \lambda ; 0\rangle=E_{n}^{(0)}|n i \lambda ; 0\rangle ;  \tag{6.4a}\\
a^{1}: & H_{0}|n i \lambda\rangle^{(1)}+H^{\prime}|n i \lambda ; 0\rangle=E_{n}^{(0)}|n i \lambda\rangle^{(1)}+E_{n i}^{(1)}|n i \lambda ; 0\rangle ;  \tag{6.4b}\\
\ldots & \cdots \ldots \ldots \ldots \ldots \ldots \ldots \ldots  \tag{6.4c}\\
a^{p}: & H_{0}|n i \lambda\rangle^{(p)}+H^{\prime}|n i \lambda\rangle^{(p-1)}=\sum_{s=0}^{p} E_{n i}^{(s)}|n i \lambda\rangle^{(p-s)} .
\end{align*}
$$

The first of these equations is immediate consequence of the assumption that the zero order eigen problem of $H$ is the eigen problem of $H_{0}$. In this sense, this trivial equation is a consistency condition for the method. Next we come to the first order equation. If $P_{n}$ is the projector onto the unperturbed hamiltonian eigenspace $\mathcal{S}_{n}$, then $P_{n}|n i \lambda ; 0\rangle=|n i \lambda ; 0\rangle$ and $P_{n} H_{0}=P_{n} H_{0} P_{n}=E_{n}^{(0)} P_{n}$. When (6.4b) are projected into $\mathcal{S}_{n}$, i.e. its both sides multiplied by $P_{n}$, one gets:

$$
E_{n}^{(0)} P_{n}|n i \lambda\rangle^{(1)}+P_{n} H^{\prime} P_{n}|n i \lambda ; 0\rangle=E_{n}^{(0)} P_{n}|n i \lambda\rangle^{(1)}+E_{n i}^{(1)} P_{n}|n i \lambda ; 0\rangle
$$

Cancelling the first term on the both sides, this becomes the eigenequation for the operator $P_{n} H^{\prime} P_{n}$, which is the perturbation $H^{\prime}$ cut in the subspace $\mathcal{S}_{n}$ :

$$
\begin{equation*}
P_{n} H^{\prime} P_{n}|n i \lambda ; 0\rangle=E_{n i}^{(1)}|n i \lambda ; 0\rangle . \tag{6.5}
\end{equation*}
$$

Both the first order correction of energy levels and the zero order state vectors are obtained by solving this eigenproblem. Note further that each unperturbed level is up to the first order correction treated independently. In practice this means that one takes some important levels (usually the ground one, and sometimes a few above it). Taking any eigenbasis $|n \mu ; 0\rangle$ for each of these levels $E_{n}^{(0)}$, the secular equation (6.5) is obtained in the matrix form $\langle n \mu ; 0| H^{\prime}|n \nu ; 0\rangle=E_{n}^{(1)} \delta_{\mu \nu}$. Its different eigenvalues $E_{n i}^{(1)}$ are the first order corrections in energy, while the corresponding eigenvectors $|n i, \lambda ; 0\rangle$ are the zero
order eigenvectors (for the perturbation $H^{\prime}$ ) obtained as the representative columns in the chosen basis $|n \mu ; 0\rangle$.

This procedure can be further developed for higher order corrections. We will analyze only the case of non-degenerate unperturbed level (the complete theory is given by Kato [?]). Let us mention here that when $E_{n}^{(0)}$ is a discrete eigenvalue with very close other energies, or a point of the continual spectrum are to be corrected, one should solve equation (6.5), but with $P_{n}$ being the sum of projectors of the relevant neighboring eigenvalues (e.g. from some interval centered at $E_{n}^{(0)}$ ).

It is very convenient to have a simple estimate of the error of the involved when energy level is approximated. In fact, the mean value of the hamiltonian terms which are excluded in the considered level of perturbation, taken in the obtained corrected eigenvectors gives such an estimate.

### 6.1.3 Higher Corrections for Non-degenerate Level

If $E_{n}^{(0)}$ is non-degenerate, the indices $i$ and $\lambda$ are redundant. While obviously the zero order state must be $|n ; 0\rangle$ (the only one in $\mathcal{S}_{n}$ ), (6.5) shows that the first order energy correction is simply the mean of the perturbation in this state:

$$
\begin{equation*}
\langle n ; 0| H^{\prime}|n ; 0\rangle=E_{n}^{(1)} . \tag{6.6}
\end{equation*}
$$

On the other hand, the projector $P_{n}^{\perp}=\sum_{m(\neq n), \mu}|m \mu ; 0\rangle\langle m \mu ; 0|$ onto the orthocomplement to $\mathcal{S}_{n}$ is (we assumed non-degeneracy for $n^{\text {th }}$ unperturbed level only, while the other ones may be degenerate) is applied to the first order equation, giving the projection onto $\mathcal{S}_{n}^{\perp}$ :

$$
\sum_{m(\neq n), \mu}\left(E_{m}^{(0)}\langle m \mu ; 0 \mid n\rangle^{(1)}+\langle m \mu ; 0| H^{\prime}|n ; 0\rangle\right)|m \mu ; 0\rangle=E_{n}^{(0)} \sum_{m(\neq n), \mu}\langle m \mu ; 0 \mid n\rangle^{(1)}|m \mu ; 0\rangle .
$$

Consequently, finding the projections onto the basis vectors $|m \mu ; 0\rangle$, and assuming that the first order correction $|n\rangle^{(1)}$ of the state $|n ; 0\rangle$ is orthogonal to $|n ; 0\rangle$ :

$$
\begin{equation*}
\langle m \mu ; 0 \mid n\rangle^{(1)}=\frac{\langle m \mu ; 0| H^{\prime}|n ; 0\rangle}{E_{n}^{(0)}-E_{m}^{(0)}} \text {, i.e. }|n\rangle^{(1)}=\sum_{m(\neq n), \mu} \frac{\langle m \mu ; 0| H^{\prime}|n ; 0\rangle}{E_{n}^{(0)}-E_{m}^{(0)}}|m \mu ; 0\rangle \tag{6.7}
\end{equation*}
$$

The second order correction of the energy can be found analogously, projecting (6.4c) for $p=2$ onto $\mathcal{S}_{n}$ :

$$
E_{n}^{(0)} P_{n}|n\rangle^{(2)}+P_{n} H^{\prime}|n\rangle^{(1)}=E_{n}^{(0)} P_{n}|n\rangle^{(2)}+E_{n}^{(1)} P_{n}|n\rangle^{(1)}+E_{n}^{(2)}|n ; 0\rangle .
$$

The first terms on the both sides cancel, the second term on the right vanishes due to the orthogonality of $|n\rangle^{(1)}$ and $|n ; 0\rangle$ :

$$
|n ; 0\rangle\langle n ; 0| H^{\prime}|n\rangle^{(1)}=E_{n}^{(2)}|n ; 0\rangle .
$$

Finally, substituting (6.7) one gets

$$
\begin{equation*}
E_{n}^{(2)}=\sum_{m(\neq n), \mu} \frac{\left.\left|\langle m \mu ; 0| H^{\prime}\right| n ; 0\right\rangle\left.\right|^{2}}{E_{n}^{(0)}-E_{m}^{(0)}} \tag{6.8}
\end{equation*}
$$

Usually it is interesting to find the corrections to the energy of the ground state, $E_{0}^{(0)}$, being the minimal among the eigenvalues of $H_{0}$. Obviously, the second order correction is always negative, independently of the system. In fact, there is an estimate of this correction:

Theorem 6.1 If there is finite difference $\triangle E_{n}^{(0)}$ from the level $E_{n}^{(0)}$ to the closest to it level, and the dispersion $\triangle H^{\prime}$ in the state $|n ; 0\rangle$ is finite, then

$$
\begin{equation*}
\left|E^{(2)}\right| \leq \frac{\triangle^{2} H^{\prime}}{\left|\triangle E_{n}^{(0)}\right|} \tag{6.9}
\end{equation*}
$$

-Proof: Taking the absolute value of (6.8) one gets the proposed estimation directly:

$$
\begin{aligned}
\left|E_{n}^{(2)}\right| \leq & \sum_{m(\neq n), \mu} \frac{\left.\left|\langle m \mu ; 0| H^{\prime}\right| n ; 0\right\rangle\left.\right|^{2}}{\left|E_{n}^{(0)}-E_{m}^{(0)}\right|} \leq \frac{\left.\sum_{m(\neq n), \mu}\left|\langle m \mu ; 0| H^{\prime}\right| n ; 0\right\rangle\left.\right|^{2}}{\left|\Delta E_{n}^{(0)}\right|}= \\
& \frac{\langle n ; 0| H^{\prime}(I-|n ; 0\rangle\langle n ; 0|) H^{\prime}|n ; 0\rangle}{\left|\Delta E_{n}^{(0)}\right|}=\frac{\left\langle H^{\prime 2}\right\rangle-\left\langle H^{\prime}\right\rangle^{2}}{\left|\Delta E_{n}^{(0)}\right|} .
\end{aligned}
$$

### 6.2 Adiabatic approximation

Complex matter consists of ions and electrons. Their great difference in masses enables specific approximate separations of degrees of freedom related to the two classes. Let $\boldsymbol{R}=\left(R_{1 x}, \ldots, R_{N z}\right)$ denotes the set of all the coordinates of ions, while $\boldsymbol{r}=\left(r_{1 x}, \ldots, r_{N z}\right)$ are the coordinates of electrons. Total hamiltonian is:

$$
\begin{equation*}
H=T_{\boldsymbol{R}}+T_{\boldsymbol{r}}+V(\boldsymbol{R}, \boldsymbol{r})=H_{r}+T_{\boldsymbol{R}} \tag{6.10}
\end{equation*}
$$

where $T_{\boldsymbol{R}}=\sum_{I=1}^{N} \frac{\hbar^{2}}{2 M_{I}} \triangle_{\boldsymbol{R}}$, and $T_{\boldsymbol{r}}=\sum_{r=1}^{n} \frac{\hbar^{2}}{2 m} \triangle_{r}$ are ionic and electronic kinetic energies, while the total interaction $V(\boldsymbol{R}, \boldsymbol{r})=V_{\text {ion }}(\boldsymbol{R})+V_{\mathrm{el}}(\boldsymbol{r})+V_{\mathrm{ie}}(\boldsymbol{R}, \boldsymbol{r})$ is the total interaction consisting of ionic, electronic and electron-ion part. This hamiltonian gives total Schrödinger equation:

$$
\begin{equation*}
H \Psi(\boldsymbol{R}, \boldsymbol{r})=E \Psi(\boldsymbol{R}, \boldsymbol{r}) \tag{6.11}
\end{equation*}
$$

Electronic hamiltonian

$$
\begin{equation*}
H_{r}=T_{\boldsymbol{r}}+V(\boldsymbol{R}, \boldsymbol{r}), \tag{6.12}
\end{equation*}
$$

does not include ionic kinetic energy, and all other terms commute with ionic coordinate operators ${ }^{2}$. Therefore, it can be understood as a family of the electronic hamiltonians, parameterized by the ionic configurations $\boldsymbol{R}$. In this sense, the solutions of the electronic Schrödinger equation

$$
\begin{equation*}
H_{r} \eta_{i}(\boldsymbol{R}, \boldsymbol{r})=\epsilon_{i}(\boldsymbol{R}) \eta_{i}(\boldsymbol{R}, \boldsymbol{r}), \tag{6.13}
\end{equation*}
$$

[^6]are parameterized by $\boldsymbol{R}$. This means that for each ionic position $\boldsymbol{R}$ there are electronic states, making basis in the electronic state space, and energies. Accordingly, total eigenfunction can be expanded in the form
\[

$$
\begin{equation*}
\Psi(\boldsymbol{R}, \boldsymbol{r})=\sum_{i} \chi_{i}(\boldsymbol{R}) \eta_{i}(\boldsymbol{R}, \boldsymbol{r}) . \tag{6.14}
\end{equation*}
$$

\]

Here, for each $\boldsymbol{R}$ the expansion coefficients give ionic functions $\chi_{i}(\boldsymbol{R})$. To find them, we substitute (6.14) in (6.11), apply (6.13), multiply by $\eta_{j}^{*}$ and finally integrate over $\boldsymbol{r}$, to find out:

$$
\left[T_{\boldsymbol{R}}+\epsilon_{j}(\boldsymbol{R})-E\right] \chi_{j}(\boldsymbol{R})=\sum_{i} \Lambda_{j i} \chi_{i}(\boldsymbol{R})
$$

Matrix $\Lambda$ has the elements

$$
\Lambda_{j i}=\sum_{I} \frac{\hbar}{2 M_{I}} \int \mathrm{~d} \boldsymbol{r} \eta_{j}^{*}(\boldsymbol{R}, \boldsymbol{r}) \frac{\partial^{2} \eta_{i}(\boldsymbol{R}, \boldsymbol{r})}{\partial \boldsymbol{R}_{I}^{2}}+\left(\sum_{I} \frac{\hbar}{M_{I}} \int \mathrm{~d} \boldsymbol{r} \eta_{j}^{*}(\boldsymbol{R}, \boldsymbol{r}) \frac{\partial \eta_{i}(\boldsymbol{R}, \boldsymbol{r})}{\partial \boldsymbol{R}_{I}} \frac{\partial \chi_{i}(\boldsymbol{R})}{\partial \boldsymbol{R}_{I}}\right)
$$

Obviously, for large ionic masses it is small. Adiabatic approximation consists in neglecting $\Lambda$. Then the previous equation becomes ionic Schrödinger equation:

$$
\left[T_{\boldsymbol{R}}+\epsilon_{j}(\boldsymbol{R})\right] \chi_{m j}(\boldsymbol{R})=E_{m j} \chi_{m j}(\boldsymbol{R})
$$

This equation contains no electronic degrees of freedom. On the other hand, the electrons appears here through their eigenenergies, which have the role of the external potential defining the ion dynamics.

To summarize, adiabatic approximation suffices to separate electron and aions. Within this model, the electronic equation is to be solved first, and their energies give the potential in which ions move. To study the ground state of the total system, one takes the ground state of electrons, with energi $\epsilon_{0}$, and then solves equation of ions in the external potential $\epsilon_{0}$.

Despite the success in reduction the problem to separate equations in electrons and ions, each of these subsystems is usually complex enough, and further approximations are employed. While for the electrons various approximations will be considered later, here we mention that for ions it is usual to adopt Born approximation. It assumes that the ions are in the position $\boldsymbol{R}_{0}$ corresponding to minimum of the potential $\epsilon_{0}$. Then, their dynamics is by harmonic approximation reduced to vibrations around this minimum.

### 6.3 Variational Method

### 6.3.1 General characteristics

## Foundation

The variational method is based on the well known mathematical theorem on the coincidence of the eigenvectors of an observable, with the stationary vectors of the functional defined over the domain of $A$ as the mean $\langle\psi| A|\psi\rangle$ of $A$. We are going to prove it, in the form suitable for further purposes.

Theorem 6.2 Let the Hamiltonian $H$ be with purely discrete spectrum, and let the energy functional be defined as:

$$
\begin{equation*}
E[|\psi\rangle] \stackrel{\text { def }}{=} \frac{\langle\psi| H|\psi\rangle}{\langle\psi \mid \psi\rangle} \tag{6.15}
\end{equation*}
$$

Then the stationary points of $E[|\psi\rangle]$ coincide with the eigenvectors of $H$; the corresponding stationary values are the eigenvalues of $H$.
-Proof: The variation of (6.15) in the form $\langle\psi|(H-E[|\psi\rangle])|\psi\rangle=0$, with the stationary condition $\delta E[|\psi\rangle]=0$ gives equation $\langle\delta \psi|(H-E[|\psi\rangle])|\psi\rangle+\langle\psi|(H-E[|\psi\rangle])|\delta \psi\rangle=0$. Due to the independence of the variation of the vectors, $|\delta \psi\rangle$, and functionals, $\langle\delta \psi|$, both terms in the last equation must vanish, imposing the eigenvalue condition: $(H-E[|\psi\rangle])|\psi\rangle=0$.

In addition, the same proof shows that in any invariant subspace in $\mathcal{S}$, the stationary points of $E$ in any invariant subspace are the eigen vectors of $H$ from that subspace (note that besides the vectors, also the variations $|\delta \psi\rangle$ in the proof are from this subspace). Altogether, the variational problem over the entire state space $\mathcal{S}$ is equivalent to the eigenvalue problem of $H$, and can be used in the context of the dynamics of the system. Still, in this form it actually reduces to eigenproblem, giving no advantage to its solution. Nevertheless, the following trivial consequence of the statement 6.2 is fruitful.
Theorem 6.3 The ground state energy (minimal eigenvalue of $H$ ) is less than or equal to the mean value of the hamiltonian in any state:

$$
\begin{equation*}
E_{0} \leq E[|\psi\rangle]=\frac{\langle\psi| H|\psi\rangle}{\langle\psi \mid \psi\rangle} \tag{6.16}
\end{equation*}
$$

-Proof: With help of probability $v_{n}=v\left(E_{n}, H,|\psi\rangle\right)$ mean is $E[|\psi\rangle]=\sum_{n} E_{n} v_{n} \geq E_{0} \sum_{n} v_{n}=E_{0}$. $\mathbf{I}$
Now one uses variational formulation of the eigenproblem to introduce the following approximation. Instead of the whole state space $\mathcal{S}$ only its part $\mathcal{T}$, called trial set, is considered. Then the variational method gives the stationary points within $\mathcal{T}$. According to the theorem 6.3, it is expected that the ground state $|0\rangle$ is best approximated within $\mathcal{T}$ by the state $\left|\psi_{0}\right\rangle$, giving the minimal $E\left[\left|\psi_{0}\right\rangle\right]$. Of course, the quality of the approximation is determined by the choice of $\mathcal{T}$.

This stress out that the choice of $\mathcal{T}$ is the critical point of the approximation. Therefore, to improve the results all a priori known properties of the ground state should be included. For example, frequently the symmetry or number of zeros of the ground state wave function are a priory clear, and the trial functions with such characteristics are to be considered. This is also important in approximations of the excited states within this method. If the ground state, either exact $|0\rangle$, or approximate $\left|\psi_{0}\right\rangle$, is known, the trial set is to be chosen among the orthogonal vectors.

## Standard realizations

To enable variational calculus, i.e. analysis, the trial set wuite generally must be a manifold. Therefore, the states are taken in the parameterized form $\left|\psi\left(a_{1}, a_{2}, \ldots\right)\right\rangle$. Then the variational derivative is found as the derivative over the parameters, and the stationary state is obtained as a value of the parameters $a_{1}, a_{2}$, etc., satisfying all the equations

$$
\begin{equation*}
\frac{\partial}{\partial a_{i}} \frac{\left\langle\psi\left(a_{1}, a_{2}, \ldots\right)\right| H\left|\psi\left(a_{1}, a_{2}, \ldots\right)\right\rangle}{\left\langle\psi\left(a_{1}, a_{2}, \ldots\right) \mid \psi\left(a_{1}, a_{2}, \ldots\right)\right\rangle}=0 \tag{6.17}
\end{equation*}
$$

Most frequently, this procedure is performed in the coordinate representation.
Particularly, if the trial set is a subspace, then the parameters are naturally chosen as the coefficients $a_{i}$ in the expansion $|\psi\rangle=\sum_{i} a_{i}\left|\phi_{i}\right\rangle$ of the arbitrary vector over some basis $\left|\phi_{i}\right\rangle$ of $\mathcal{T}$. Then, denoting the Hamiltonian matrix elements in this subbasis by $H_{i j}=\left\langle\phi_{i}\right| H\left|\phi_{j}\right\rangle$, the state norm in the equations (6.17) may be restricted to 1 , which introduces the mean energy as the Lagrange multiplier:

$$
\frac{\partial}{\partial a_{i}}\left(\sum_{p q} a_{p}^{*} a_{q} H_{p q}-E \sum_{p q} a_{p}^{*} a_{q}\right), \text { i.e. } \sum_{p} a_{p}^{*}\left(H_{p q}-E\right)=0 \forall q \text {. }
$$

Since $\left|\phi_{i}\right\rangle$ form the basis in the subspace $\mathcal{T}$, in the last equation only the matrix elements of the restricted operator $H_{\mathcal{T}}=P_{\mathcal{T}} H P_{\mathcal{T}}$ are involved (as usual, $P_{\mathcal{T}}$ is the projector onto $\mathcal{T}$ ). Hence, this equation is exactly the eigenvalue problem $\langle\psi|\left(H_{\mathcal{T}}-E\right)=0$, its $q$-th component written in the $\left|\phi_{i}\right\rangle$ representation. Indeed, this is anticipated when immediately after the Theorem 6.2 it is generalized to any invariant subspace of $H$ : the only relevant operator in the variational procedure turns out to be the restriction $\mathcal{H}_{\mathcal{T}}$, for which the subspace $\mathcal{T}$ is invariant. Therefore, the variational problem with linear trial set reduces to the eigenvalue problem of the restriction.

It is also interesting to compare the variational method to the stationary perturbations. To this end let Hamiltonian be $H=H_{0}+H^{\prime}$, while the trial set is chosen as an eigen space of $H_{0}: \mathcal{T}=\mathcal{S}_{n}^{0}$. The stationary states are found as the eigen states of the restricted Hamiltonian $P_{n}\left(H_{0}+H^{\prime}\right) P_{n}=E_{n}^{0}+P_{n} H^{\prime} P_{n}$. Obviously, the states $|n i \lambda ; 0\rangle$ are obtained, with encountered correction of energy of the first order: $E[|n i \lambda ; 0\rangle]=E_{n}^{0}+E_{n i}^{(1)}$.

## Error estimates

There is no precise estimate of the error, but the following results are useful in this context.

Theorem 6.4 Let $\{|0\rangle,|1\rangle, \ldots\}$ and $\left\{E_{0}, E_{1}, \ldots\right\}$ be the exact eigenvectors and the corresponding eigenvalues of the Hamiltonian, in the increasing energy order.
(i) For each physical state $|\psi\rangle$ there is at least one $n$ such that

$$
\mid E_{n}-E[|\psi\rangle] \mid \leq \triangle_{|\psi\rangle} H
$$

(ii) If the difference between the variational and the exact ground states $\left|\psi_{0}\right\rangle$ and $|0\rangle$ is estimated by $\epsilon \stackrel{\text { def }}{=} 1-\left|\left\langle\psi_{0} \mid 0\right\rangle\right|$, then for each $|\psi\rangle$ orthogonal to $\left|\psi_{0}\right\rangle$ :

$$
\epsilon \geq|\langle\psi \mid 0\rangle|^{2}, \quad E[|\psi\rangle] \geq E_{1}-\epsilon\left(E_{1}-E_{0}\right)
$$

-Proof: (i) According to the previous theorems, there surely exists $E_{n}$ closest to the $E[|\psi\rangle]$. Then, using the mean in $|\psi\rangle$ of the spectral form $(H-E[|\psi\rangle])^{2}=\sum_{i}\left(E_{i}-E[|\psi\rangle]\right)^{2} P_{i}$, one gets

$$
\triangle_{|\psi\rangle}^{2} H \geq \sum_{i}\langle\psi|\left(E_{n}-E[|\psi\rangle]\right)^{2} P_{i}|\psi\rangle=\left(E_{n}-E[|\psi\rangle]\right)^{2} .
$$

(ii) For any vector $|\psi\rangle$ orthogonal to $\left|\psi_{0}\right\rangle$ one finds

$$
1=\langle 0 \mid 0\rangle=\left\langle 0 \mid \psi_{0}\right\rangle\left\langle\psi_{0} \mid 0\right\rangle+\langle 0 \mid \psi\rangle\langle\psi \mid 0\rangle+\cdots \leq 1-\epsilon+|\langle 0 \mid \psi\rangle|^{2},
$$

which proves the left inequality. The right one follows from: $E[|\psi\rangle]=\sum_{n n^{\prime}}\langle\psi \mid n\rangle\langle n| H\left|n^{\prime}\right\rangle\left\langle n^{\prime} \mid \psi\right\rangle=\sum_{n} E_{n}|\langle\psi \mid n\rangle|^{2}=$
$E_{0}|\langle\psi \mid 0\rangle|^{2}-E_{1}|\langle\psi \mid 0\rangle|^{2}+E_{1}|\langle\psi \mid 0\rangle|^{2}+\sum_{n(>0)} E_{n}|\langle\psi \mid n\rangle|^{2} \geq$
$\left(E_{0}-E_{1}\right)|\langle\psi \mid 0\rangle|^{2}+E_{1} \sum_{n}|\langle\psi \mid n\rangle|^{2} \geq-\left(E_{0}-E_{1}\right) \epsilon+E_{1}$. I
The first part of the theorem asserts that any state, including any state stationary over $\mathcal{T}$, is close to an eigenstate (in the $\triangle H$ neighborhood), but it is not specified to which eigenstate. Hence, using this criterion it cannot be estimated how far from the ground state the result is.

### 6.3.2 Hartree-Fock method

Correlation of many particle states is the main obstacle to the exact dynamical approach. They cannot be avoided, as the (anti)symmetrization and interaction make the noncorrelated states not physically real. As it has been discussed, the minimal correlation is pure kinematical exchange one, introduced by the indistinguishability of the particles. As for the electrons, this Pauli correlation is included by the Hartree-Fock variational method: the trial set comprises all the Slater determinants. It should be immediately stressed out that the dynamical correlation is completely neglected, which restricts the scope of the applications: the typical error in the obtained ground state energy is 0.1 eV is good enough for prediction of the most of the physical (e.g. electro-optical properties), but too large for reliable predictions of the stable molecular configuration. Therefore, Hartree-Fock method has been during the past decade progressively substituted by the density functional methods, to accommodate requirements of high-technology and materials science.

## Energy functional

The cornerstone of the Hartree-Fock method is the restriction to the trial set of Slater determinants. This results in the extremely simple form (5.22b) of the corresponding single- and two-particle reduced states, which are in turn enough to find convenient general form (for this trial set!) of the energy functional. Indeed, as at most pairwise interactions are assumed, the hamiltonian is $H=H^{[1]}+H^{[2]}$, with the single-particle term $H^{[1]}=$ $\sum_{p=1}^{N} H_{p}$; here $H_{p}=T_{p}+U_{p}$ includes kinetic energy and potential of the possible external field. The pairwise Coulomb interactions $V_{p q}$, determine $H^{[2]}=\frac{1}{2} \sum_{p_{1} \neq p_{2}} V_{p_{1} p_{2}}$. Therefore, for the Slater determinant composed of the states $\left|\phi_{1}\right\rangle, \ldots,\left|\phi_{N}\right\rangle$, making the single particle space projector $P=\sum_{i=1}^{N}\left|\phi_{i}\right\rangle\left\langle\phi_{i}\right|=N \rho_{1}$, the energy functional, i.e. the mean value of the hamiltonian, by (5.28) reads:

$$
\begin{equation*}
E[P]=\operatorname{Tr}_{1}\left(H_{1}+\frac{1}{2} U_{1}^{\mathrm{HF}}\right) P=\sum_{p=1}^{N}\left\langle\phi_{p}\right|\left(H_{1}+\frac{1}{2} U_{1}^{\mathrm{HF}}\right)\left|\phi_{p}\right\rangle, \tag{6.18}
\end{equation*}
$$

with $V=V_{12}$. The crucial point is that the two-particle terms reduce to the formally single particle operator:

$$
\begin{equation*}
U_{1}^{\mathrm{HF}} \stackrel{\text { def }}{=} 2 \operatorname{Tr}_{2} P_{2}^{(-)} V P=2 N \operatorname{Tr}_{2} P_{2}^{(-)} V \rho_{2} \tag{6.19}
\end{equation*}
$$

here, $P$ and $\rho_{1}$ are the operators in the second particle space. In other words, the Slater determinants approximation, change the dynamical model to that of the noninteracting particles in the external field $U+\frac{1}{2} U^{\mathrm{HF}}$.

Explicitly, the terms related to the two-particle interactions in (6.18) are:

$$
\begin{equation*}
\left\langle\phi_{p}\right| \frac{1}{2} U_{1}^{\mathrm{HF}}\left|\phi_{p}\right\rangle=\frac{1}{2} \sum_{q}\left(\left\langle\phi_{p} \phi_{q}\right| V\left|\phi_{p} \phi_{q}\right\rangle-\left\langle\phi_{p} \phi_{q}\right| V\left|\phi_{q} \phi_{p}\right\rangle\right) \tag{6.20a}
\end{equation*}
$$

(note that the term $p=q$ is cancelled). The first term is called direct or Hartree's and the second one exchange or Fock's (the ordinal of the space is given within the vector for clarity):

$$
\begin{align*}
U_{1}^{\mathrm{H}} & =\sum_{p=1}^{\infty} \sum_{q=1}^{N}\left|\phi_{p} 1\right\rangle\left\langle\phi_{p} 1\right|\left\langle\phi_{q} 2\right| V\left|\phi_{q} 2\right\rangle=\sum_{q=1}^{N}\left\langle\phi_{q} 2\right| V\left|\phi_{q} 2\right\rangle  \tag{6.20b}\\
U_{1}^{\mathrm{F}} & =\sum_{p=1}^{\infty} \sum_{q=1}^{N}\left|\phi_{p} 1\right\rangle\left\langle\phi_{q} 1\right|\left\langle\phi_{p} 2\right| V\left|\phi_{q} 2\right\rangle . \tag{6.20c}
\end{align*}
$$

Their coordinate representation (assuming position dependent interaction potential) is

$$
\begin{align*}
U_{1}^{\mathrm{H}}(\boldsymbol{r}) & =\int \mathrm{d} \boldsymbol{r}_{2} \sum_{q} \phi_{q}^{*}\left(\boldsymbol{r}_{2}\right) \phi_{q}\left(\boldsymbol{r}_{2}\right) V\left(\boldsymbol{r}, \boldsymbol{r}_{2}\right),  \tag{6.20d}\\
U_{1}^{\mathrm{F}}\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right) & =\sum_{q} \phi_{q}^{*}\left(\boldsymbol{r}^{\prime}\right) \phi_{q}(\boldsymbol{r}) V\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right) \tag{6.20e}
\end{align*}
$$

The Hartree contribution is diagonal in the coordinates, i.e. it is usual multiplicative single-particle potential. In fact, this is simply averaged interaction with other electrons, and due to the exchange it must be corrected by the Fock's term, which is not of this type. The method in which exchange term is neglected (and possibly treated perturbatively later on) is therefore called Hartree's, while the Hartree-Fock method proceeds with both terms.

## Equation of extrema

Having thus the energy functional expressed as the single particle operator mean value, we can perform variational calculus to find the extremes. At first, we emphasize again that the Slater determinants are considered only, which results in the approximation to the ground state with the minimal, Pauli, correlation. On the other hand, due to the biunique relation of the N -particle Slater determinants and N -dimensional projectors in the single particle state space, the trial set is formed by the $N$-dimensional projectors (or subspaces) in $\mathcal{S}$. Such a set is manifold (thus the variational approach is eligible), but the realization of the variational problem is more transparent within the constrained (conditional) variation scheme: we perform the variation over the linear space of the hermitian operators (apparently enlarging the trial set by the physically senseless operators, not corresponding to the states), but compensate this by the constraint $P^{2}-P=0$. This introduces the Lagrange multiplier $\Lambda$, and the variation of the energy functional becomes:

$$
\delta E[P]=\delta \operatorname{Tr}_{1}\left[\left(H_{1}+\operatorname{Tr}_{2}\left(P_{2}^{(-)} V P\right)\right) P+\Lambda\left(P^{2}-P\right)\right]=0
$$

Here, $\Lambda$ is matrix of the dimension of the single particle space, in which the $N$-dimensional projectors $P$ are trial set. The dimension of the projector is automatically provided in the course of the realization of the variational method. Indeed, the initial condition should be appropriately chosen, i.e. such $N$-dimensional projector is started with, and in the view of the continuity of the procedure, only the manifold of the same dimensional projectors is taken into account. Since $\Delta(\tau) P_{2}^{(-)} V(\delta P \otimes P) \Delta(\tau)=P_{2}^{(-)} V(P \otimes \delta P)$, and $\operatorname{Tr}_{1} \Lambda\left(P^{2}-P\right)=\operatorname{Tr}_{1}(P \Lambda-\Lambda) P$, the equation for the variational extreme is

$$
\operatorname{Tr}[(h+\Lambda P+P \Lambda-\Lambda) \delta P]=0
$$

with the single particle effective hamiltonian $h=H_{1}+U_{1}^{\mathrm{HF}}$. The trace is over the single particle space, which is the total space of the equation. The variational equation has the form of the operator scalar product; thus, since the variation is arbitrary hermitian operator, one gets the equivalent algebraic equation for the extremes:

$$
\begin{equation*}
h+\Lambda P+P \Lambda-\Lambda=0 . \tag{6.21}
\end{equation*}
$$

Multiplying it by $P$, firstly from left and then from right, and subtracting the obtained equations, we get:

$$
\begin{equation*}
[h, P]=0 . \tag{6.22}
\end{equation*}
$$

However, we easily verify that whenever (6.22) is satisfied by some $h$ and $P$, then also the extremal equation (6.21) is solved by $\Lambda=h-h P-P h$. Therefore, (6.22) is the solution which is looked for, and it is called the Hartree-Fock equation. If in $U_{1}^{\mathrm{HF}}$ the exchange potential is omitted, one gets the Hartree equation of the same form. Note again that the considerably simple algebraic form of the Hartree-Fock equation (6.22) originates in the effectively single particle procedure, enabled by the neglecting of the dynamical correlations.

## Self consistent solution

It is important that $U_{1}^{\mathrm{HF}}$, and this way $h$, is itself determined by the state $P$ (precisely, by the Slater determinant corresponding to $P$ ), which is looked for. This makes the equation (6.22) nonlinear in $P$, and the procedure for solving it is necessarily iterative. To understand this approach, consider the final solution: we obtained the Hartree-Fock hamiltonian $h$ commuting with the $N$-dimensional projector $P$. Consequently, within the range of $P$ there are $N$ states (single particle) which are eigenstates of $h$ :

$$
\begin{equation*}
h\left|i_{p}\right\rangle=\epsilon_{p}\left|i_{p}\right\rangle, \quad p=1, \ldots, N . \tag{6.23}
\end{equation*}
$$

As the dynamical model is reduced to the independent (noninteracting) electrons, we may assume that the electrons, due to the Pauli principle, occupy these $N$ states. In this sense, the operator $h$ takes the role of the hamiltonian; nevertheless, the total energy is not the sum of the above eigenvalues, as $U^{\mathrm{HF}}$ is multiplied by $\frac{1}{2}$ in (6.18). Further, if there was not mutual interaction of electrons, i.e. if $U=0$, these states would be real single particle states (and $h$ would be real single particle hamiltonian).

Thus one starts with a Slater determinant constructed with some states $\left|i_{1}^{0}\right\rangle, \ldots,\left|i_{N}^{0}\right\rangle$, giving the projector $P^{0}$. Usually, they correspond to the noninteracting case of the considered problem; as Hartree-Fock method is usually used to find the atomic or molecular orbitals, one can use the first $N$ hydrogen like orbitals. Then, for this state $U^{\mathrm{HF}}$ and $h$ are to be calculated: $U^{\mathrm{HF}}\left[P^{0}\right]=2 \operatorname{Tr}_{2} P_{2}^{(-)} V^{(12)} P^{0}, h^{0}=T+U+U^{\mathrm{HF}}\left[P^{0}\right]$. Next, the eigen problem of $h^{0}$ is solved, to get the eigenvectors $\left|i_{p}^{1}\right\rangle$, and the eigenergies $\epsilon_{p}^{1}(p=1, \ldots,|\mathcal{S}|)$. The lowest $N$ energies select the eigenstates $\left|i_{1}^{1}\right\rangle, \ldots,\left|i_{N}^{1}\right\rangle$ forming new Slater determinant and projector $P^{1}$. The procedure is then repeated: in the $s$-th step we calculate single particle hamiltonian $h^{s}$, with eigenvectors determining the next step:

$$
h^{s}\left|i_{p}^{s+1}\right\rangle=\epsilon_{p}^{s+1}\left|i_{p}^{s+1}\right\rangle, \quad p=1, \ldots, N .
$$

The solution is the fixed point of the procedure: after some iterations, new state is same as the previous. If the initial state is considerably well estimated, the procedure converges, providing self consistency of the solution. The obtained solution surely is the extreme of energy (within the trial set of the Slater determinants), but it is not obvious if it is minimum. However, the choice of the lowest energy eigenvectors usually leads to the minimum.

As mentioned above, the total energy of the resulting state is not the sum of the obtained eigenenergies; indeed by (6.18) one finds:

$$
E[P]=\sum_{i} \epsilon_{i}-\frac{1}{2}\left\langle U^{\mathrm{HF}}\right\rangle
$$

This is because the sum includes each pairwise interaction twice. Therefore, the eigenvalues $\epsilon_{i}$ are not to be interpreted as the energies of the electrons in the corresponding states $|i\rangle$, but the ionization energies necessary to separate particle in that state (Koopmans' theorem).

### 6.4 Density Functional Theory

The Hartree-Fock method essentially reduces the interacting many-body problem to the noninteracting one, with suitably chosen single particle external potential. In fact, this includes a part of the inter-particle interaction as the mean field in the assumed state. The cost for the reduction to the single particle problem is neglecting of correlations (except the Pauli exchange), due to the trial function of the Slater determinant type.

The same concept inspired much more precise method called density functional theory. It is based on the following simple:

Theorem 6.5 [Hohenberg-Kohn, 1964] Reduced density matrix $\rho_{1}$ of the ground state uniquely determines the external potential $U$. Particularly, for the position dependent potential, the density function $\rho_{1}(\boldsymbol{r})=\langle\boldsymbol{r}| \rho_{1}|\boldsymbol{r}\rangle$ determines $U$.
-Proof: We assume that in the external potential $U$ the ground state of the system is $\rho$, with the reduced density matrix $\rho_{1}$. Let there is another external potential $U^{\prime}$ giving the ground state $\rho^{\prime}$ with the same
reduced statistical operator $\rho_{1}$. With $H_{0}=T+V$, the minimal energy (??) of the ground states in the potentials $U$ and $U^{\prime}$ are related as:

$$
E_{U}[\rho]=\operatorname{Tr} H_{0} \rho+N \operatorname{Tr}_{1} U \rho_{1}<E_{U}\left[\rho^{\prime}\right]=\operatorname{Tr} H_{0} \rho^{\prime}+N \operatorname{Tr}_{1} U \rho_{1}=E_{U^{\prime}}\left[\rho^{\prime}\right]+N \operatorname{Tr}_{1}\left(U-U^{\prime}\right) \rho_{1}
$$

Analogously, $E_{U^{\prime}}\left[\rho^{\prime}\right]<E_{U}[\rho]+N \operatorname{Tr}_{1}\left(U^{\prime}-U\right) \rho_{1}$. Then, adding these two equations, one gets contradiction $E_{U^{\prime}}\left[\rho^{\prime}\right]+E_{U}[\rho]<E_{U^{\prime}}\left[\rho^{\prime}\right]+E_{U}[\rho]$, proving that if in the potential $U$ the reduced ground state density is $\rho_{1}$, there is no other potential producing the same ground state density. This conclusion is consequence of the same $\rho_{1}$ in the last term; When the external potential is position dependent only, this term reads $\operatorname{Tr}_{1}\left(U-U^{\prime}\right) \rho_{1}=\int \mathrm{d} \boldsymbol{r} \rho_{1}(\boldsymbol{r})\left(U(\boldsymbol{r})-U^{\prime}(\boldsymbol{r})\right)$, and the equality of density functions is sufficient. I

Although in the above proof the non-degenerate ground state is considered, the generalization to degenerate ground state exists in literature. Further, there are densities $\rho_{1}(\boldsymbol{r})$ being not reduced states of the ground state for any potential. Thus, the theorem means that there is biunique correspondence between potentials $U$ and the densities of the ground states of a many-particle system in some external potential; such densities are called $U$-representable.

In fact, a system of many particles of some sort defines itself an inter-particle interaction $V$, being universal in the sense that it has the same form for these particles independently on the specific circumstances. Analogously, the kinetic energy is universal, and only the external potential defines different physical situations. Now, the HohenbergKohn theorem emphasizes the importance of the densities: they biuniquely correspond to different physical situations, i.e. to the associated ground states. When this is understood, the theorem implies that the ground state is functional of $\rho_{1}(\boldsymbol{r})$, as well as the ground state mean value of any many body observable $A$ :

$$
\begin{equation*}
\langle A\rangle \stackrel{\text { def }}{=} \operatorname{Tr} A \rho=A\left[\rho_{1}(\boldsymbol{r})\right] . \tag{6.24}
\end{equation*}
$$

The last conclusion applied to the energy functional immediately gives:
Theorem 6.6 [Hohenberg-Kohn, 1964] The energy functional

$$
\begin{equation*}
E_{U}\left[\rho_{1}\right]=F\left[\rho_{1}\right]+\operatorname{Tr}_{1} U \rho_{1} \tag{6.25}
\end{equation*}
$$

reaches its minimum for the ground state reduced density $\rho_{1}$ corresponding to $U$. Here, $F$ is functional universal for the given type of particles.

Instead of $\rho_{1}$, to fix the number of particles $N$, we use $\varrho=N \rho_{1}$. The function $\varrho(\boldsymbol{r})=\langle\boldsymbol{r}| \varrho|\boldsymbol{r}\rangle$ is real $N$-particle density, due to the trace relation

$$
\begin{equation*}
\operatorname{Tr}_{1} \varrho=\int \mathrm{d} \boldsymbol{r} \varrho(\boldsymbol{r})=N \tag{6.26}
\end{equation*}
$$

## Density variational principle

It is well known that the $N$-particle ground state $|\Psi\rangle$ is the minimal state of the energy mean. The variational procedure may be performed in two steps. Firstly, for each fixed $\varrho(\boldsymbol{r})$, one may find $\left|\Psi_{\varrho}\right\rangle$ as the minimum of the energy mean over the states giving same $\varrho(\boldsymbol{r})$. Thus this mean is the functional of $\varrho(\boldsymbol{r})$ :

$$
\begin{equation*}
E\left[\left|\Psi_{\varrho}\right\rangle\right]=\min _{\Phi \rightarrow \varrho(\boldsymbol{r})}\langle\Phi| H|\Phi\rangle=F[\varrho(\boldsymbol{r})]+\int \mathrm{d} \boldsymbol{r} U(\boldsymbol{r}) \varrho(\boldsymbol{r}) \tag{6.27}
\end{equation*}
$$

This defines the universal (exterior field independent) functional

$$
\begin{equation*}
F[\varrho(\boldsymbol{r})] \stackrel{\text { def }}{=} \min _{\Phi \rightarrow \varrho(\boldsymbol{r})}\langle\Phi| T+V|\Phi\rangle=\left\langle\Psi_{\varrho}\right| T+V\left|\Psi_{\varrho}\right\rangle . \tag{6.28}
\end{equation*}
$$

In the next step, the ground state is found as the minimum of $E\left[\left|\Psi_{\varrho}\right\rangle\right]$ over densities, i.e. the states $\left|\Psi_{\varrho}\right\rangle$. The Euler-Lagrange equation for the constrained variational extremes ${ }^{3}$ is

$$
\begin{equation*}
\frac{\delta F[\varrho(\boldsymbol{r})]}{\delta \varrho(\boldsymbol{r})}+U(\boldsymbol{r})=\mu, \tag{6.29}
\end{equation*}
$$

where the chemical potential $\mu$ is the Lagrange coefficient for the density constraint (6.26).
Determination of the functional $F[\varrho(\boldsymbol{r})]$ would transform (6.29) into equation in $\varrho$, making $N$-particle problem exactly solvable. However, the functional is not known. In principle, it is a sum of the kinetic energy $(T[\varrho(\boldsymbol{r})])$ and potential containing the direct (Hartree) term $(D[\varrho(\boldsymbol{r})])$, Fock exchange $(X[\varrho(\boldsymbol{r})])$ and correlation $(C[\varrho])$ functionals ${ }^{4}$. The last two terms are usually joined into the exchange-correlation functional $E_{\mathrm{xc}}[\varrho(\boldsymbol{r})]$. These potentials will be precisely determined later. The kinetic energy and correlation functional are not known exactly.

To handle the kinetic energy, we consider a system of noninteracting particles, i.e. $V=0$, when essentially the single particle approximation is fulfilled, and instead of $F$ the kinetic energy functional $T_{0}$ is obtained. Thus, the hamiltonian is $H_{0}=T+U$, and for fixed $\varrho(\boldsymbol{r})$ one can define $\left|\Psi_{\varrho 0}\right\rangle$ and the ground state $\left|\Psi_{0}\right\rangle$ :

$$
\begin{equation*}
T_{0}[\varrho(\boldsymbol{r})] \stackrel{\text { def }}{=} \min _{\Phi \rightarrow \varrho(\boldsymbol{r})}\langle\Phi| T|\Phi\rangle=\left\langle\Psi_{\varrho 0}\right| T\left|\Psi_{\varrho 0}\right\rangle \tag{6.30}
\end{equation*}
$$

giving the variational minimum by the equation:

$$
\begin{equation*}
\frac{\delta T_{0}[\varrho(\boldsymbol{r})]}{\delta \varrho(\boldsymbol{r})}+U_{0}(\boldsymbol{r})=\mu \tag{6.31}
\end{equation*}
$$

Now, assuming that $\varrho(\boldsymbol{r})$ is both interacting and noninteracting system $U$-representable, one can think of (6.29) and (6.31) as describing equivalent (via $\varrho(\boldsymbol{r})$ ) interacting and noninteracting systems. Their equivalence is achieved by the redefined (with respect to $U(\boldsymbol{r})$ ) external field $U_{0}(\boldsymbol{r})$ (which is also suited to get the same Lagrange multiplier $\mu$ ). Thus, the Euler-Lagrange equation for the noninteracting system give Kohn-Sham equations:

Theorem 6.7 [Kohn-Sham, 1965] The exact $N$-particle ground state density $\varrho(\boldsymbol{r})$ is $\varrho=\sum_{p=1}^{N}|p\rangle\langle p|$, where $|p\rangle$ are the $N$ lowest eigenvectors of the Kohn-Sham single-particle hamiltonian $H_{0}=T+U_{0}$ :

$$
\begin{equation*}
H_{0}|p\rangle=\epsilon_{p}|p\rangle, \quad U_{0}(\boldsymbol{r})=U(\boldsymbol{r})+\frac{\delta D[\varrho(\boldsymbol{r})]}{\delta \varrho(\boldsymbol{r})}+\frac{\delta E_{\mathrm{xc}}[\varrho(\boldsymbol{r})]}{\delta \varrho(\boldsymbol{r})} . \tag{6.32}
\end{equation*}
$$

[^7]This equation can be solved self consistently, in analogy with Hartree-Fock method if all the potentials are known. For the noninteracting Kohn-Sham orbitals (giving $\left|\Psi_{0}\right\rangle$ as their Slater determinant) $T_{0}[\varrho(\boldsymbol{r})]$ is found exactly, as well as the Hartree direct term, $D[\varrho(\boldsymbol{r})]=\frac{1}{2} \int \mathrm{~d} \boldsymbol{r} \mathrm{~d} \boldsymbol{r}^{\prime} \varrho(\boldsymbol{r}) \varrho\left(\boldsymbol{r}^{\prime}\right) v\left(\boldsymbol{r}-\boldsymbol{r}^{\prime}\right)$, introducing the potential (6.20d). Since Slater determinant $\left|\Psi_{0}\right\rangle$ involves no correlation, $X[\varrho(\boldsymbol{r})] \stackrel{\text { def }}{=}\left\langle\Psi_{0}\right| V\left|\Psi_{0}\right\rangle-D[\varrho(\boldsymbol{r})]$, giving $\left\langle\Psi_{0}\right| T+$ $V\left|\Psi_{0}\right\rangle=T_{0}[\varrho(\boldsymbol{r})]+D[\varrho(\boldsymbol{r})]+X[\varrho(\boldsymbol{r})]$. As $F[\varrho(\boldsymbol{r})]=\langle\Psi| T+V|\Psi\rangle$, and $E_{x c}[\varrho(\boldsymbol{r})]=$ $X[\varrho(\boldsymbol{r})]+C[\varrho(\boldsymbol{r})]$ it remains:

$$
\begin{equation*}
C[\varrho]=F[\varrho]-\left(T_{0}[\varrho]+D[\varrho]+X[\varrho]\right)=\langle\Psi|(T+V)|\Psi\rangle-\left\langle\Psi_{0}\right|(T+V)\left|\Psi_{0}\right\rangle . \tag{6.33}
\end{equation*}
$$

Note that when $V=0$, single particle case, $X[\varrho(\boldsymbol{r})]=-D[\varrho(\boldsymbol{r})]$, making $C[\varrho(\boldsymbol{r})]=0$ (since a priori $F[\varrho(\boldsymbol{r})]=T_{0}[\varrho(\boldsymbol{r})]$ ). For $V \neq 0$, since $|\Psi\rangle$ is the minimal point of $\langle T+V\rangle$, $C[\varrho(\boldsymbol{r})] \leq 0$. On the contrary, $\left|\Psi_{0}\right\rangle$ minimalize $\langle T\rangle$, making $\langle T\rangle_{\Psi}-\langle T\rangle_{\Psi_{0}}$ positive. Consequently, the rest $\langle V\rangle_{\Psi}-\langle V\rangle_{\Psi_{0}}$ of $C[\varrho(\boldsymbol{r})]$ is negative.

Thus the main problem is determination of correlation functional. This is performed within some approximations. It appears that it is easier to approximate together exchange and correlation. To understand this energy, we consider family of the hamiltonians, $H_{\lambda}=T+\lambda V+U$, yielding $\left|\Psi_{\lambda}\right\rangle$ with fixed $\varrho(\boldsymbol{r})$ (adiabatic connection). For $\lambda=0$ the Kohn-Sham system is obtained, and for $\lambda=1$ the actual interacting system. Further, we introduce two-particle density operator $\varrho_{2}^{\lambda}=N(N-1) \operatorname{Tr}_{3, \ldots, N}\left|\Psi_{\lambda}\right\rangle\left\langle\Psi_{\lambda}\right|$, and two particle density $\varrho_{2}^{\lambda}\left(\boldsymbol{r}^{\prime}, \boldsymbol{r}\right)=\left\langle\boldsymbol{r}, \boldsymbol{r}^{\prime}\right| \varrho_{2}^{\lambda}\left|\boldsymbol{r}, \boldsymbol{r}^{\prime}\right\rangle$, giving the joint probability to find particle around $\boldsymbol{r}^{\prime}$ and another one around $\boldsymbol{r}$. Then

$$
\begin{equation*}
\langle V\rangle_{\Psi_{\lambda}}=\frac{1}{2} \operatorname{Tr} V \varrho_{2}=\frac{1}{2} \int \mathrm{~d} \boldsymbol{r} \mathrm{~d} \boldsymbol{r}^{\prime} \varrho_{2}^{\lambda}\left(\boldsymbol{r}^{\prime}, \boldsymbol{r}\right) v\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right) \tag{6.34}
\end{equation*}
$$

If the conditional probability to find particle near $\boldsymbol{r}^{\prime}$ when one is fixed at $\boldsymbol{r}$ is $n_{2}^{\lambda}\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right)$, then $\varrho_{2}^{\lambda}\left(\boldsymbol{r}^{\prime}, \boldsymbol{r}\right)=\varrho(\boldsymbol{r}) n_{2}^{\lambda}\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right)$. Obviously, $n_{2}^{\lambda}\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right)$ is the density of the remaining particles, when one is at $\boldsymbol{r}$, giving $\int \mathrm{d} \boldsymbol{r}^{\prime} n_{2}^{\lambda}\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right)=N-1$. Finally, to extract the exchangecorrelation part, one writes $n_{2}^{\lambda}\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right)=\varrho\left(\boldsymbol{r}^{\prime}\right)+n_{\text {xc }}^{\lambda}\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right)$. The first term is the direct one, for the noninteracting particles. This can be used in (6.36), written in the form

$$
\begin{equation*}
E_{\mathrm{xc}}[\varrho(\boldsymbol{r})]=\int_{0}^{1} \mathrm{~d} \lambda \frac{\mathrm{~d}}{\mathrm{~d} \lambda}\left\langle\Psi_{\lambda}\right|(T+\lambda V)\left|\Psi_{\lambda}\right\rangle-D[\varrho]=\int_{0}^{1} \mathrm{~d} \lambda\left\langle\Psi_{\lambda}\right|(T+V)\left|\Psi_{\lambda}\right\rangle-D[\varrho] \tag{6.35}
\end{equation*}
$$

(Hellman-Faynman theorem is used on the right). Altogether, one gets:

$$
\begin{equation*}
E_{\mathrm{xc}}[\varrho(\boldsymbol{r})]=\frac{1}{2} \int \mathrm{~d} \boldsymbol{r} \mathrm{~d} \boldsymbol{r}^{\prime} \varrho(\boldsymbol{r}) \bar{n}_{\mathrm{xc}}^{\lambda}\left(\boldsymbol{r}^{\prime}, \boldsymbol{r}\right) V\left(\boldsymbol{r}, \boldsymbol{r}^{\prime}\right), \quad \bar{n}_{\mathrm{xc}}\left(\boldsymbol{r}^{\prime}, \boldsymbol{r}\right)=\int_{0}^{1} \mathrm{~d} \lambda n_{\mathrm{xc}}^{\lambda}\left(\boldsymbol{r}^{\prime}, \boldsymbol{r}\right) \tag{6.36}
\end{equation*}
$$

This shows that the exchange-correlation energy is the interaction between a particle and averaged (with respect to the coupling constant $\lambda$ ) exchange-correlation hole around it. This hole appears due to three factors: (i) self-interaction correction, (ii) Pauli principle, (3) potential $V$. First two of them contribute to the exchange, and the last to correlation.

### 6.4.1 Uniform electron gas

We consider $d$-dimensional noninteracting electron gas of the uniform density $\varrho$. When only translational subgroup of the lattice with all periods equal to $L$ is used, the corresponding eigenorbitals are plane waves $\mathrm{e}^{\mathrm{i} \boldsymbol{k} \cdot \boldsymbol{r}} / \sqrt{L^{d}}$, with eigenenergies $\hbar k^{2} / 2 m$. Assuming $g$
dimensional interior space ( $g=2$ for spin) the density of orbitals in $\boldsymbol{k}$-space is $g L^{d} /(2 \pi)^{d}$. On the other hand, $N=\varrho L^{d}$ electrons from the elementary cell occupy all the states with the momentum $k \leq k_{\mathrm{F}}$ (each with the degeneracy $g$ ). This defines the Fermi sphere with radius $k_{\mathrm{F}}$, such that its volume $S_{d}=\pi^{d / 2} k_{\mathrm{F}}^{d} / \Gamma(d / 2+1)$ contains exactly $N$ states. Therefore the density is:

$$
\begin{equation*}
\varrho=k_{\mathrm{F}}^{d} \frac{g}{2^{d}} \frac{1}{\pi^{d / 2} \Gamma(d / 2+1)}=\frac{\Gamma(d / 2+1)}{\pi^{d / 2} r_{\mathrm{S}}^{d}}, \tag{6.37}
\end{equation*}
$$

where the Seitz radius $r_{\mathrm{S}}$ defines the sphere containing single electron.
The average kinetic energy per electron of these orbitals is $t_{0}=\frac{g}{N} \frac{\hbar^{2}}{2 m} \sum_{k<k_{\mathrm{F}}} k^{2}$; using $\boldsymbol{k}$-space orbital density $g L^{d} /(2 \pi)^{d}$, and $\boldsymbol{k}$-space volume element $\mathrm{d} \boldsymbol{k}^{d}=d \pi^{d / 2} k^{d-1} / \Gamma(d / 2+$ 1) $\mathrm{d} k$, one finds:

$$
\begin{equation*}
t_{0}[\varrho]=\frac{\hbar^{2}}{2 m} \frac{d}{d+2} k_{\mathrm{F}}^{2}=\frac{\hbar^{2}}{2 m} \frac{d}{d+2} \frac{4 \pi}{g} \Gamma^{2 / d}(d / 2+1) \varrho^{2 / d}=\frac{\hbar^{2}}{2 m} \frac{d}{d+2} \frac{4 \Gamma^{2 / d}(d / 2+1)}{g^{2 / d} r_{\mathrm{S}}^{2}} \tag{6.38}
\end{equation*}
$$

Particularly, for $g=2$ and $d=3$ one gets:

$$
\begin{equation*}
T_{0}[\varrho]=\int \mathrm{d} \boldsymbol{r} \varrho(\boldsymbol{r}) t_{0}[\varrho]=\frac{\hbar^{2}}{m} \frac{3^{2 / 3} 3 \pi^{4 / 3}}{10} \int \mathrm{~d} \boldsymbol{r} \varrho^{5 / 3}(\boldsymbol{r}) \tag{6.39}
\end{equation*}
$$

### 6.4.2 Exchange energy

The Fock exchange energy for the Slater determinant $\Psi$ of doubly occupied orbitals $\psi_{i}$ is functional:

$$
\begin{equation*}
E_{\mathrm{x}}[\Psi]=-\frac{1}{2} \sum_{\sigma} \sum_{i, j} \iint \mathrm{~d} \boldsymbol{r} \mathrm{~d} \boldsymbol{r}^{\prime} \frac{\phi_{i \sigma}^{*}(\boldsymbol{r}) \phi_{j \sigma}^{*}\left(\boldsymbol{r}^{\prime}\right) \phi_{i \sigma}\left(\boldsymbol{r}^{\prime}\right) \phi_{j \sigma}(\boldsymbol{r})}{\left|\boldsymbol{r}-\boldsymbol{r}^{\prime}\right|} \tag{6.40}
\end{equation*}
$$

To this end, all the functionals should be known, which is not the case. However, various approximations are applied. These are based on the assumptions on thee locality of thee used functionals. In fact, the functionals having Rietzs-Frechet representations, $F[\rho] \stackrel{\text { def }}{=} \int \mathrm{d} \boldsymbol{r} \phi(\boldsymbol{r}) \rho(\boldsymbol{r})$, are called local. However, there is no direct argument resulting in the locality, and therefore the expansion over derivatives of $\rho_{1}$ must be involved to mimic nonlocality. The zero-order term, without derivatives, thus of the local form, in this expansion gives the local density approximation (LDA). Neglecting of gradient terms corresponds to the uniform gas of particles. Therefore, to get this approximation (giving the leading terms) one starts looks for the relevant energy terms in the uniform gas, where all of them are function of density (being the only parameter of the system). Then, one comes back to the slowly varying density (becoming function of $\boldsymbol{r}$ ), and construct the average of the energy terms with such density; this average is itself a functional of $\rho_{1}$.

### 6.5 Time dependent perturbations

The goal of the time dependent perturbation theory is to describe the response of a quantum system to the external perturbation. We assume that, when isolated, the considered
system is described by the hamiltonian $H_{0}$, with eigenproblem

$$
\begin{equation*}
H_{0}|n\rangle=E_{n}|n\rangle . \tag{6.41}
\end{equation*}
$$

Then, during time interval $[0, t]$, it is exposed to the external field (in a very broad sense, e.g. this includes interaction with some other system). This is described as an additional interaction term in the hamiltonian, the perturbation $H^{\prime}(t)$. After $t$, the field vanishes, and the system is isolated again. Therefore, $H^{\prime}(t)$ is a function of time, nonvanishing during the interval $[0, t]$ only. Very frequent example is periodic perturbation, describing the influence of the electromagnetic radiation.

The response of the system is manifested in excitations: $H^{\prime}(t)$ induces transitions between the various eigenstates of the unperturbed hamiltonian. Thus, the main task is to find these transition probabilities. Precisely, for the system prepared at $t=0$ in the eigenstate $|n\rangle$ (of the isolated hamiltonian $H_{0}$ ), we look for the probability

$$
\begin{equation*}
\left.v\left(1,\left|n^{\prime}\right\rangle\left\langle n^{\prime}\right|,|n, t\rangle\right)=v_{n \rightarrow n^{\prime}}(t)=\left|\left\langle n^{\prime}\right| U(t)\right| n\right\rangle\left.\right|^{2} \tag{6.42}
\end{equation*}
$$

to find the system at $t$ in some other eigenstate $\left|n^{\prime}\right\rangle$.

### 6.5.1 Expansion of the evolution

The task is most efficiently solved within Dirac's picture (Sec. 2.5.2). The total evolution governed by

$$
\begin{equation*}
H(t)=H_{0}+H^{\prime}(t) \tag{6.43}
\end{equation*}
$$

is factorized in the form

$$
\begin{equation*}
U(t)=U^{(0)}(t) U^{\prime}(t) \tag{6.44}
\end{equation*}
$$

While the total evolution $U(t)$ satisfies the differential equation (2.2), the unperturbed evolution $U^{(0)}$ and the remained factor $U^{\prime}(t)$ obey (2.33) and (2.34) (we shorten notation: $\left.U\left(t-t_{0}, t_{0}\right)=U(t-0,0)=U(t)\right):$

$$
\begin{aligned}
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} U^{(0)}(t) & =H_{0} U^{(0)}(t), \quad U^{(0)}(0) \\
\mathrm{i} \hbar \frac{\mathrm{~d}}{\mathrm{~d} t} U^{\prime}(t) & =H_{D}^{\prime}(t) U^{\prime}(t), \quad U^{\prime}(0)
\end{aligned}=I .
$$

The formal integral solution of the last equation

$$
\begin{equation*}
U^{\prime}(t)=I+\frac{1}{\mathrm{i} \hbar} \int_{0}^{t} \mathrm{~d} t_{1} H_{D}^{\prime}\left(t_{1}\right) U^{\prime}\left(t_{1}\right) \tag{6.45}
\end{equation*}
$$

is iterated to get the series

$$
U^{\prime}(t)=\sum_{p=0}^{\infty} U^{\prime(p)}(t), \quad U^{\prime(p)}(t)=\frac{1}{\mathrm{i}^{p} \hbar^{p}} \int_{0}^{t} \cdots \int_{0}^{t_{p-1}} \mathrm{~d} t_{1} \cdots \mathrm{~d} t_{p} H_{D}^{\prime}\left(t_{1}\right) \cdots H_{D}^{\prime}\left(t_{p}\right)
$$

Note the time ordering $t \geq t_{1} \geq t_{2} \geq \cdots \geq 0$. We assume that the series is convergent; as we will use only the first few terms of the series, neglecting the reminder, this convergence
is particularly important, and restrict the considerations to sufficiently small $t$. When $U^{\prime}(t)$ is substituted (6.44) becomes:

$$
\begin{equation*}
U(t)=\sum_{p=0}^{\infty} U^{(p)}(t), \quad U^{(p)}(t)=U^{(0)}(t) U^{\prime(p)}(t) \tag{6.46}
\end{equation*}
$$

Taking into account $H_{D}^{\prime}(t)=U^{(0)^{\dagger}}(t) H^{\prime}(t) U^{(0)}(t)$ and $U^{(0)}\left(t_{i}\right) U^{(0)^{\dagger}}\left(t_{j}\right)=U^{(0)}\left(t_{i}-t_{j}\right)$, the convenient form of the terms is obtained:

$$
\begin{equation*}
U^{(p)}(t)=\frac{1}{\mathrm{i} p \hbar^{p}} \int_{0}^{t} \cdots \int_{0}^{t_{p-1}} \mathrm{~d} t_{1} \cdots \mathrm{~d} t_{p} U^{(0)}\left(t-t_{1}\right) H^{\prime}\left(t_{1}\right) U^{(0)}\left(t_{1}-t_{2}\right) \cdots H^{\prime}\left(t_{p}\right) U^{(0)}\left(t_{p}\right) \tag{6.47}
\end{equation*}
$$

### 6.5.2 Transition amplitudes

As the transition probability (6.42) involves the probability amplitudes $\left\langle n^{\prime}\right| U(t)|n\rangle$, we have to find the matrix elements $\left\langle n^{\prime}\right| U^{(p)}(t)|n\rangle$. To simplify this, we insert the identical operators $I=\sum\left|n_{i}\right\rangle\left\langle n_{i}\right|$ after each $H^{\prime}\left(t_{i}\right)$ in (6.47), which reduces the unperturbed evolution to the phases:

$$
\begin{align*}
\left\langle n^{\prime}\right| U^{(p)}(t)|n\rangle & =\sum_{n_{1}, \ldots, n_{p-1}} \frac{1}{\mathrm{i}^{p} \hbar^{p}} \int_{0}^{t} \cdots \int_{0}^{t_{p-1}} \mathrm{~d} t_{1} \cdots \mathrm{~d} t_{p} \times \\
& \times \mathrm{e}^{-\frac{\mathrm{i}}{\hbar}\left[E_{n^{\prime}}\left(t-t_{1}\right)+E_{n_{1}}\left(t_{1}-t_{2}\right)+\cdots+E_{n_{p-1}}\left(t_{p-1}-t_{p}\right)+E_{n} t_{p}\right]} \times  \tag{6.48}\\
& \times\left\langle n^{\prime}\right| H^{\prime}\left(t_{1}\right)\left|n_{1}\right\rangle\left\langle n_{1}\right| H^{\prime}\left(t_{2}\right)\left|n_{2}\right\rangle \cdots\left\langle n_{p-1}\right| H^{\prime}\left(t_{p}\right)|n\rangle .
\end{align*}
$$

In practice, the series (6.42) is truncated to a few first terms. We explicitly find them for $p=0,1,2$ :

$$
\begin{align*}
\left\langle n^{\prime}\right| U^{(0)}(t)|n\rangle & =\mathrm{e}^{-\frac{\mathrm{i}}{\hbar} E_{n} t} \delta_{n n^{\prime}} ;  \tag{6.49}\\
\left\langle n^{\prime}\right| U^{(1)}(t)|n\rangle & =\frac{\mathrm{e}^{-\frac{\mathrm{i}}{\hbar} E_{n^{\prime}} t}}{\mathrm{i} \hbar} \int_{0}^{t} \mathrm{~d} t_{1} \mathrm{e}^{\frac{\mathrm{i}}{\hbar}\left(E_{n^{\prime}}-E_{n}\right) t_{1}}\left\langle n^{\prime}\right| H^{\prime}\left(t_{1}\right)|n\rangle ;  \tag{6.50}\\
\left\langle n^{\prime}\right| U^{(2)}(t)|n\rangle & =\sum_{n_{1}} \frac{\mathrm{e}^{-\frac{\mathrm{i}}{\hbar} E_{n^{\prime}} t}}{\mathrm{i}^{2} \hbar^{2}} \int_{0}^{t} \int_{0}^{t_{1}} \mathrm{~d} t_{1} \mathrm{~d} t_{2} \mathrm{e}^{\frac{\mathrm{i}}{\hbar}\left(\left(E_{n^{\prime}}-E_{n_{1}}\right) t_{1}+\left(E_{n_{1}}-E_{n}\right) t_{2}\right]} \times \\
& \times\left\langle n^{\prime}\right| H^{\prime}\left(t_{1}\right)\left|n_{1}\right\rangle\left\langle n_{1}\right| H^{\prime}\left(t_{2}\right)|n\rangle . \tag{6.51}
\end{align*}
$$

The last expressions, as well as the general term (6.48) can be nicely interpreted by diagrams Fig. 6.1. In fact, it turns out that the perturbation in the expression for $U^{(p)}$ acts only in the instances $t_{1}, \ldots, t_{p}$, and meanwhile the system evolves as isolated. Consequently, the state $|i\rangle$ during the interval $\left(t_{i}, t_{i+1}\right)$ obtains pure phase factor. At the instance $t_{i}$, affected by the perturbation $H^{\prime}\left(t_{i}\right)$, the system abruptly changes state from $\left|n_{i-1}\right\rangle$ to $\left|n_{i}\right\rangle$, not conserving its energy (of isolated system). Thus, for $p \geq 2$ there appear intermediate states $\left|n_{1}\right\rangle, \ldots,\left|n_{p-1}\right\rangle$, which are called virtual states. Transitions to all unperturbed eigenstates at any instance from $[0, t]$ are included by summations over the unperturbed basis, and integration over all intermediate instances.


Figure 6.1: Scheme of the terms of the transition amplitudes: wave lines represent perturbation acting in the instances $t_{i}$ only, and causing system transition from the state $\left|n_{i-1}\right\rangle$ to $\left|n_{i}\right\rangle$, while meanwhile the system evolves as isolated.

### 6.5.3 Transition probabilities

The preceding results enable to expand the transition probabilities. Substituting (6.46) in (6.42) we find

$$
\begin{equation*}
v_{n \rightarrow n^{\prime}}(t)=\sum_{p, p^{\prime}=0}^{\infty}\left\langle n^{\prime}\right| U^{(p)}(t)|n\rangle\left\langle n^{\prime}\right| U^{\left(p^{\prime}\right)}(t)|n\rangle^{*} . \tag{6.52}
\end{equation*}
$$

Using (6.49) we find the first two terms in the transition probabilities:

$$
\begin{align*}
v_{n \rightarrow n^{\prime}}^{(0)}(t) & =\delta_{n n^{\prime}} ;  \tag{6.53}\\
v_{n \rightarrow n^{\prime}}^{(1)}(t) & =\delta_{n n^{\prime}} \frac{2}{\hbar} \operatorname{Im}\left[\mathrm{e}^{\frac{\mathrm{i}}{\hbar}\left(E_{n}-E_{n^{\prime}}\right) t} \int_{0}^{t} \mathrm{~d} t_{1} \mathrm{e}^{\frac{1}{\hbar}\left(E_{n^{\prime}}-E_{n}\right) t_{1}}\left\langle n^{\prime}\right| H^{\prime}\left(t_{1}\right)|n\rangle\right] ;  \tag{6.54}\\
v_{n \rightarrow n^{\prime}}^{(2)}(t) & =v_{n \rightarrow n^{\prime}}^{(2)}(t)+\delta_{n n^{\prime}} v_{n \Rightarrow n^{\prime}}^{(2)}(t) . \tag{6.55}
\end{align*}
$$

Obviously, in the zeroth and the first orders, probabilities forbid transitions (to other states then the initial one $|n\rangle$ ). Only in the second order, besides the state preserving term $v_{n \neq n}^{(2)}(t)=2 \operatorname{Re}\left[\langle n| U^{(2)}(t)|n\rangle\langle n| U^{(0)}(t)|n\rangle^{*}\right]$, there is a probability of the transition to other states. Introducing Bohr's frequencies $\omega_{n^{\prime} n}=\frac{E_{n^{\prime}}-E_{n}}{\hbar}$ one finds the transition probability (in the first effective order):

$$
\begin{equation*}
\left.v_{n \rightarrow n^{\prime}}^{(2)}(t)=\frac{1}{\hbar^{2}}\left|\int_{0}^{t} \mathrm{~d} t_{1} \mathrm{e}^{\mathrm{i} \omega_{n^{\prime} n} t_{1}}\left\langle n^{\prime}\right| H^{\prime}(t)\right| n\right\rangle\left.\right|^{2} . \tag{6.56}
\end{equation*}
$$

As $H^{\prime}(t)$ is hermitian, $v_{n \rightarrow n^{\prime}}^{(2)}(t)=v_{n^{\prime} \rightarrow n}^{(2)}(t)$, i.e. the opposite transitions occur with same probability. This characteristic of the first order transition probability remains in the higher orders only for the systems with time reversal symmetry.

## Constant perturbation and Fermi golden rule

When $H^{\prime}$ is constant during the considered time interval, the matrix elements can be pulled out of the integral, and the remaining time dependent part is easily integrated


Figure 6.2: Function $f(\omega, t)$ : note its periodicity of $2 \pi / t$ dumped by the enveloping function $4 / \omega^{2}$. Bellow: intervals of frequency corresponding to Eqns. (6.62) and (6.61).
giving squared "cardinal sine" function:

$$
\begin{equation*}
f(\omega, t)=\left|\int_{0}^{t} \mathrm{~d} t_{1} \mathrm{e}^{\mathrm{i} \omega t_{1}}\right|^{2}=2 \frac{1-\cos \omega t}{\omega^{2}}=\left(\frac{\sin t \frac{\omega}{2}}{\frac{\omega}{2}}\right)^{2} . \tag{6.57a}
\end{equation*}
$$

Two asymptotic properties of this function (Fig. 6.2) are important for further analysis:

$$
\begin{array}{r}
\int_{-\infty}^{\infty} f(\omega, t) \mathrm{d} \omega=2 \pi t \\
f(\omega, t \rightarrow \infty)=2 \pi t \delta(\omega) \tag{6.57c}
\end{array}
$$

In the first (effective) order, the transition probability becomes:

$$
\begin{equation*}
\left.v_{n \rightarrow n^{\prime}}^{(2)}=\frac{1}{\hbar^{2}}\left|\left\langle n^{\prime}\right| H^{\prime}\right| n\right\rangle\left.\right|^{2} f\left(\omega_{n^{\prime} n}, t\right) . \tag{6.58}
\end{equation*}
$$

For experimental applications, it is useful to consider the probability of transition not to a single state, but to a group of states $N^{\prime}$. Then this probability is sum of the probabilities for each of these states:

$$
\begin{equation*}
\left.v_{n \rightarrow N^{\prime}}^{(2)}=\sum_{n^{\prime} \in N^{\prime}} \frac{1}{\hbar^{2}}\left|\left\langle n^{\prime}\right| H^{\prime}\right| n\right\rangle\left.\right|^{2} f\left(\omega_{n^{\prime} n}, t\right) . \tag{6.59}
\end{equation*}
$$

Particularly frequent is the situation when the energies (eigen-energies of $H^{(0)}$ ) of the final states are (quasi)continually distributed within the interval ( $E_{0}-\varepsilon / 2, E_{0}+\varepsilon / 2$ ). Then the last sum becomes integral with the contribution of each energy $E$ pondered by its weight, i.e. by the density of states $\rho(E)$ :

$$
\begin{equation*}
\left.v_{n \rightarrow N^{\prime}}^{(2)}=\frac{1}{\hbar^{2}} \int_{E_{0}-\varepsilon / 2}^{E_{0}+\varepsilon / 2}\left|\left\langle n^{\prime}\right| H^{\prime}\right| n\right\rangle\left.\right|^{2} f\left(\frac{E-E_{n}}{\hbar}, t\right) \rho(E) \mathrm{d} E . \tag{6.60}
\end{equation*}
$$

For sufficiently small $\varepsilon$ the matrix elements and density may be considered almost constant, and taken outside the integral. On the other side, for large $t$ (but still within the applicability of the approximation!), $\varepsilon \gg 2 \pi \hbar / t$, and the interval of integration contains a number of periods $2 \pi / t$ of $f(\omega, t)$. Then two opposite cases can be analyzed.
(i) $E_{n} \in\left(E_{0}-\varepsilon / 2, E_{0}+\varepsilon / 2\right)$ (Fig. 6.2). Then, since the contribution of the central peak is highly dominant (according to (6.57c)), the integration may be approximately extended to the whole energy axis, and (6.57b) is used to obtain:

$$
\begin{equation*}
\left.v_{n \rightarrow N^{\prime}}^{(2)}=\frac{2 \pi}{\hbar}\left|\left\langle n^{\prime}\right| H^{\prime}\right| n\right\rangle\left.\right|^{2} \rho\left(E_{n}\right) t . \tag{6.61}
\end{equation*}
$$

This conclusion, that the transition probability to the nearby states linearly increases with time is known as the Fermi's golden rule.
(ii) $E_{n} \notin\left(E_{0}-\varepsilon / 2, E_{0}+\varepsilon / 2\right)$ (Fig. 6.2). Now the central peak is excluded, but the integral may be approximated by its mean over a number of periods. As the mean of the denominator is $\left(E_{0}-E_{n}\right)^{2}$, while the mean of $2(1-\cos \omega t)$ over multiple period $\left[2 \pi k / t, 2 \pi\left(k+k^{\prime}\right) / t\right]$ ( $k$ and $k^{\prime}$ are integers) is 2 , the estimate of the transition amplitude becomes:

$$
\begin{equation*}
v_{n \rightarrow N^{\prime}}^{(2)}=\frac{\left.2 \varepsilon\left|\left\langle n^{\prime}\right| H^{\prime}\right| n\right\rangle\left.\right|^{2}}{\left(E_{0}-E_{n}\right)^{2}} \rho\left(E_{0}\right) \tag{6.62}
\end{equation*}
$$

To conclude, the last equation shows that the transition rate per unit time is constant for the transitions conserving the energy of the isolated system, while it vanishes ( $t$ is large) if the energy is changed.

## Periodic perturbation and resonances

Another frequent situation is that the perturbation is periodic in time (e.g. optical transitions induced by the electromagnetic waves). Then its general form is

$$
\begin{equation*}
H^{\prime}(t)=A \mathrm{e}^{\mathrm{i} \omega t}+A^{\dagger} \mathrm{e}^{-\mathrm{i} \omega t} \tag{6.63}
\end{equation*}
$$

where $A$ is time independent operator. Instead of the complete matrix elements, only time independent part can be put outside the integral in (6.56), resulting in

$$
\begin{equation*}
\left.v_{n \rightarrow n^{\prime}}^{(2)}=\frac{1}{\hbar^{2}}\left|\left\langle n^{\prime}\right| A\right| n\right\rangle \int_{0}^{t} \mathrm{e}^{\mathrm{i}\left(\omega_{n^{\prime} n}+\omega\right) t_{1}} \mathrm{~d} t_{1}+\left.\left\langle n^{\prime}\right| A^{\dagger}|n\rangle \int_{0}^{t} \mathrm{e}^{\mathrm{i}\left(\omega_{n^{\prime} n}-\omega\right) t_{1}} \mathrm{~d} t_{1}\right|^{2} . \tag{6.64}
\end{equation*}
$$

Again, the case when $t$ is sufficiently large is considered. Then, the first term in the probability amplitude is negligible unless $E_{n^{\prime}}=E_{n}-\hbar \omega$, i.e. only the process of emitting of the energy quant $\hbar \omega$ gives significant contribution to the transition probability. Analogously, the second term describes dominantly the absorption of the same quant: $E_{n^{\prime}}=E_{n}+\hbar \omega$. Under the same assumption of large $t, t \gg 2 \pi / \omega$, these two regions are separated. Therefore, the amplitudes of emission and absorption do not interfere, and the probability is large only for these two processes, which can be analyzed separately. For emission, one gets analogously to (6.58):

$$
\begin{equation*}
\left.\left.v_{n \rightarrow n^{\prime}}^{(2) e}=\frac{1}{\hbar^{2}}\left|\left\langle n^{\prime}\right| A\right| n\right\rangle\left.\right|^{2} f\left(\omega_{n^{\prime} n}+\omega, t\right), \quad v_{n \nrightarrow n^{\prime}}^{(2) \mathrm{a}}=\frac{1}{\hbar^{2}}\left|\left\langle n^{\prime}\right| A^{\dagger}\right| n\right\rangle\left.\right|^{2} f\left(\omega_{n^{\prime} n}-\omega, t\right) . \tag{6.65}
\end{equation*}
$$

This obviously differs with respect to the constant perturbation only in the shift for $\pm \omega$ of the function $f(\omega, t)$. Therefore, all the conclusions are analogous to that case: instead of the conservation of energy (corresponding to the central peak of $f(\omega, t)$ ) the transition probability per unit time to the group of states around $E_{n} \mp \omega$ is constant, while vanishes for different final energies. All these results are easily generalized for the more complex periodic perturbations, having the Fourier form

$$
H^{\prime}(t)=\sum_{s} A_{s} \mathrm{e}^{\mathrm{i} s \omega t}+A_{s}^{\dagger} \mathrm{e}^{-\mathrm{i} s \omega t}
$$

each term contributes to the emission/absorption of the quant $s \hbar \omega$.

### 6.6 Elementary scattering theory

Collisions of some sort of the particles with a target is one of the most frequent experimental methods to examin properties of target and interactions. The scheme of the collision is simple (Fig. ??): a collimated beam of particles is directed toward the target. After that the scattered particles are detected and studied, and according to this many conclusions about the target are derived. There are various types of the collision processes: spallation and reactions (when some completely new particles are scattered), ZAHVAT (the target and the incoming particle together form a new particle) and scattering (the simplest one, when the outgoing particle is the same one as incoming). Particularly, here we are interested in the elastic scattering, in which the kinetic energy of the incoming particle is preserved.

In the experiments with elastic scattering the distribution $\mathrm{d} n(\theta, \varphi)$ of the outgoing particles over the directions is measured. This is obviously proportional to the flux $\Phi_{\text {in }}$ of the incoming (along the $z$-axis) particles:

$$
\begin{equation*}
\mathrm{d} n(\theta, \varphi)=\sigma(\theta, \varphi) \Phi_{\mathrm{in}} \mathrm{~d} \Omega \tag{6.66}
\end{equation*}
$$

where the differential cross section $\sigma(\theta, \varphi)$ is the proportionality factor, which comprises all the details on the types of the beam and target, interactions, etc. Note that $\sigma(0, \varphi)$ is not measured quantity, since experimentally it is not easy to distinguish between the scattered particles with $\theta=0$ scattering angle, and the particles which simply passed by target without interaction (thus in the same direction). If the interaction between the beam and target depends only on their distance, then it is spherically symmetric, and the cross section is $\varphi$-independent. Therefore, it can be integrated over $\varphi$, giving $\sigma(\theta)=\int_{0}^{2 \pi} \sigma(\theta, \varphi) \mathrm{d} \varphi=2 \pi \sigma(\theta, 0)$. Total cross section is the integral $\sigma=\int_{0}^{\pi} \int_{0}^{2 \pi} \sigma(\theta, \varphi) \mathrm{d} \Omega$. Cross sections have the dimension of surface, and usually are measured in Barns: 1Barn $=10^{-24} \mathrm{~cm}^{2}$.

Within the quantum theory, it is possible to give some general relations on the scattering cross section, without detailed analysis of the interaction. To this end the stationary regime is to be considered, when the state of beam is a stationary state $\psi(\boldsymbol{r}, t)=\psi(\boldsymbol{r}) \mathrm{e}^{-\mathrm{i} E t / \hbar}$ of the hamiltonian $H=T+V(r)$ describing a beam particle interacting with target only. The interaction potential is assumed to decrease with $r$. The
time independent function is of the form (it is not normalized)

$$
\begin{equation*}
\psi(\boldsymbol{r})=\mathrm{e}^{\mathrm{i} k z}+f(\theta, \varphi) \frac{\mathrm{e}^{\mathrm{i} k r}}{r} \tag{6.67}
\end{equation*}
$$

being a superposition of the incoming state (plane wave along $z$-axis) and the scattered state (spherical wave). It asymptotically (for large $r$ ) satisfies time independent Schrödinger equation. The function $f(\theta, \varphi)$ (or $f(\theta)$ for the spherically symmetric interaction) is called scattering amplitude. The number of scattered particles is given by (6.66), with the incoming beam particles in the state of plane wave. Thus their flux is proportional to the probability current $\Phi_{\text {in }}=c \hbar k / m$. On the other side this number is proportional to the flux of outgoing particles:

$$
\mathrm{d} n(\theta, \varphi)=c \boldsymbol{j}^{\text {out }} \cdot \mathrm{d} \boldsymbol{S}=c j_{r}^{\text {out }} r^{2} \mathrm{~d} \Omega
$$

The outgoing probability current is easily calculated applying (2.19) to the outgoing state $\psi_{\mathrm{sw}}(\boldsymbol{r})=f(\theta, \varphi) \mathrm{e}^{\mathrm{i} k r} / r$. The leading term in $r$ is $r^{-2}$, and in the asymptotic region (which is studied here) only $r$-component of current appears:

$$
j_{r}^{\text {out }}=\frac{1}{m} \operatorname{Re}\left(\psi^{*}(\boldsymbol{r}, t) \frac{\hbar}{\mathrm{i}} \partial_{r} \psi(\boldsymbol{r}, t)\right)=\frac{\hbar}{m} k \frac{|f(\theta, \varphi)|^{2}}{r^{2}} .
$$

Using these two expressions for the number of outgoing particles one finally finds the differential cross section:

$$
\begin{equation*}
\sigma(\theta, \varphi)=|f(\theta, \varphi)|^{2} \tag{6.68}
\end{equation*}
$$

## Appendix A

## Technical support

## A. 1 Separation of Variables

Here we give simple algebraic spectral theorem on the operators acting in the direct product space $\mathcal{S}=\mathcal{S}_{1} \otimes \mathcal{S}_{2}$. The aim is to facilitate the solution of the eigenproblem, reducing it to the factor spaces for two special, but frequent in quantum mechanical problems forms of the considered operators: $A=B_{1} \otimes B_{2}$ and $A=B_{1} \otimes B_{2}+C_{1} \otimes C_{2}$. While the former is given for the completeness only, in the view of its self obvious outcome and considerably rare physical situations when it is applicable, the latter one refers to wide range of examples.

Lema A. 1 (Separation of variables) Let $\mathcal{S}=\mathcal{S}_{1} \otimes \mathcal{S}_{2}$.
(a) If $A=B_{1} \otimes B_{2}$ with spectral forms $B_{i}=\sum_{b_{i}} b_{i} P_{b_{i}}\left(B_{i}\right)(i=1,2)$, where $P_{b_{i}}\left(B_{i}\right)$ are the spectral projectors of the factor operators, then the spectral form of $A$ is:

$$
\begin{equation*}
A=\sum_{b_{1}, b_{2}} b_{1} b_{2} P_{b_{1}}\left(B_{1}\right) \otimes P_{b_{2}}\left(B_{2}\right) . \tag{A.1}
\end{equation*}
$$

Geometrically, this means that the space $\mathcal{S}$ is decomposed as the direct sum $\mathcal{S}=$ $\oplus_{b_{1}, b_{2}} \mathcal{S}_{1}^{\left(b_{1}\right)} \otimes \mathcal{S}_{2}^{\left(b_{2}\right)}$ of the eigenspaces $\mathcal{S}_{1}^{\left(b_{1}\right)} \otimes \mathcal{S}_{2}^{\left(b_{2}\right)}$ of $A$ for the eigenvalues $b_{1} b_{2}$ (not necessarily different; hence, to achieve the spectral decomposition one should rearrange the sum firstly, to group together the same eigenvalues), each of these subspaces being the product of the eigenspaces of the factor operators.
(b) If $A=B_{1} \otimes B_{2}+C_{1} \otimes C_{2}$, and $B_{2}$ and $C_{2}$ are compatible, and the projectors to their common eigensubspaces $\mathcal{S}_{2}^{\left(b_{2}, c_{2}\right)}$ are $P_{\left(b_{2}, c_{2}\right)}$, then

$$
\begin{equation*}
A=\sum_{\left(b_{2}, c_{2}\right)} \sum_{a\left(b_{2}, c_{2}\right)} a\left(b_{2}, c_{2}\right) P_{a\left(b_{2}, c_{2}\right)}\left(b_{2} B_{1}+c_{2} C_{1}\right) \otimes P_{b_{2} c_{2}} \tag{A.2}
\end{equation*}
$$

In other words, $\mathcal{S}$ is decomposed to the direct sum $\mathcal{S}=\sum_{\left(b_{2}, c_{2}\right)} \sum_{a\left(b_{2}, c_{2}\right)} \mathcal{S}_{1}^{a\left(b_{2}, c_{2}\right)} \otimes$ $\mathcal{S}_{2}^{\left(b_{2}, c_{2}\right)}$ of the products of the eigenspaces $\mathcal{S}_{1}^{a\left(b_{2}, c_{2}\right)}$ for the eigenvalues a $\left(b_{2}, c_{2}\right)$ (again, not necessarily different) of the operators $A_{1}\left(b_{2}, c_{2}\right)=b_{2} B_{1}+c_{2} C_{1}$ and common eigenspaces $\mathcal{S}_{2}^{\left(b_{2}, c_{2}\right)}$ of $B_{2}$ and $C_{2}$.
=Proof: (a) The direct products $\left|b_{1} ; \lambda_{1}\right\rangle \otimes\left|b_{1} ; \lambda_{2}\right\rangle$ of the eigenvectors of $B_{1}$ and $B_{2}$ are basis in $\mathcal{S}$. Thus they are an eigenbasis of $A$, since $A\left|b_{1} ; \lambda_{1}\right\rangle \otimes\left|b_{1} ; \lambda_{2}\right\rangle=b_{1} b_{2}\left|b_{1} ; \lambda_{1}\right\rangle \otimes\left|b_{1} ; \lambda_{2}\right\rangle$. Finally, for fixed $b_{1}$ and $b_{2}$ they span the subspaces $\mathcal{S}_{1}^{\left(b_{1}\right)} \otimes \mathcal{S}_{2}^{\left(b_{2}\right)}$ in the decomposition $\mathcal{S}=\oplus b_{1}, b_{2} \mathcal{S}_{1}^{\left(b_{1}\right)} \otimes \mathcal{S}_{2}^{\left(b_{2}\right)}$.
(b) Operator $A$ is reduced to $A_{1}\left(b_{2}, c_{2}\right) \otimes P_{b_{2} c_{2}}$ in each of the spaces $\mathcal{S}_{1} \otimes \mathcal{S}_{2}^{\left(b_{2}, c_{2}\right)}$, since $A|x\rangle\left|\left(b_{2}, c_{2}\right), \lambda\right\rangle=$ $\left(\left(b_{2} B_{1}+c_{2} C_{1}\right)|x\rangle\right) \otimes\left|\left(b_{2}, c_{2}\right), \lambda\right\rangle$ for each $|x\rangle$ from $\mathcal{S}_{1}$ and $\left|\left(b_{2}, c_{2}\right), \lambda\right\rangle$ from $\mathcal{S}_{2}^{\left(b_{2}, c_{2}\right)}$. Therefore, it remains to find the spectral form of $A_{1}\left(b_{2}, c_{2}\right)$ in $\mathcal{S}_{1}$ for each compatible pair $\left(b_{2}, c_{2}\right)$.

## A. 2 Hypergeometric equation

The most important equations of quantum dynamics are second order differential one. Among them, the class of the generalized hypergeometric equation covers most of the problems considered in this textbook:

$$
\begin{equation*}
u^{\prime \prime}(t)+\frac{\tilde{\tau}(t)}{\sigma(t)} u^{\prime}(t)+\frac{\tilde{\sigma}(t)}{\sigma(t)^{2}} u(t)=0 \tag{A.3}
\end{equation*}
$$

where $\sigma(t)$ and $\tilde{\sigma}(t)$ are polynomials of the order less then 3 , and $\tilde{\tau}(t)$ a polynomial of the order less than 2. This class is solvable, with the solution standardized within so called special functions. Here we give the three step algorithm to find out these solutions. The method is based on the theory of differential operators in the underlying Lebesgue space $\mathcal{L}([a, b])$.

Step 1. Find all possible constants $k$ such that

$$
\pi(t)=\frac{\sigma^{\prime}-\tilde{\tau}}{2} \pm \sqrt{\left(\frac{\sigma^{\prime}-\tilde{\tau}}{2}\right)^{2}-\tilde{\sigma}+k \sigma}
$$

is polynomial (of the order less than 2). This means that we look for $k$ giving the rooted expression the form $(a t+b)^{2}$. For each (at most 2) such $k$, we find two polynomials $\pi_{k \pm}$, and the next steps are to be performed for each of them.

Step 2. For each $\pi_{k \pm}$ solve the first order differential equation in $\phi(t)$ :

$$
\frac{\phi^{\prime}}{\phi}=\frac{\pi_{k \pm}}{\sigma}
$$

and calculate polynomials $\tau_{k \pm}=2 \pi_{k \pm}-\widetilde{\tau}$ and constants $\lambda_{k \pm}=-k-\pi_{k \pm}^{\prime}$. Substituting the solutions $\phi_{k \pm}$ in (A.3) by the change $u=\phi_{k \pm} y_{k \pm}$ we get hypergeometric equation in $y=y_{k \pm}$ :

$$
\begin{equation*}
\sigma(t) y^{\prime \prime}(t)+\tau(t) y^{\prime}(t)=\lambda y(t) \tag{A.4}
\end{equation*}
$$

which obviously has the eigenproblem form.
Step 3. Solve Pearson's equation

$$
(\sigma \rho)^{\prime}=\tau_{k \pm} \rho, \text { i.e. } \frac{\rho^{\prime}}{\rho}=\frac{\tau_{k \pm}-\sigma^{\prime}}{\sigma}
$$

for each $k$, and among the solutions $\rho_{k \pm}$ select only those being positive (perhaps vanishing on a countable set) and bounded. Then (A.4) can be rewritten in the eigenequation form:

$$
\begin{equation*}
H y(t)=\lambda y(t), \quad \text { with } H=\frac{1}{\rho(t)} \frac{d}{d t}\left[\sigma(t) \rho(t) \frac{d}{d t}\right] \tag{A.5}
\end{equation*}
$$

The rest of the algorithm is based on the
Theorem A. 1 In the subspace $\left\{f(t) \in \mathcal{L}^{2}((a, b), \rho) \mid \sigma(a) \rho(a) f(a)=\sigma(b) \rho(b) f(b)=0\right\}$ of $\mathcal{L}^{2}((a, b), \rho)$, with $\rho$ being bounded positive function on $(a, b)$ satisfying $(\sigma \rho)^{\prime}=\tau \rho$, operator $H$ has purely discrete spectrum $\left\{\left.\lambda_{n}=n \tau^{\prime}+\frac{n(n-1)}{2} \sigma^{\prime \prime} \right\rvert\, n=0,1, \ldots\right\}$ with the corresponding eigenfunctions being orthogonal polynomial given by Rodrigues equation $x_{n}(t)=\frac{B_{n}}{\rho(t)} \frac{d^{n}}{d t^{n}}\left[\sigma^{n}(t) \rho(t)\right]$ (here $B_{n}$ is arbotrary constant, to be determined by the normalization).

To resume, only for $\lambda=\lambda_{n}$ (for some $n$ ) (A.3) has solution $u=\phi x_{n}$ satisfying the conditions of the theorem, which is unique (up to the constant). For other values of $\lambda$ there are no solutions (at least of this type).

It is easy to verify that the obtained functions $x_{n}(t)$ are polynomials of the order $n$. Therefore, they can be obtained by the orthonormalization of the set $\left\{1, t, t^{2}, \ldots\right\}$ in the space $\mathcal{L}^{2}((a, b), \rho)$. These classical orthogonal polynomials are well studied and standardized (see Table A. 1 and [?] for details) according to $\sigma$ and $\tau$.

Table A.1: Classical orthogonal polynomials. Among Jacobi's polynomials are Legendre's polynomials $P_{n}(t)=P_{n}^{(0,0)}(t)$ and associated Legendre's polynomials $P_{n}^{(m)}(t)=P_{n}^{(m, m)}(t)$.

| Name | $(a, b)$ | $\sigma(t)$ | $\tau(t)$ | $\rho(t)$ | $x_{n}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Jacobi | $(-1,1)$ | $1-t^{2}$ | $\begin{aligned} & \beta-\alpha- \\ & -(\alpha+\beta+2) t \end{aligned}$ | $\begin{aligned} & \hline \hline(1-t)^{\alpha} \times \\ & \times(1+t)^{\beta} \end{aligned}$ | $\begin{aligned} & \hline P_{n}^{(\alpha, \beta)}(t)=\frac{(-1)^{n} \times}{2^{n} n!(1-t)^{\alpha \alpha}(1+t)} \\ & \frac{d^{n}}{d t^{\beta}}\left[(1-t)^{n+\alpha}(1+t)^{n+\beta}\right] \end{aligned}$ |
| Laguerre | $(0, \infty)$ | $t$ | $-t+\alpha+1$ | $t^{\alpha} e^{-t}$ | $L_{n}^{\alpha}(t)=\frac{e^{t} t^{-\alpha}}{n!} \frac{d^{n}}{d t^{n}}\left[e^{-t} t^{n+\alpha}\right]$ |
| Hermite | $(-\infty, \infty)$ | 1 | $-2 t$ | $e^{-t^{2}}$ | $H_{n}(t)=\frac{(-1)^{n}}{\sqrt{2^{n} n!\sqrt{\pi}}} e^{t^{2}} \frac{d^{n}}{d t^{n}}\left[e^{-t^{2}}\right]$ |

## Appendix B

## Solutions of Exercises

## B.0.1 Quantum Kinematics

Solution ??. The following equalities are valid: $[A B, C]=A B C-C A B=A B C-$ $A C B+A C B-C A B=A(B C-C B)+(A C-C A) B=A[B, C]+[A, C] B$.

Solution ??. Writing expression as series in powers of $\lambda$ we have: $(A-\lambda B)^{-1}=$ $\sum_{n=0}^{\infty} \lambda^{n} L_{n}$, where operators $L_{n}$ are to be determined. Multiplying by $(A-\lambda B)$ one gets $I=\sum_{n=0}^{\infty} \lambda^{n}(A-\lambda B) L_{n}=A L_{0}+\sum_{n=0}^{\infty} \lambda^{n+1}\left(A L_{n+1}-B L_{n}\right)$. It is obvious that expression becomes $(A-\lambda B)^{-1}=\sum_{n=0}^{\infty} \lambda^{n}\left(A^{-1} B\right)^{n} A^{-1}=A^{-1} \sum_{n=0}^{\infty} \lambda^{n}\left(B A^{-1}\right)^{n}$.

Solution ??. $\quad\left\langle(A-\langle A\rangle)^{2}\right\rangle=\left\langle A^{2}-2 A\langle A\rangle+\langle A\rangle^{2}\right\rangle=\left\langle A^{2}\right\rangle-2\langle A\rangle\langle A\rangle+\langle A\rangle^{2}=\left\langle A^{2}\right\rangle-$ $\langle A\rangle^{2}$

Solution ??. For $m=1$ equality holds. Supposing that equality is valid for some $m=n$, for $m=n+1$ using solution ??: $\left[A^{n+1}, B\right]=A\left[A^{n}, B\right]+[A, B] A^{n}=n A^{n}[A, B]+[A, B] A^{n}$, using condition that operator $A$ (and any its positive power) commutes with commutator $[A, B]$, we have $\left[A^{n+1}, B\right]=(n+1) A^{n}[A, B]$.

Solution ??. Assuming that orthonormal basis $\beta=\{|i\rangle\}$ exists, trace can be expressed as $\operatorname{Tr}(|u\rangle\langle v|)=\sum\langle i \mid u\rangle\langle v \mid i\rangle=\sum\langle v \mid i\rangle\langle i \mid u\rangle=\langle v|\left(\sum|i\rangle\langle i|\right)|u\rangle=\langle v| I|u\rangle=\langle v \mid u\rangle$.

Solution ??. Operator $H$ is Hermitian iff $(x, H y)=(H x, y)$ holds for any two vectors $x$ and $y$. To shorten notation we'll write $A^{\dagger}=A^{+}$and $A=A^{-}$. The following set of equalities for any $x$ and $y\left(x, A^{ \pm} A^{\mp} y\right)=\left(A^{\mp} x, A^{\mp} y\right)=\left(A^{ \pm} A^{\mp} x, y\right)$ proves that operators $A^{ \pm} A^{\mp}$ are Hermitian. For any vector $x$, mean value $\left(x, A^{ \pm} A^{\mp} x\right)=\left(A^{\mp} x, A^{\mp} x\right)=\left\|A^{\mp} x\right\|^{2}$ is always positive, and for any orthonormal basis $\beta=\{|i\rangle\}, \operatorname{Tr} A^{ \pm} A^{\mp}=\sum\left(i, A^{ \pm} A^{\mp} i\right)=$ $\sum\left(A^{\mp} i, A^{\mp} i\right)=\sum\left\|A^{\mp} i\right\|^{2}$. Finally, $\operatorname{Tr} A^{ \pm} A^{\mp}=\sum\left\|A^{\mp} i\right\|^{2}=0$ iff $\left\|A^{\mp} i\right\|=0$ for any $i \Rightarrow A^{\mp}=0$.

Solution ??. $\mathrm{d}(A B) / \mathrm{d} \lambda=\lim _{\varepsilon \rightarrow 0}(A(\lambda+\varepsilon) B(\lambda+\varepsilon)-A(\lambda) B(\lambda)) / \varepsilon=\lim _{\varepsilon \rightarrow 0}(A(\lambda+$ ع) $B(\lambda+\varepsilon)-A(\lambda) B(\lambda+\varepsilon)+A(\lambda) B(\lambda+\varepsilon)-A(\lambda) B(\lambda)) / \varepsilon=\lim _{\varepsilon \rightarrow 0}(A(\lambda+\varepsilon)-A(\lambda)) / \varepsilon B(\lambda+$ $\varepsilon)+\lim _{\varepsilon \rightarrow 0} A(\lambda)(B(\lambda+\varepsilon)-B(\lambda)) / \varepsilon=(\mathrm{d}(A) / \mathrm{d} \lambda) B+A(\mathrm{~d}(B) / \mathrm{d} \lambda)$. Differentiating equality $A A^{-1}=I$ one gets $(\mathrm{d}(A) / \mathrm{d} \lambda) A^{-1}+A\left(\mathrm{~d}\left(A^{-1}\right) / \mathrm{d} \lambda\right)=0 \Rightarrow \mathrm{~d}\left(A^{-1}\right) / \mathrm{d} \lambda=$ $-A^{-1}(\mathrm{~d}(A) / \mathrm{d} \lambda) A^{-1}$.

Solution ??. Introducing operator $g(A, B, \alpha)=e^{\alpha B} A e^{-\alpha B}$ and expanding it around $\alpha=0$ one gets $g(A, B, \alpha)=\sum \alpha^{n} \mathrm{~d}^{n} g / \mathrm{d} \alpha^{n} / n!$ where $\mathrm{d}^{0} g / \mathrm{d} \alpha^{0}=A, \mathrm{~d} g / \mathrm{d} \alpha=B g-g B=$ $[B, A], \mathrm{d}^{2} g / \mathrm{d} \alpha^{2}=[B, \mathrm{~d} g / \mathrm{d} \alpha]=[B,[B, A]], \ldots$ Setting $\alpha=1$ one gets the equality.

Solution ??. Using the solution of the previous problem with conditions given in this problem one gets: $\mathrm{d} T(s) / \mathrm{d} s=\mathrm{d}\left(e^{s A} e^{s B}\right) / \mathrm{d} s=A T(s)+T(s) B=\left(A+e^{s A} B e^{-s A}\right) T(s)=$ $(A+B+[A, B] s) T(s) \Rightarrow T(s)=\mathrm{e}^{(A+B) s+\frac{1}{2}[A, B] s^{2}}$.

Solution 1.1. For an arbitrary vector $|\psi\rangle$ its image under $\rho$ is $\rho|\psi\rangle=\sum_{i} \omega_{i}\left|\psi_{i}\right\rangle\left\langle\psi_{i} \mid \psi\right\rangle$. Thus, any vector from the range of $\rho$ is a linear combination of the involved states $\left|\psi_{i}\right\rangle$, and $\mathcal{R}(\rho)<\mathcal{V} \stackrel{\text { def }}{=}$ span $\{|\psi\rangle\}$. If $\mathcal{R}(\rho)$ is a proper subspace in $\mathcal{V}$, there exists a nonzero vector $|x\rangle$ from $\mathcal{V}$ such that $\rho|x\rangle=0$. Then we can find an orthonormal basis $|j\rangle$ in $\mathcal{V}$, such that $|1\rangle=|x\rangle$. In this basis $\left|\psi_{i}\right\rangle=\sum_{j} c_{j i}|j\rangle$ and $\rho=\sum_{j j^{\prime}} \sum_{i} c_{j i} c_{j^{\prime} i}^{*} \omega_{i}|j\rangle\left\langle j^{\prime}\right|$, with $c_{j i}=\left\langle j \mid \psi_{i}\right\rangle$. Since $|1\rangle$ is an eigenvector of $\rho$ for the eigenvalue 0 , it spans onedimensional invariant subspace for $\rho$, meaning that for the matrix elements in the first row and the first column of $\rho$ vanish. Particularly, for the first diagonal matrix element one has $\sum_{i} c_{1 i} c_{1 i}^{*} \omega_{i}=0$, and in the view of positivity of $\omega_{i}$ this means that all the coefficients $c_{1 i}$ vanish, i.e. $|x\rangle=0$, which contradicts to the assumption.

Solution 1.2. From (1.13b) one finds superoperator $\Pi_{A} X \stackrel{\text { def }}{=} \sum_{a} P_{a}(A) X P_{a}(A)$. This superoperator is hermitian: $\left(X, \Pi_{A} Y\right)=\sum_{a} \operatorname{Tr} X^{\dagger} P_{a}(A) Y P_{a}(A)=\sum_{a} \operatorname{Tr} P_{a}(A) X^{\dagger} P_{a}(A) Y=$ $\operatorname{Tr}\left(\sum_{a} P_{a}(A) X P_{a}(A)\right)^{\dagger} Y=\left(\Pi_{A} X, Y\right)$. It is also idempotent:

$$
\Pi_{A} \Pi_{A} X=\sum_{a, b} P_{a}(A) P_{b}(A) X P_{b}(A) \Pi_{A}=\sum_{a} P_{a}(A) X \Pi_{A}=\Pi_{A} X
$$

Solution 1.4. Using partial integration one verifies: $(\psi, p \varphi)=\int_{a}^{b} \psi^{*} p \varphi \mathrm{~d} x=$ $\int_{a}^{b} \psi^{*}\left(-\mathrm{i} \hbar \frac{\mathrm{d}}{\mathrm{d} x} \varphi\right) \mathrm{d} x=\int_{a}^{b} \psi^{*}\left(-\mathrm{i} \hbar \varphi^{\prime}\right) \mathrm{d} x=-\left.\mathrm{i} \hbar \psi^{*} \varphi\right|_{a} ^{b}+\int_{a}^{b}\left(-\mathrm{i} \hbar \psi^{\prime}\right)^{*} \varphi \mathrm{~d} x=(p \psi, \varphi)$.
Solution 1.5. Applying obvious relation $[A, B C]=B[A, C]+[A, B] C$ successively in $\left[\hat{x}, \hat{p}^{n}\right]$ one finds series $\left[\hat{x}, \hat{p}^{n}\right]=\hat{p}^{n-1}[\hat{x}, \hat{p}]+\left[\hat{x}, \hat{p}^{n-1}\right] \hat{p}=\mathrm{i} \hbar \hat{p}^{n-1}+\hat{p}^{n-2}[\hat{x}, \hat{p}] \hat{p}+\left[\hat{x}, \hat{p}^{n-2}\right] \hat{p}^{2}=$ $2 \mathrm{i} \hbar \hat{p}^{n-1}+\left[\hat{x}, \hat{p}^{n-2}\right] \hat{p}^{2}=\cdots=n \mathrm{i} \hbar \hat{p}^{n-1}$.

Solution 1.6. ???
Solution 1.7. Since for each $|x\rangle$ one has $\langle x| A_{1}|x\rangle=\sum_{b}(\langle x| \otimes\langle b|) A(|x\rangle \otimes|b\rangle)$, the reality and positivity of the terms on the right implies reality and positivity of the left hand side. Finally, if $A$ is statistical, $\operatorname{Tr} A_{1}=\sum_{a b}(\langle a| \otimes\langle b|) A(|a\rangle \otimes|b\rangle)=\operatorname{Tr} A=1$.

Solution 1.8. Consider the general, i.e. mixed state $\rho$. Being positive operator, $\rho$ has unique positive root $\rho^{1 / 2}$. Let us define the auxiliary operators $A^{\prime}=A-\langle A\rangle_{\rho}$, $B^{\prime}=B-\langle B\rangle_{\rho}$, and $A "=\rho^{1 / 2} A^{\prime}, B^{\prime \prime}=\rho^{1 / 2} B^{\prime}$. Using the standard scalar product in the operators space $(X, Y)=\operatorname{Tr} A^{\dagger} B$, it is easy to recognize the Schwartz inequality in the expression $\operatorname{Tr}\left(A^{\prime \dagger} A^{\prime \prime}\right) \operatorname{Tr}\left(B^{\prime \prime} B^{\prime \prime}\right) \geq\left|\operatorname{Tr} B^{\prime \dagger} A^{\prime \prime}\right|^{2}$. Note that the left hand side is equal to $\Delta_{\rho}^{2}(A) \Delta_{\rho}^{2}(B)$, while the right one substituting $B$ " becomes mean $\left|\left\langle A^{\prime} B^{\prime}\right\rangle_{\rho}\right|^{2}=$ $\left|\left\langle\frac{A^{\prime} B^{\prime}+B^{\prime} A^{\prime}}{2}\right\rangle_{\rho}+\left\langle\frac{A^{\prime} B^{\prime}-B^{\prime} A^{\prime}}{2}\right\rangle_{\rho}\right|^{2}$. Recall that the obtained symmetrized and antisymmetrized
products are hermitian and scew-hermitian operators, and their means real and pure imaginary, respectively. Thus, the absolute value if the whole obtained expression is less that square of the imaginary part, or altogether, $\Delta_{\rho}^{2}(A) \Delta_{\rho}^{2}(B) \geq\left|\left\langle A^{\prime} B^{\prime}\right\rangle_{\rho}\right|^{2}=\frac{1}{4}\left|\left\langle\left[A^{\prime}, B^{\prime}\right]\right\rangle_{\rho}\right|^{2}=$ $\frac{1}{4}\left|\langle[A, B]\rangle_{\rho}\right|^{2}$.
Solution 1.9. ???

## B.0.2 Quantum dynamcs

Solution 2.1. Differentiating the expectation value one finds:

$$
\frac{\mathrm{d}}{\mathrm{~d} a} E(a)=\langle E(a)| \frac{\mathrm{d}}{\mathrm{~d} a} H(a)|E(a)\rangle+\left(\frac{\mathrm{d}}{\mathrm{~d} a}\langle E(a)|\right) H(a)|E(a)\rangle+\langle E(a)| H(a)\left(\frac{\mathrm{d}}{\mathrm{~d} a}|E(a)\rangle\right) .
$$

Therefore:

$$
\frac{\mathrm{d}}{\mathrm{~d} a} E(a)=\langle E(a)| \frac{\mathrm{d}}{\mathrm{~d} a} H(a)|E(a)\rangle+E(a)\left(\frac{\mathrm{d}}{\mathrm{~d} a}\langle E(a)|\right)|E(a)\rangle+E(a)\langle E(a)|\left(\frac{\mathrm{d}}{\mathrm{~d} a}|E(a)\rangle\right) .
$$

The last two terms cancel, as the derivative of the norm of the eigenstate, being 1.
Solution 2.2. ???
Solution 2.3. ???
Solution ??. It is easy to see that if energy $(E)$ is greater than any potential in infinity $\left(V_{ \pm}\right)$than eigenfunction is plane-wave with momentum $k=\sqrt{2 m\left(E-V_{ \pm}\right) / \hbar^{2}}$ without boundary condition in that infinity. This means that particle is free iff energy is greater than minimum of the potentials in infinity. In conclusion, the state is bound iff $V_{+}>E$.

Solution ??. Suppose that we have two physically different states (nonlinear) $\psi_{1}(x)$ and $\psi_{2}(x)$ heaving the same energy. Than, for any potential $V(x)$ we have $\psi_{1}^{\prime \prime} / \psi_{1}=2 m(E-$ $V(x)) / \hbar^{2}=\psi_{2}^{\prime \prime} / \psi_{2}$ leading to equation $\left(\psi_{1}^{\prime} \psi_{2}-\psi_{1} \psi_{2}^{\prime}\right)^{\prime}=0$. Integrating the last equation we have $\psi_{1}^{\prime} \psi_{2}-\psi_{1} \psi_{2}^{\prime}=C$ for any $x$. Constant $C$ vanishes because eigenfunction and its first derivative for boundary states vanishes in both infinities. Furthermore, integrating $\psi_{1}^{\prime} / \psi_{1}=\psi_{2} / \psi_{2}^{\prime}$ we have $\psi_{1}=A \psi_{2}$ and because states are normalized constant $A$ is phase factor specifying that states $\psi_{1}(x)$ and $\psi_{2}(x)$ are the same physical states which contradicts assumptions.

Solution ??. Consider two eigenfunctions $\psi_{n}(x)$ and $\psi_{n+1}(x)$ with eigenvalues $E_{n}<$ $E_{n+1}$. From the equations $\psi_{n}^{\prime \prime}+2 m\left(E_{n}-V\right) \psi_{n} / \hbar^{2}=0$ and $\psi_{n+1}^{\prime \prime}+2 m\left(E_{n+1}-V\right) \psi_{n+1} / \hbar^{2}=$ 0 , after simple calculation we obtain $\left.\left(\psi_{n}^{\prime} \psi_{n+1}-\psi_{n+1}^{\prime} \psi_{n}\right)\right|_{a \ldots ? ? ?}$

Solution ??. Adding and subtracting normal and space inverted Schrödinger equation we get two equalities $(V(x)-E)(\psi(x) \pm \psi(-x))=0, \forall x$, giving that eigenfunctions are either even or odd $(\psi(-x)= \pm \psi(x))$.

Solution ??. ???
Solution ??. ???

Solution ??. ???
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Solution ??. ???
Solution ??. ???
Solution 2.4. ???
Solution ??. ???
Solution 2.5. ???
Solution 2.6. ???
Solution 2.7. ???
Solution 2.8. ???
Solution 2.9. ???

## B.0.3 Galilean transformations

## Solution 3.1. ???

Solution 3.2. For even $n, \mathcal{I}^{n}$ is unity while for odd $n, \mathcal{I}^{n}=\mathcal{I}$. Expanding operator $e^{\mathrm{i} \pi \mathcal{I}}$ one gets $e^{\mathrm{i} \pi \mathcal{I}} \psi(\mathbf{r})=\sum_{n}(\mathrm{i} \pi)^{n} / n!\mathcal{I}^{n} \psi(\mathbf{r})=\sum_{k}(-)^{k} \pi^{2 k} /(2 k)!\psi(\mathbf{r})+\mathrm{i} \sum_{k}(-)^{k} \pi^{2 k+1} /(2 k+$ 1)! $\mathcal{I} \psi(\mathbf{r})=\cos (\pi) \psi(\mathbf{r})+\mathrm{i} \sin (\pi) \psi(-\mathbf{r})=-\psi(\mathbf{r})$

Solution 3.3. Introducing unity $I$ and space inversion $\mathcal{I}$ operators, one can write any function as sum of even and odd functions: $\psi(\mathbf{r})=(\psi(\mathbf{r})+\psi(-\mathbf{r})) / 2+(\psi(\mathbf{r})-\psi(-\mathbf{r})) / 2=$ $(I+\mathcal{I}) / 2 \psi(\mathbf{r})+(I-\mathcal{I}) / 2 \psi(\mathbf{r})=P_{+} \psi(\mathbf{r})+P_{-} \psi(\mathbf{r})$. Operator $\mathcal{I}$ is Hermitian as it follows from: $(\psi(\mathbf{r}), \mathcal{I} \varphi(\mathbf{r}))=\int_{-\infty}^{\infty} \mathrm{d} x \int_{-\infty}^{\infty} \mathrm{d} y \int_{-\infty}^{\infty} \mathrm{d} z \psi(\mathbf{r}) \varphi(-\mathbf{r})=$
$-\int_{\infty}^{-\infty} \mathrm{d} x \int_{\infty}^{-\infty} \mathrm{d} y \int_{\infty}^{-\infty} \mathrm{d} z \psi(-\mathbf{r}) \varphi(\mathbf{r})=\int_{-\infty}^{\infty} \mathrm{d} x \int_{-\infty}^{\infty} \mathrm{d} y \int_{-\infty}^{\infty} \mathrm{d} z \psi(-\mathbf{r}) \varphi(\mathbf{r})=(\mathcal{I} \psi(\mathbf{r}), \varphi(\mathbf{r}))$. Consequently, operators $P_{ \pm}$are Hermitian, and idempotent $P_{ \pm}^{2}=(I \pm \mathcal{I})^{2} / 4=(I \pm 2 \mathcal{I}+$ $\left.\mathcal{I}^{2}\right) / 4=P_{ \pm}$.

Solution 3.4. Operator $U_{\alpha \beta}$ is unitary if $\alpha$ and $\beta$ are pure imaginary numbers. Operator $U_{\alpha \beta}$ is space inversion if $S_{x}=U_{\alpha \beta} x U_{\alpha \beta}^{\dagger}=-x$ and $S_{p}=U_{\alpha \beta} p U_{\alpha \beta}^{\dagger}=-p$. Using equality
from Prob. ?? we have: $S_{x}=x+\left[\alpha x^{2}+\beta p^{2}, x\right]+\frac{1}{2!}\left[\alpha x^{2}+\beta p^{2},\left[\alpha x^{2}+\beta p^{2}, x\right]\right]+\cdots=$ $\sum_{i=0}^{\infty} \frac{1}{2!} K_{i}$, where $K_{i+1}=\left[K_{i}, x\right]$. It is easy to calculate commutators: $K_{0}=x, K_{1}=$ $-2 \mathrm{i} \hbar \alpha p, K_{2}=-2 \mathrm{i} \hbar \alpha 2$

Solution ??. ???

## B.0.4 Angular momentum

Solution 4.2. Obviously, if $K_{p}|x\rangle=0(p=1,2,3)$, then also $\left\langle K_{p}^{2}\right\rangle_{|x\rangle}=\langle x| K_{p} K_{p}|x\rangle=$ $\| K_{p}|x\rangle \|=0$, and $\left\langle K^{2}\right\rangle_{|x\rangle}=0$ (and $k=m=0$ ).
In the opposite direction, from $\left\langle K_{i}^{2}\right\rangle_{|x\rangle}=\left\langle K_{j}^{2}\right\rangle_{|x\rangle}=0$, the same argument implies that $\| K_{i}|x\rangle\|=\| K_{j}|x\rangle \|=0$, i.e. $K_{i}|x\rangle=K_{j}|x\rangle=0$. However, then $0=\left[K_{i}, K_{j}\right]|x\rangle=$ $K_{k}|x\rangle$, for the remaining component $K_{k}$.

Solution 4.1. Straightforward application of (4.8) in the representation formula $K_{i}|k m\rangle=$ $\sum_{m^{\prime}=-k}^{k} D_{m^{\prime} m}^{(k)}\left(K_{i}\right)\left|k m^{\prime}\right\rangle$.
Solution 4.2. Use the form (4.12) for $K^{2}$, and then substitute the expressions (4.8) for the action of $K_{3}$ and $K_{ \pm}$on the standard basis.

Solution 4.3. As squares of Hermitean operators, $K_{i}^{2}$ are positive, implying that their average vanishes if and only if $K_{i}|x\rangle=0$. Then, if $\left\langle K_{i}^{2}\right\rangle_{|x\rangle}=\left\langle K_{j}^{2}\right\rangle_{|x\rangle}=0$, also $\left[K_{i}, K_{j}\right]|x\rangle=0$, and $\left\langle K^{2}\right\rangle_{|x\rangle}=0$. This means that $k=m=0$.

Solution 4.4. ???

## B.0.5 Identical particles

Solution 5.1. All the transpositions (which appear only for $N>1$ are in the same conjugation class of $\boldsymbol{S}_{N}$, since any pair of them is mutually conjugated: $\tau_{p p^{\prime}}=\tau_{p q} \tau_{p^{\prime} q^{\prime}} \tau_{q q^{\prime}} \tau_{p q} \tau_{p^{\prime} q^{\prime}}$ (conveniently, $\tau_{i i}=e$ ). Therefore, they have the same character in each representation, and as for the one dimensional ones, they are represented by the same numbers. As obviously $\tau_{p p^{\prime}}^{2}=e$ (identity element), homomorphism gives for any representation $D^{2}\left(\tau_{p p^{\prime}}\right)=1$, giving for the one dimensional representations $D\left(\tau_{p p^{\prime}}\right)= \pm 1$. Taking one of these two choices coincidentally for all transpositions, one builds up the numbers representing the other permutations as the product of even or odd number of $\pm$ factors.

Solution 5.2. The two bases are necessarily related as $\left|j_{p}\right\rangle=\sum_{l} U_{l p}\left|i_{l}\right\rangle$, with $U$ being an unitary operator (thus $\operatorname{det}(U)$ is pure phase). The new Slater determinant is $\left|\boldsymbol{n}^{-} ; j\right\rangle=\frac{1}{\sqrt{N!}} \operatorname{det}\left|j_{p}\right\rangle_{q}=\frac{1}{\sqrt{N!}} \operatorname{det} \sum_{l} U_{l p}\left|i_{l}\right\rangle_{q}$. Noticing that the form of the $p q$-th element is $\left(U^{T}\left|i_{l}\right\rangle\right)_{p q}$, the homomorphism of the determinant functional, $\operatorname{det} A \operatorname{det} B=\operatorname{det} A B$, directly gives $\left|\boldsymbol{n}^{-} ; j\right\rangle=\operatorname{det}(U)\left|\boldsymbol{n}^{-} ; i\right\rangle$.

Solution 5.3. All the terms in the sum (5.13) contain the same quantum numbers, and obviously the wanted partial scalar product vanishes if the set of quantum numbers $\left\{i_{1}, \ldots, i_{N}\right\}$ is not a superset of $\left\{j_{L+1}, \ldots, j_{N}\right\}$, i.e. when some of the components of $\boldsymbol{n}_{L}=\boldsymbol{n}-\boldsymbol{n}_{N-L}$ are negative. Otherwise, there exists a permutation $\sigma$ defined in the exercise, rearranging the quantum numbers $\left\{i_{1} \leq \cdots \leq i_{N}\right\}$ such that the last
$N-L$ coincide with $\left|j_{L+1}, \ldots, j_{N}\right\rangle$ and the first $L$ are ordered. Then the $L$ particle state $\left|\Psi_{L}\right\rangle=\left\langle j_{L+1}, \ldots, j_{N} \mid \boldsymbol{n}^{ \pm}\right\rangle$is:

$$
\left|\Psi_{L}\right\rangle=\frac{( \pm)^{\tilde{\sigma}}}{\sqrt{n!N!}} \sum_{\pi}( \pm)^{\tilde{\pi}}\left\langle j_{L+1}, \ldots, j_{N}\right| \Delta(\pi)\left|i_{\sigma^{-1}} \leq \cdots \leq i_{\sigma^{-1} L}, j_{L+1}, \ldots, j_{N}\right\rangle .
$$

Since bra-vector is fixed, the partial scalar product involving the permutation $\pi$ does not vanish only when $\pi$ leaves invariant the states of the particles $L+1, \ldots, N$. Each such permutation may be written as a product $\pi=\pi_{L} \pi_{n}$, where $\pi_{L} \in \boldsymbol{S}_{L}$ permutes only the first $L$ particles, and $\pi_{\boldsymbol{n}} \in \boldsymbol{S}_{\boldsymbol{n}}$. However, the intersection of the groups $\boldsymbol{S}_{L}$ and $\boldsymbol{S}_{\boldsymbol{n}}$ is just the group $\boldsymbol{S}_{n_{L}}$, the stabilizer of the $L$ particle state $\left|i_{\sigma^{-1}} \leq \cdots \leq i_{\sigma^{-1} L}\right\rangle$. Therefore, the factorization $\pi_{L} \pi_{n}$ is not unique, and in the set of the products $\pi_{L} \pi_{n}$ each permutation appears exactly $\boldsymbol{n}_{L}!$ times. One gets

$$
\left|\Psi_{L}\right\rangle=\sum_{\pi_{L}, \pi_{n}} \frac{( \pm)^{\tilde{\sigma}+\tilde{\pi}_{L}+\tilde{\pi}_{n}}}{\boldsymbol{n}_{L}!\sqrt{\boldsymbol{n}!N!}}\left\langle j_{L+1}, \ldots, j_{N}\right| \Delta\left(\pi_{L} \pi_{\boldsymbol{n}}\right)\left|i_{\sigma^{-1}} \leq \cdots \leq i_{\sigma^{-1} L}, j_{L+1}, \ldots, j_{N}\right\rangle
$$

Finally, one realizes that

$$
\left\langle j_{L+1}, \ldots, j_{N}\right| \Delta\left(\pi_{L} \pi_{n}\right)\left|i_{\sigma^{-1}} \leq \cdots \leq i_{\sigma^{-1} L}, j_{L+1}, \ldots, j_{N}\right\rangle=\Delta\left(\pi_{L}\right)\left|i_{\sigma^{-1}} \leq \cdots \leq i_{\sigma^{-1} L}\right\rangle
$$

independently of $\pi_{\boldsymbol{n}}$; hence there are $\boldsymbol{n}$ ! such terms (all with the same sign: for fermions $\pi_{n}$ is the identical permutation only, and for bosons all terms are with the same sign), giving

$$
\left|\Psi_{L}\right\rangle=\frac{( \pm)^{\tilde{\sigma}} \boldsymbol{n}!}{\boldsymbol{n}_{L}!\sqrt{\boldsymbol{n}!N!}} \sum_{\pi_{L}}( \pm)^{\tilde{\pi}_{L}} \Delta\left(\pi_{L}\right)\left|i_{\sigma^{-1} 1} \leq \cdots \leq i_{\sigma^{-1} L}\right\rangle .
$$

Summing over $\pi_{L}$, and including normalization factor, one immediately gets (5.18a). Analogously one gets (5.18b). As for $L=1$ case, the permutation $\sigma$ moves the first particle in the state $|j\rangle$, to the first place (where this state is cancelled by $\langle j|$ ). So, there are $l_{j}=\sum_{i<j} n_{i}$ states before, which are effectively cycled by $\sigma=\tau_{12} \cdots \tau_{l_{j}-1, l_{j}} \tau_{l_{j}, l_{j}+1}$, showing that $l_{j}=\tilde{\sigma}$.

Solution 5.4. First, note that the permutation $\sigma$ from the Exercise 5.5 may be given a canonical form $\sigma=\pi_{N-L} \lambda$, where $\lambda$ is a "riffle shuffle" permutation, which on the set $\left\{i_{1}, \ldots, i_{n}\right\}$ with singled out $L$ (and therefore the remaining $N-L$ ) elements simply separates them, putting the $L$ elements first and $N-L$ after them, without changing the order within subsets, while $\pi_{N-L}$ permutes only the second subset. Taking into account (5.13) for $N-L$ particles, one gets $\left\langle\boldsymbol{n}_{N-L}^{ \pm} \mid \boldsymbol{n}^{ \pm}\right\rangle=\sqrt{\frac{1}{\boldsymbol{n}_{N-L}!(N-L)!}} \sqrt{\frac{n!L!}{n_{L}!N!}}(N-L)!( \pm)^{\tilde{\lambda}}\left|\boldsymbol{n}_{L}^{ \pm}\right\rangle$.

Solution 5.5. From (5.18) it immediately follows that
$\rho_{1, \ldots, L}=\sum_{j_{L+1}, \ldots, j_{N}}\left\langle j_{L+1}, \ldots, j_{N} \mid \boldsymbol{n}^{ \pm}\right\rangle\left\langle\boldsymbol{n}^{ \pm} \mid j_{L+1}, \ldots, j_{N}\right\rangle=\binom{N}{L}^{-1} \sum_{\boldsymbol{n}_{L} \leq \boldsymbol{n}} C_{\boldsymbol{n}_{L}} \frac{\boldsymbol{n}!L!}{\boldsymbol{n}_{L}!N!}\left|\boldsymbol{n}_{L}^{ \pm}\right\rangle\left\langle\boldsymbol{n}_{L}^{ \pm}\right|$,
where the weight $C_{\boldsymbol{n}_{L}}$ shows how many times occupation number $\boldsymbol{n}_{L}$ appears during the summation in $\boldsymbol{j}=\left(j_{L+1}, \ldots, j_{N}\right)$. To find it we observe that $\boldsymbol{n}_{L}$ appears for $\boldsymbol{j}$ such that
$\boldsymbol{n}_{L}=\boldsymbol{n}-\boldsymbol{n}(\boldsymbol{j})$. For fixed $\boldsymbol{n}_{L}$, also $\boldsymbol{n}(\boldsymbol{j})$ is fixed, and from $(N-L)$ ! permutations of the set $\boldsymbol{j}$ altogether $\boldsymbol{n}(\boldsymbol{j})$ ! give the same arrangement of the indices $\boldsymbol{j}$, i.e. they correspond to the same term in sum over $\boldsymbol{j}$. So, there are $C_{\boldsymbol{n}_{L}}=(N-L)!/ \boldsymbol{n}(\boldsymbol{j})$ ! different terms with same $\boldsymbol{n}_{L}$.

Solution 5.6. As now $\boldsymbol{n}!=\boldsymbol{n}_{L}!=1$ (5.21a) becomes

$$
\rho_{1, \ldots, L}=\sum_{n_{L} \leq \boldsymbol{n}} \frac{\left|\boldsymbol{n}_{L}^{ \pm}\right\rangle\left\langle\boldsymbol{n}_{L}^{ \pm}\right|}{\binom{N}{L}}=\frac{L!}{\binom{N}{L}} \sum_{i_{1}<\cdots<i_{L}=1}^{N} P_{L}^{(-)}\left|i_{1}, \ldots, i_{L}\right\rangle\left\langle i_{1}, \ldots, i_{L}\right| P_{L}^{(-)} .
$$

The primed sum means that all the indices run over the $N$ single particle states defined by $n_{s}=1$ in $\boldsymbol{n}$ (alternatively, we may assume that the single particle space basis $|i\rangle$ starts with these $N$ occupied states and indices run from 1 to $N$ ). Due to the antisymmetrizer, the restrictions $i_{1}<\cdots<i_{L}$ in the summation may be omitted: the repeated indices are neglected by the antisymmetrizers, and only unordered permutations of ( $i_{1}<\cdots<i_{L}$ ) are effectively added. However, corresponding vectors $P^{(-)}\left|i_{1}, \ldots, i_{L}\right\rangle$ are up to the sign same as for the ordered one, and the same sign appears both in bra and ket. Hence there are $L$ ! same terms, instead of one which must be compensated by divison:

$$
\rho_{1, \ldots, L}=\frac{1}{\binom{N}{L}} \sum_{i_{1}, \ldots, i_{L}=1}^{N} P_{L}^{(-)}\left|i_{1}, \ldots, i_{L}\right\rangle\left\langle i_{1}, \ldots, i_{L}\right| P_{L}^{(-)} .
$$

It remains to notice that the summations in $i_{1}, \ldots, i_{L}$ become independent, each giving the projector $P_{\boldsymbol{n}}$, and that the direct product of $L$ these projectors commutes with the projector $P_{L}^{(-)}$, which can be omitted from one side.

Solution 5.7. By definition $\left(\hat{\hat{P}}^{(+)} A\right) P^{( \pm)}=\frac{1}{N^{2}} \sum_{\pi \pi^{\prime}}( \pm)^{\tilde{\pi}} \Delta^{-1}\left(\pi^{\prime}\right) A \Delta\left(\pi^{\prime}\right) \Delta(\pi)$. Denoting $\pi^{\prime} \pi$ by $\pi^{\prime \prime}$, one finds $\Delta\left(\pi^{\prime}\right) \Delta(\pi)=\Delta\left(\pi^{\prime \prime}\right)$ and $\Delta^{-1}\left(\pi^{\prime}\right)=\Delta^{-1}(\pi) \Delta^{-1}\left(\pi^{\prime \prime}\right)$. Substituting summation in $\pi^{\prime}$ and $\pi$ by summation in $\pi^{\prime \prime}$ and $\pi^{-1}$, respectively (rearrangement lema), one obtains $\left(\hat{\hat{P}}^{(+)} A\right) P^{( \pm)}=P^{( \pm)}\left(\hat{\hat{P}}^{(+)} A\right)$.

Solution 5.8. The operators $K_{ \pm}=\sum_{p=1}^{2 k+1} k_{p \pm}$ are symmetric (additive) operators, thus commuting with the antisymmetrizer (recall that otherwise the relevant space $\mathcal{S}_{-}^{2 k+1}$ would not be invariant; formally since $\hat{\hat{P}} K_{ \pm}=K_{ \pm}$, this follows from (5.24)). Thus the $2 k+1$ particle state (5.11) is defined by the occupation numbers $n_{m}=1, m=-k, \ldots, k$ (using single particle basis $|Q, k, m\rangle)$ :

$$
\left|\boldsymbol{n}^{-}\right\rangle=\sqrt{N!} P_{n}^{(-)}|\boldsymbol{n}\rangle=\sqrt{\frac{1}{N!}} \sum_{\pi}(-)^{\tilde{\pi}} \Delta(\pi)|Q, k,-k ; \ldots ; Q, k, k\rangle,
$$

giving

$$
K_{ \pm}\left|\boldsymbol{n}^{-}\right\rangle=\sqrt{\frac{1}{N!}} \sum_{p=1}^{2 k+1} \sum_{\pi}(-)^{\tilde{\pi}} k_{p \pm} \Delta(\pi)|Q, k,-k ; \ldots ; Q, k, k\rangle .
$$

Now note that each summand $k_{p \pm} \Delta(\pi)|Q, k,-k ; \ldots ; Q, k, k\rangle$ is zero: either at position $p$ after action of $\Delta(\pi)$ appears $k_{p \pm}|Q, k, m= \pm k\rangle=0$, or $k_{p \pm}|Q, k, \pm m\rangle \sim|Q, k, \pm m \pm 1\rangle$,
the state in which is another electron already. However, $K_{ \pm}\left|\boldsymbol{n}^{-}\right\rangle=K_{ \pm} P^{(-)}\left|\boldsymbol{n}^{-}\right\rangle=$ $P^{(-)} K_{ \pm}\left|\boldsymbol{n}^{-}\right\rangle$, and we can apply on each summand another antisymmetrization, resulting in the vanishing vector.

Solution 5.9. Acting on the Slater determinant $|1, \ldots, N\rangle$ (as $N=|\mathcal{S}|$ ) one gets $\left(\otimes_{p=1}^{N} D_{p}\right)|1, \ldots, N\rangle=\frac{1}{N!} \sum_{\pi}(-1)^{\pi} \sum_{j_{1}, \ldots, j_{N}} D_{j_{1}, \pi^{-1} 1} \cdots D_{j_{N}, \pi^{-1} N}\left|j_{1}\right\rangle \otimes \cdots \otimes\left|j_{N}\right\rangle$, which is (reordering the matrix element) $\frac{1}{N!} \sum_{\pi}(-1)^{\pi} \sum_{j_{1}, \ldots, j_{N}} D_{j_{\pi 1} 1} \cdots D_{j_{\pi N} N}\left|j_{1}\right\rangle \otimes \cdots \otimes\left|j_{N}\right\rangle$. Changing $j_{\pi p}$ to $j_{p}$ gives $\frac{1}{N!} \sum_{\pi}(-1)^{\pi} \sum_{j_{1}, \ldots, j_{N}} D_{j_{1} 1} \cdots D_{j_{N} N}\left|j_{\pi^{-1} 1}\right\rangle \otimes \cdots \otimes\left|j_{\pi^{-1} N}\right\rangle=$ $\sum_{j_{1}, \ldots, j_{N}} D_{j_{1} 1} \cdots D_{j_{N} N} P^{-}\left|j_{1}, \ldots, j_{N}\right\rangle$. The antisymmetrizer annihilates the terms with repeated indices $j_{p}$; as $P^{-}\left|j_{1}, \ldots, j_{N}\right\rangle=(-1)^{\pi\left(j_{1}, \ldots, j_{N}\right)}|1, \ldots, N\rangle$, this becomes
$\left(\sum_{j_{1}, \ldots, j_{N}}(-1)^{\pi\left(j_{1}, \ldots, j_{N}\right)} D_{j_{1} 1} \cdots D_{j_{N} N}\right)=(\operatorname{det} D)|1, \ldots, N\rangle$. The second part of the Exercise is an obvious generalization of the well known equality $\operatorname{det} A \otimes B=(\operatorname{det} A)^{|B|}(\operatorname{det} B)^{|A|}$.

Solution 5.10. Compare the definition (5.31) to (5.18c). As for $a_{s}^{\dagger}$, we again consider its left action in $\left|\boldsymbol{n}^{\prime \pm}\right\rangle a_{s}^{\dagger}\left|\boldsymbol{n}^{ \pm}\right\rangle=\sqrt{N}\left(\left\langle\boldsymbol{n}^{\prime \pm} \mid s\right\rangle\right)\left|\boldsymbol{n}^{ \pm}\right\rangle$. Thus, the expression vanishes unless $\left|\boldsymbol{n}^{\prime p m}\right\rangle$ is the unique one satisfying $\left\langle s \mid \boldsymbol{n}^{\prime \pm}\right\rangle \sim\left|\boldsymbol{n}^{ \pm}\right\rangle$. Apparently, this is just $\left|\boldsymbol{n}^{ \pm}\right\rangle|s\rangle$ (everywhere it is assumed that $|s\rangle$ refers to the last particle). However, although it has the correct occupation number, it is not (anti)symmetrized, thus it is not in the Fock space: the adjoint of the operator in $\mathcal{F}_{ \pm}$must be also an operator in the same space. Thus this vector is to be (anti)symmetrized, but previously quantum number $s$ is to be put (i.e. permuted by $\sigma$ ) to the position $p_{s}$, such that the states after it are with greater quantum numbers. Altogether, this exactly gives the same result as (5.32). The generalization (5.33) is obtained by the successive application of the first relation. N independent form is achieved using the number of particles operator: $\left\langle j_{1}, \ldots, j_{L} \mid \boldsymbol{n}^{ \pm}\right\rangle=$ $a_{j_{L}} \frac{1}{\sqrt{\hat{N}}} \cdots a_{j_{1}} \frac{1}{\sqrt{\hat{N}}}\left|\boldsymbol{n}^{ \pm}\right\rangle$.

Solution 5.11. Consider a particle with one dimensional state space. Due to this peculiarity, all the observables are real numbers, among them the hamiltonian being $\hbar \omega(\omega \geq 0)$. Further, any $N$-particle space is again one dimensional, spanned by the vector $|N\rangle$ (as there is only one single particle basis vector, occupation number vector is simply the total number of particles). The particles do not interact, making the $N$-particle hamiltonian $N \hbar \omega$. The particles may be created or annihilated, thus they must be considered within Fock's space, which is exactly the state space of the harmonic oscillator. It only remains to shift the energy scale for $\bar{\omega} / 2$, and the oscillator dynamics is completely reproduced by the described noninteracting particles, which really correspond to the excitations of the harmonic oscillator. The commutation relation (2.26) shows that such excitations are of Bose type.

Solution 5.12. Consider an arbitrary matrix element $A_{\boldsymbol{n}^{\prime}}^{\boldsymbol{n}}=\left\langle\boldsymbol{n}^{ \pm}\right| A^{[L]}\left|\boldsymbol{n}^{\prime \pm}\right\rangle$ in the occupation number basis:

$$
A_{\boldsymbol{n}^{\prime}}^{n}=\sum_{p_{1}<\cdots<p_{l}} \sum_{\substack{s_{1}, \ldots, s_{L} \\ s_{1}^{\prime}, \ldots, s_{L}^{L}}} A_{s_{1}^{\prime}, \ldots, s_{L}^{\prime}}^{s_{1}, \ldots, s_{L}}\left\langle\boldsymbol{n}^{ \pm} \mid s_{1}^{\left(p_{1}\right)}, \ldots, s_{L}^{\left(p_{L}\right)}\right\rangle\left\langle s_{1}^{\left(p_{1}\right)}, \ldots, s_{L}^{\prime\left(p_{L}\right)} \mid \boldsymbol{n}^{\prime \pm}\right\rangle
$$

superscript is particle counter, i.e. noncorrelated vector $\left|s_{1}^{\left(p_{1}\right)}, \ldots, s_{L}^{\left(p_{L}\right)}\right\rangle$ describes a system of particles $p_{1}<\cdots<p_{L}$, with the particle $p_{j}$ in the state $s_{j}$. To apply (5.33), one
previously should move particles $p_{1}<\cdots<p_{L}$ to the first $L$ places; however, the sign arising (for fermions only) in such a permutation, appears both for bra and ket. Thus, summation in particles gives $\binom{N}{L}$ same terms:

$$
A_{\boldsymbol{n}^{\prime}}^{n}=\binom{N}{L} \sum_{\substack{s_{1}, \ldots, s_{L} \\ s_{1}^{\prime}, \ldots, s_{L}^{L}}} A_{s_{1}^{\prime}, \ldots, s_{L}^{\prime}}^{s_{1}, \ldots, s_{L}}\left\langle\boldsymbol{n}^{ \pm} \mid s_{1}, \ldots, s_{L}\right\rangle\left\langle s_{1}^{\prime}, \ldots, s_{L}^{\prime} \mid \boldsymbol{n}^{ \pm}\right\rangle
$$

Now, (5.33) directly gives the equality of the arbitrary matrix elements of the operators in (5.39c).

Solution 5.13. Defining operators $a_{Q k m}$ corresponding to the basis $|Q k m\rangle$, and noticing that $K_{ \pm}$is additive single particle operator such that $\langle Q k m| k_{ \pm}\left|Q^{\prime} k^{\prime} m^{\prime}\right\rangle=\alpha_{m^{\prime}}^{ \pm} \delta_{Q Q^{\prime}} \delta_{k k^{\prime}} \delta_{m \mp 1, m^{\prime}}$, one finds $K_{ \pm}=\sum_{m} \alpha_{m}^{ \pm} a_{m}^{\dagger} a_{m \mp 1}$. The closed shell corresponds to the state $\left|\boldsymbol{n}^{-}\right\rangle$, such that the occupation numbers of all the states are 0 except that $n_{Q k m}=1$, for $m=-l, \ldots, l$. Obviously, $K_{ \pm}$maps this state into the vector with occupation numbers greater than 1, i.e. into the vanishing vector.

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[^0]:    ${ }^{1}$ Here the opposite of equivalence $\nLeftarrow$ emphasizes that exactly one of two possible paths is realized.

[^1]:    ${ }^{2}$ Probability measure $\mu$ over the subspaces is a function $\mu$ mapping each subspace of $\mathcal{S}$ into nonnegative real number, such that $\sum_{i} \mu\left(\mathcal{S}_{i}\right)=1$ for any orthogonal decomposition $\mathcal{S}=\oplus_{i} \mathcal{S}_{i}$.

[^2]:    ${ }^{1}$ On the other hand, this view introduced new mathematical techniques, particularly the Wienner measures, which are still developing. This resembles to the Dirac's $\delta$-function, and its influence to functional analysis.
    ${ }^{2}$ It is well known that $\int_{-\infty}^{\infty} \mathrm{e}^{\mathrm{i}\left(\alpha p+\beta p^{2}\right)} \mathrm{d} p=\sqrt{\frac{\pi}{\beta}} \mathrm{e}^{\mathrm{i}\left(\frac{\pi}{4}-\frac{\alpha^{2}}{4 \beta}\right)}$; it is assumed that $\beta$ is positive, while in our case it is imaginary. This will be discussed after (2.42).

[^3]:    ${ }^{1}$ Mathematical definition of the exponential relation between Lie group and its algebra (real!) does not include the imaginary unit, and therefore unitary representations of the group are obtained from the skew hermitian representation of the algebra. This factor enables that the Lie algebra elements in physics are hermitian (physical observables), but non-closeness of the commutator in the space of hermitian operators must be cured by another factor $\mathrm{i} \hbar$ in the commutators, as in (4.1). This may be also understood as a change of basis in the complexified Lie algebra.

[^4]:    ${ }^{2}$ Sometimes, instead of Clebsch-Gordan coefficients, their particular biunique function called $3 j$ symbols are used.

[^5]:    ${ }^{1}$ In principle, it is also possible that several unperturbed levels $E_{n}(0), E_{n^{\prime}}(0), \ldots$, give paths coinciding in $a=1$. For differentiable curves $E_{n}(a), E_{n^{\prime}}(a), \ldots$, the coincidence may occur in the countable number of points, and this is tacitly assumed in the forthcoming expansion.

[^6]:    ${ }^{2}$ Alternatively, for some purposes also $V_{\text {ion }}(\boldsymbol{R})$ can be omitted from $H_{\boldsymbol{r}}$, with the same conclusions.

[^7]:    ${ }^{3}$ Variation of the functional $F[\varrho]$ is defined as $\delta F[\varrho] \stackrel{\text { def }}{=} \frac{\partial}{\partial \alpha}(F[\varrho(\boldsymbol{r})+\alpha \delta \varrho]){ }_{\alpha=0}$. The functional derivative is the functional of $\varrho$ comprising the part of variation linear in $\delta \varrho$. For the local functional $F[\varrho]=\int \mathrm{d} \boldsymbol{r} f[\boldsymbol{r}, \varrho]$, the variation is $\int \mathrm{d} \boldsymbol{r} f_{\varrho}^{\prime}[\boldsymbol{r}, \varrho] \delta \varrho$, making $\frac{\delta F[\varrho]}{\delta \varrho}=f_{\varrho}^{\prime}[\boldsymbol{r}, \varrho(\boldsymbol{r})]$.
    ${ }^{4}$ Exchange term includes correlation due to antisymmetrization only, while the rest is gathered in correlation term; this means that Slater determinant gives $C=0$, while $C \neq 0$ is provided only by their linear combinations.

